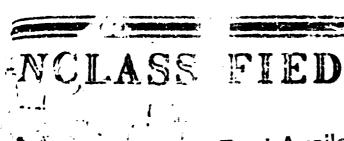
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increased Absorption by Carbon Dioxide, Water Vapor, and Minor Atmospheric Constituents

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# Infrared Absorption by Carbon Dioxide, Water Vapor, and Minor Atmospheric Constituents

DARRELL E. BURCH, DAVID GRYVNAK, EDGAR B. SINGLETON, WILBUR L. FRANCE, AND DUDLEY WILLIAMS OHIC STATE UNIVERSITY CONTRACT AF 19(604)-2633

OPTICAL PHYSICS LABORATORY PROJECT MEDIA

# PREFACE

In 1956 Citto published Geophysical Research Papers No. 40, which have a meaned originally as an Onio State University contract report, to make a convenient scally available to a wider audience. This report contained the results by J. N. Howard, D. E. Burch, and D. Williams at the Ohio State University of their studies on long path length infrared absorption of water vapor and carbodioxide. This report was almost a unique source for this type of data and was and still is in great demand.

Shortly after that time Dr. Howard left the Ohio State group, but Dr. Burch and others continued these infrared studies, extending them to other atmospheric constituents such as N<sub>2</sub>O, CH<sub>4</sub>, and CO and to smaller concentrations and lower pressure conditions more appropriate to higher altitudes. In addition part of the above work was repeated with better spectral resolution. The results of this work under contract AF19(604)-2633 with the Ohio State University were brought out in three contract reports: AFCRL-TN-60-674, AFCRL 255, and AFCRL 412. The demand for these reports has exhausted the supply. They are now being published as a single paper in the AFCRL Research Report series for wider dissembation.

# ABSTRACT

This report presents results of studies concerned with the absorption of infrared radiation by various atmospheric gases. It places under one cover three previously published centractor reports on (1) absorption by nitrous oxide, car. — monoxide, and methane; (2) absorption by carbon dioxide; and (3) absorption by water vapor and nitrous oxide.

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by

Darrell E. Durch and Dudley Williams

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# ABSORPTION BY NUTROUS OXIDE, CARBON MONOXIDE AND MITHANE

#### I. INTRODUCTION

A detailed knowledge of the infrared transmission of planets ystamapheres is of importance to space physics in connection with studies of heat balances and with the possible development of infrared proting and ignaling systems. During the past three-quarters of a century numerous

dies of infrared transmission of the earth's atmosphere have been made. Molecular absorption in the atmosphere is caused chiefly by water vapor, carbon dioxide, and ozone, but our presence of small quantities of nitrous oxide, carbon monoxide, and methane also results in significant absorption; molecular oxygen also has absorption bands in the very near infrared. Infrared rediction undergoed attenuation by particles of fug, haze, and smoke in the atmosphere, this type of attenuation, as well as uncontrollable factors such as temperature and humidity, makes it difficult to obtain quantitative data on true molecular absorption in field studies. Hence, laboratory studies of "synthetic atmospheres" seem desirable even in telluric atmospheric work, and provide a poverful method of investigating absorption in the atmospheres of ouner planets.

The absorption of mear infrared radiation by atmospheric gases results chiefly from transitions between the rotation-vibration energy levels of molecules; absorption in the far infrared is produced by transitions between rotational levels. The observed transmission by tlean samples of atmospheric gases is therefore a function of frequency v. If  $I_{\Omega}(v)$  represents the intensity of a beam of incident radiation as measured by a monochromator set for frequency win the vicinity of an absorption band, it is found that the intensity I(v) of the radiation after passage through a sample consisting of a mixture of gases is less than  $L_0(\nu)$ . The ratio  $T(\nu) = I(\nu)/I_0(\nu)$  is called the measured transmission of the gas sample at frequency w and is expressed either as a frection or a percentage. The measured fractional absorption at frequency v is given by A(v) = 1T(v); A(v) may also be stated as a percentage. It is found that the measured theorptica A(v) depends on such factors as the total number of absorbing molecules per cross-sectional area of the beam, denoted by w and called the absorber concentration or optical density; the partial pressure p of the absorbing gar; the total pressure P, which includes the markal pressure of absorbing and nonateorbing gases; the absolute temperature T of the sample; and the effective alityidth of the mono-

In 1945 Chapman and Howard' of this department began a laboratory study of the absorption of water vapor and carbon dioxide samples containing various amounts of nonabsorbing games. The gas samples were enclosed in a multiple-traversal cell giving a maximum path length of several meters; the temperature of the gas samples could be elevated by electrical heating of the entire cell. In 1953 Howard, Bursh, and Williams' (hereafter referred to as HBM), also of this laboratory, extended the earlier studies of water vapor and carbon dioxide to longer math lengths by the use of a C2-meter multiple-traversal absorption cell, in which absorption paths as great as 1-10 meters rould be obtained. The work reported by HBM was done with games at room temperature.

In these earlies abudies it was verified that although the measured values of A(v) for a given example and given juch length are strongly decomposite of a continuous and a a continuous

$$\int_{\nu_1}^{\nu_2} A(\nu) d\nu \qquad \text{Total Absorption} \tag{1}$$

is independent of slithwidth provided the rimits  $v_1$  and  $v_2$  are chosen to include an entire absorption band; the validity of this statement is discussed in some detail by HEW. The integral (1) is called the total absorption or the absorption band defined by the limits  $v_1$  and  $v_2$  and is expressed in frequency units, other authors have referred to the integral (1) as the equivalent band width of the absorption, since the same integral can be considered as applying to an equivalent band having complete absorption over a inequency interval

$$\mathcal{W} = \int_{V_1}^{V_2} A(v) dv .$$

The work of Unapman and Howard and of HBW was directed toward the determination of the total absorptions of wall is infrared bands of water rapor and carbot divide mixed with nitrogen as functions of absorber concentration w, partial pressure p of the absorbing gas, total pressure P, and absorbe the pressure T. The functional relationships between total absorption and the parameters listed were given by tables, by suitable graphs, and by empirical equations applicable to certain restricted ranges of total absorption for each band.

The present research program includes (f) further studies of absorption by water vapor and carbon dickide vita an extension to spectral regions and to ranges of w, g, P, and T not covered in earlier studies, and (ii) studies of the infrared absorption produced by the minor amospheric constituents siculate carbon tonordes, and methans. The purpose of the investigation is to establish the functional relationship

$$\int_{\psi_1}^{\psi_2} \Lambda(\psi) d\psi = \mathcal{G}(\psi, \ p - F, \ T) \tag{2}$$

for various regions of characteristic absorption. The functional relation will, as in the earlier reports, be given by tables, graphs, and empirical equations. It is happed that the results of the investigation will prove useful in the following cases:

 (a) predictions of acception for known paths through known absorping on all studes;

- (b) determination of atmospheric composition by measurements of the transmission of infrared without along known meth lengths in the atmosphere; and
- (c) test of the various theories that have been proposed to account for absorption phenomena.

The results obtained in the present reseason program will be given in three reports as follows:

NEFORT I (ITERATE): Absorption by nitrous oxide in use 2004 (at legion; postryption by methon monosthe in the region of its fundamental (2145 cm<sup>-1</sup>) band and its first overtone (~250 cm<sup>-1</sup>) band; absorption of methans bands at 3020 cm<sup>-1</sup>, 1550 cm<sup>-1</sup>, and 1506 cm<sup>-1</sup>.

REFORT II: Absorption by parton dioxide bands at 2500 cm<sup>-1</sup>, at 3600 cm<sup>-1</sup>, at 1716 cm<sup>-1</sup>, at 961 cm<sup>-1</sup>, at 1064 cm<sup>-2</sup>, and in the range 878 - 450 cm<sup>-2</sup>.

FIRAL REPORT: Absorption by water vapor in the region of its vibration-rotation bands at 1595 cm<sup>-1</sup>
and near 3700 cm<sup>-1</sup>, and in the region of its pure
rotational band in the range 480 - 1000 cm<sup>-1</sup>, absorption by nitrous oxide bands at 789 cm<sup>-1</sup>, .LLT cm<sup>-1</sup>.
LLS cm<sup>-1</sup>, 2461 cm<sup>-1</sup>, and 2463 cm<sup>-1</sup>, a shully of the
rotative effects of various foreign gases on the
total accomption of several labels.

A portion of the tata on the MgO, DO, and DOg banis were reported previously in a Ph.D. dissertation by one or usi. Sowever, additional lata which are included in Deports I and II were not included in the previous work. As a result of these additional fata some of the empirical relations which now fewered have been middled alightly.

### 11. THEORY

Theories of absorption are based on an exponential law of absorption which was proposed by Lambert nearly two hundred years ago on the basis of studies of absorption of light by inquids, and was later elaborated by Beer. According to the Lambert law, the fractional absorption  $A^*(\nu)$  for monochromatic radiation of frequency  $\nu$  is given by the expression

$$A'(v) = 1 - e^{-k(v)w}$$
 (3)

where h (:) is the normal equal the absorption operation for the absorbing material and y is the absorber concentration. Beer's elaboration of the law was concerned with a mixture of absorbing m. orials. According to Beer's law

$$k(v) = C_1 e_1 + C_2 e_2 + C_3 e_4 \dots$$
 (4)

where  $C_1$ ,  $C_2$ ,  $C_4$ ,... are the concentrations of the various components and  $e_1$ ,  $e_2$ ,  $e_3$ ,... are constants constants called "extinc" of coefficients" for the various components. Been's law is based on the resumption that each absorbing molecule absorbs radiation independently of every other molecule.

Various early evilence: If the failure of Beer's Law for infrared absorption of gates are discussed by HBW. The failure of Beer's law can be traced on two sources.

(i) In the first place, specifography have finite slithwidths and therefore direct measurement of A'(v) is impossible, since truly monochromatic radiation is not being measured. In the tare of broad, relationly attucturaless bands observed for liquids, the measured abscrption A(v) approaches A'(v); it was this clove correspondence of A(v) and A'(v) that led tumbert and he r to the formulation of their laws. The infrared spectrum of a gas consists of bands compused of individual lines with resulting rapid waterions of A'(v) with v. With Hos vidule that are usually marrow as compared with the spectral intermal Av passed by a monochromaton or spectrograph set to pass indiction of frequency  $\nu$ , the measured  $A(\nu)$  is not usually a close approximation of  $A'(\nu)$ . Thus, it is not surprising that there is apparent failure of the simple absorption laws (5) and (4) if in a rest mentured absorption A(v) is used to mplace A'(v) in (c). How mr, the simple last fail for a nore fundamental reason even when more sophisticated comparisons or theory and experiment are made.

(ii) As indicated above, Beer's law is based on the assumption that the absorption of cath molecule is independent or every other molecule. This assumption is not justified for the case in which the absorbing molecule experiences collisions with other molecules, since molecular collisions result in appreciable broadening of the spectral lines.

Even in the case of noninteracting molecules, spectral lines have figite natural widths related to the uncertainty principle which predicts an uncertainty OF in an energy level related to the lifetime At of the molecule in the level by the relation AE-Atizh; since At 10 limited for all excited levels, all spectral lines involving transitions between ground states and excited states would be expected to have an uncertainty in frequency  $\Delta v = \Delta E/h \approx 1/\Delta t$ . Natural broadening is independent of effect. external to the molecule. Usually natural broadening is negligible is compared with line broadening due to the Doppler effect, which is re sted to the temperature of the gas sample and produces noticeable effer a when absorption of gases at very low pressures is measured; Doppler or adening is independent of pressure. The contributions of natural widths and Doppler broadening to the total widths of spectral lines are usually negligible as compared with contributions produced by the large electric fields that act during molecular collisions. The frefuency of molecular collisions is related to pressure and hence the total all a within and related to pressure except at extremely low pressure. B- I using of lines beyond the widths attributable to natural widths and the toppler effect is usually referred to as "pressure broadening", a term which includes effects attributable to collision lamping, to pressure shifts, and to Stark effects produced by "close approaches" between molecules. The present work deals primarily with studies of collision broadening, since the reast of pressures does not extend to the enormor pressures required to produce saymmetry shifts, Stark shifts, and enforced dipole effects. Collision broadening is usually encountered when the times between collisions are large compared with the times spent in collisie; the other effects are encountered when time spent in collisions is large and involves radiation absorbed during collisions.

# A. LORENTZIAN LINES

Although the experimental results obtained in the present investigat, have not dependent upon assumptions as to the actual shapes of spectral fixes, it is convenient for purposes of discussion to assume a line shape first proposed by Lorentz and round to give a satisfactory approximation actual line shapes over the fainly wide range of pressures in which collistin damping or collision brackening is encountwied. The exprestion for the assumption of limit proposed by pressure is

$$x(v, \frac{s}{v}, \frac{\alpha}{(v-v_0)^2 + \alpha^2}$$
 (5)

where the line atrength  $S = \int_{-\infty}^{\infty} k(v)dv$  and x is the halfwidth of the line at half the maximum value of K(v). The line etremeth of a given nollistion-breadened line is a constant for a given temperature and is relabed to probabilities of transitions between the initial and final energy levels of the absorbing molecule and to the relative populations of these levels. The half-width  $\alpha$  is related to the pressure, the absolute temperature, and the types of molecules in the ras sample.

Ladenberg and Reichelb have shown that with the value of k(v) in (5) substituted in (5) the integral

$$\int A^{*}(v)iv = \int \left(1 - e^{-2i(v)v}\right)dv \tag{5}$$

can be evaluated in terms of 0, 3 and w provided the limits of integration include the entire absorption associated with a given line. Thus, on the assumption that with appropriate limits of integration  $\int A'(v)dv \approx \int A(v)dv, \text{ it is possible to obtain majorithmental tests of ladouberg and Reiche's theoretical results. It is results indicate that$ 

$$\int\!\! A(v) dv = 2\pi i f(x)$$

Where x = Sw/ke, and

$$\int (x) + i e^{-x} \left[ \sigma_{ij}(tx) + i \sigma_{ij}(tx) \right] \tag{7}$$

where  $I_{ij}(\{1,i\})$  and  $I_{ij}(\{1,i\})$  are the brestle functions of order 0 and 1, respectively.

Although Eq. (7) is a rather involved expression, Eleaser? has all pind two approximations which we have dealing much to determine the model absorption of single lines under contain limiting conditions:

$$\int \hat{h}(v)dv = Sv \qquad \text{for small } x \quad \text{(weak lines)} \qquad (8)$$

433

Equations (3) and (3) approximations that the espectation of the settlefectority resisted to manifestic quantities that have to measured.

As indicated earlier, the line strength S of a collision-broadened line is a constant at a given temperature. The absorber concentration as a given temperature is airectly proportional to the product of the partial pressure p and the absorption path length: word. The relation between the helf-width and the pressure, temperature, and composition of the sample is somewhat more involved.

The half-width G of a collision-broadened apentral line is proportional to the collision frequency F. From kinetic theory it can be shown that the half-width is given by the relation

$$\alpha = \frac{y}{2\pi} = \frac{1}{4\pi} \frac{1}{2} N_1 \left[ u_{0,1} \right]^2 \left[ 2n\kappa^2 \left( \frac{1}{N_0} + \frac{1}{N_1} \right) \right]^{\frac{1}{2}/2}$$
 (10)

where  $E_1$  is the number of molecules of the 1-th type per unit volume,  $D_{n,1}$  is the sum of the optical collision diameters of the absorbing molecule and a molecule of the 1-th type, k is Boltzman's constant, T is the absorbing tolerale, and  $E_1$  is the mass of a molecule or the 1-th type.

The dependence of a on pressure F can readily be determined for a given mixture of games in which the ratios of the partial pressures of the components are the same for all values of the total pressure. Recalling that the partial pressure py of one of the i-th type is given by py = Bylf and assuming that collision diameters are independent of temperature, one can readily show that for a given mixture of games the half-width of a collision-broadened absorption line can be written

$$\alpha = \omega_n(P/P_n)\sqrt{T_n/T} \tag{11}$$

where  $\alpha_0$  is the half-width of the also under standard conditions  $\Gamma_0$  and  $\Gamma_0$ ; at a given temperature, the half-width  $\alpha$  would be expected to be directly proportional to the total pressure  $\Gamma$ .

Substitution of the value of a given by (11) in relations (8) and (9) leads to the following predictions for a given mixture of games at a given temperature:

Equations (0') and (p') account a very simple form for the special case of a sample consisting of a single absorbing gas with no admixture of other gases, since for this special case g a wand the equations become

$$\int A(x)dx \cdot SY \cdot SY = Px \cdot Smil(x) \quad \text{(weak line)} \qquad (87)$$

The expressions of total absorption , (Non in (5') and (5') apply to the very bladued situation in which a given mixture of gases is studied at the continuous continuous of total absorption of sixtures of different proportions of absorbing and non-absorbing gases. The appearance of D<sub>a</sub> , and M<sub>b</sub> in (10) indicates that different gases have different broadening effects on spectral lines. Since most of the gas samples employed in the experimental work constated of binary mixtures of absorbing gases and nitrogen, which was used as a nonabsorbing "broadening gase", it is desirable to consider (10) for the special case of a binary mixture of an absorbing gas and a broadening gas b. The summation in (10) will in this case have two terms; a self-broadening term in which is a, and a second term for the effects of broadening by the inert gas is b. Equation (10) assumes the form

$$2 = \frac{1}{4\pi} (2\pi h^2)^{1/2} \left\{ H_b(\Omega_{b,a})^2 \left[ \frac{1}{H_b} \right]^{1/2} + H_b(\Omega_{b,b})^2 \left[ \frac{H_b \cdot H_b}{H_b \cdot H_b} \right]^{1/2} \right\}$$
(101)

substitution for Ma and My in terms of the partial provides on Ma and My gives

$$1 = \frac{1}{k_T} \cdot \frac{2\pi^{-1/2}}{kT} \left\{ c_{n,A} \, \, p_n + c_{n,b} \, \, p_0 \right\} \tag{12}$$

where  $C_{n,n}$  and  $C_{n,n}$  are constants involving the optical collision dismeters and masses of absorbing and breadening gases. Recalling that the total pressure  $P = p_0 + p_0$ , one may the constant of P and  $p_0$  by rewriting (12) in the for .

$$Q = \frac{\lambda}{4\pi} \left( \frac{1}{kT} \right)^{-1} \left( \frac{1}{2a_1b_1} \cdot \frac{1}{2a_2b_2} \right) \cdot \left( \frac{1}{2a_2b_2} - \frac{1}{2a_2b_2} \right)$$

$$= \frac{\lambda}{4\pi} \left( \frac{1}{2a_2b_2} - \frac{1}{2a_2b_2} \right) \cdot \left( \frac{1}{2a_2b_2} - \frac{1}{2a_2b_2} \right)$$

$$= \frac{\lambda}{4\pi} \left( \frac{1}{2a_2b_2} - \frac{1}{2a_2b_2} \right) \cdot \left( \frac{1}{2a_2b_2} - \frac{1}{2a_2b_2} \right)$$

$$= \frac{\lambda}{4\pi} \left( \frac{1}{2a_2b_2} - \frac{1}{2a_2b_2} \right) \cdot \left( \frac{1}{2a_2b_2} - \frac{1}{2a_2b_2} - \frac{1}{2a_2b_2} \right)$$

$$= \frac{\lambda}{4\pi} \left( \frac{1}{2a_2b_2} - \frac{1}{2a_$$

where B  $\omega_{B,B}/d_{B,B}$  is called the self-croatening coefficient of the absorbing gas and represents the ratio of the "self-broatening salitity" of the absorbing gas to the "broadening ability" of the normalsolating at the "broadening ability" of the normalsolating at the Mathoda of determining a for various absorbing gases relative to the value be described later. The term in brackets in (1) is called the equivalent pressure  $T_{B}$  of the gas cample,

$$P_{p} = P + (B - 1)p$$
 (14)

where P is the measured total pressure of the sample and p is the partial pressure of the absorbing gas. The predictions embodied in (8') and (9') should be fulfilled by a mixture of gases provided P is replaced by appropriate equivalent pressure P<sub>p</sub>. It should be remarked that the theory presented thus far might be expected to apply to likes whose width is determined primarily by collision broadening; it is not applicable to extremely now pressures, for which natural widths and Doppler broadening are of primary importance, nor to extremely high pressures, at which absorption during collisions or "close approaches" leads to Stars effects and asymmetrical pressure shifts.

# B. ARSORPTION BANDS

In order to extend the theory to include total absorption of entire bands it is necessary to consider the spacings and intensities of the individual lines. If there is overlapping of the spactral lines so that more than one line contributes to the absorption at frequency v,  $A^*(v)$  is still given by  $m_1 \in \mathbb{N}$  if the value of k(v) used is the sum of the k(v) for each of the individual lines. Equation (6) also yields the total absorption of an entire tend if the proper value of k(v) is used. For the simple case of no overlapping of the lines, the total absorption of a band can be treated as the sum of the total absorptions of each of the lines.

In order to make theoretical calculations of the dependence of total absorption of an entire band on pressure and absorber concentration, it is usually necessary to make some assumptions about the spacings and intensities of the individual lines. Two different idealized band models have received general recognition:

- (a) A band consisting of an infinite array of equally intense, equally apaced lines has been treated by Eluassics;
- (b) A disordered bond, usually referred to as the stational mode, which is composed of lines where positions come at random and where in tensities can be represented by a probability distribution, he have worked out independently by Cooly and by Mayer.

Proof one recently discussed in deball where two models as well as a third-

(c) The limits Elsauser model which is composed of several groups of lines each of which individually forms on Elsauser band, but where these groups are superposed to give a nearly landom spacing.

The report by Plass also includes curves which relate absorption to absorber concentration and pressure for the three different band models.

Colourse, no real band is examily represented by any of these inerversal models. But it we become the in the risk report that at its the inerval dependence of the total absorption on y and Pe for CO2 ands can be explained by the Eleasser model\*. It was also demonstrated that the absorption by H2O bands is similar to that of a statistical band model. It is not subprising that the absorption by CO2 is best explained by an Eleasser band, since the changes in the intensities, half-widths, and spacings of the lines are gradual over the bands. The intensities and spacing of the H2O lines vary almost at random over the band, and the absorption is therefore similar to that of the statistical model.

Exact theoretical expressions for total absorption of even the idealized band models are complex. However, Plassy has discussed three different approximations for each of the three band models listed above. Each approximation permits the absorption to be expressed in terms of elementary functions and is valid over a considerable range of pressure and absorber concentration. These three approximations are: (i) strongline approximation; (ii) weak-line approximation; and (iii) non-overlapping approximation. Mothods (i) and (ii) are valid regardless of whether or not the spectral lines overlap, and method (iii) is valid whether or not the absorption lines are weak or strong.

(i) The strong-line approximation is will when there is virtually complete absorption near the centers of the strongest lines, and an increase in either v or  $P_e$  results in increased absorption in the wings of the lines only. The requirement for the validity of the strong line approximation is given in Eq. (9) and corresponds to large values of v and small values of  $P_e$ . For the case of no overlapping of the lines the total absorption of a band can be treated as the sum or the total absorptions of all of the lines, and for strong lines the total absorption of the band is seen from Eqs. (9), (13), and (14) to be proportional to  $(vP_e)^{1/2}$ . The range of values of v and v are ever which this relationship is valid is referred to as the square-root region. When there is everlapping of the lines the total absorption no longer follows a simple square-root dependence; however, the

<sup>\*</sup>The definition of Pc in this report differs from the equivalent presence employed by abd.

total absorption is still a function of the product wie. The distinguishing feature of the strong line approximation is that the total absorption can be expressed as a function of the single variable wie. The distinction between the square-root dependence encountered for strong in-overlapping lines and the more general strong-line approximation with reference to entire bands is emphasized; the square-root dependence can be considered as a special case of the strong-line approximation, and is wild only when the spectral lines do not overlap appreciably.

- (ii) The weak-line approximation is valid when the absorption is small at all frequencies in the bend, including those near the centers of the strongest absorption lines. The requirement for the validity of the weak-line approximation with reference to an entire band is similar to the requirement for the weak-line approximation for a single line given in Eq. (8). This requirement is satisfied in the case of small values of y and large values of P<sub>d</sub>. The total absorption of a given band is a function of y only; however, the simple linear relationship given by Eq. (8) is not valid in the case of over-lapping.
- (iii) The non-overlapping approximation is valid when there is not appreciable overlapping of the lines; and the total absorption is given by Eq. (7) with the contribution of all the individual lines added together.

It is possible to test the usefulness of a given band model by comparing the results of total absorption measurements with the theoretical predictions for ranges of values of w and Pe over which the strong-line approximation is expected to be valid. The strong-line approximation is useful for testing different band models since the total absorption depends on the distribution of the strengths of the lines as well as upon the errangement of the spectral times, and is quite different for different band models. Economy, the weak-line approximation and the non-overlapping approximation cannot be used to test different band models since the total absorption is virtually independent of the line specing and the distribution of the strengths of the lines.

Of the gases treated in the present report, 00 and H20 are linear molecules whose absorption bands are composed of lines having gradually changing intensity and spacing. The structure of these bands is best approximated by an Elasaser model; and is is therefore useful to compere the N20 and C0 data in the present study with results based on the Elasaser theoretical model for values of y and Pe for which the strongline approximation would be expected to be valid.

It will be shown below in the discussion of each of the bands inwolved that the Elsasser model is sufficient to predict, in a general way, the dependence of total absorption on absorber concentration and pressure. However, it will be seen that the experimentally observed effect of increasing pressure is significantly less than that predicted for an Elsasser band for certain values of y and  $P_{\rm e}$ .

## C. BAND INTENSITIES

One quantity of fundamental importance in studies of molecular spectra is a quantity called band strength or hand intensity,  $\int k(\nu) d\nu$ , a quantity analogous to the line strength defined in connection with (5) except for the fact that the limits of integration are chosen to include an entire band. If  $\int k(\nu) d\nu$  is known for a vibration-rotation basic information concerning transition moments appropriate to transitions between vibrational levels can be obtained; the value of  $\int k(\nu) d\nu$  also depends upon the population differences between the lower and upper levels and is therefore also a function of temperature.

From the data on total absorption to be presented later in the report, it is sometimes possible to determine a value of band intensity for the various bands to be investigated. This quantity can be evaluated from total absorption measurements provided  $k(\nu)\ll 1$  for all frequencies included in absorption. Under this condition, an integral similar to (6) but with the limits of integration chosen to include an entire band becomes

$$\int A(\nu) d\nu = \nu \int k(\nu) d\nu \qquad (6')$$

for  $k(v)w\ll 1$  throughout the entire band. Since  $\int A(v)dv$  and w can both be measured, the band intensity  $\int k'(v)dv$  can be determined. The choice of absorber concentrations and pressure for which  $(6^{\circ})$  applies will be discussed later in some detail:

#### III. EXPERIMENTAL METHODS

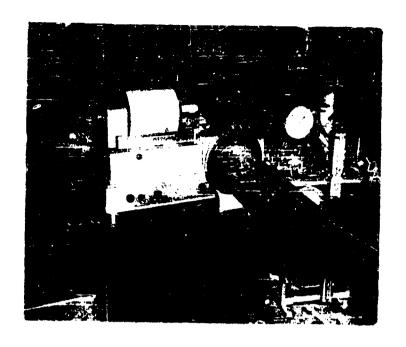
A. SPECTROGRAPES, ARSORPTION CLILS, AND GAS-HARDLING EQUIPMENT

In the present study, a double-beam, single-pass Perkin-Elmer Model 21 spectrometer with NaCl prism was employed for frequencies between 4000 cm<sup>-1</sup> and 670 cm<sup>-1</sup>. This spectrograph was equipped with a drive mechanism and recorder which provided a direct plot of A(v) vs. v on the recorder chart. In Fig. 1 are shown two views of the Model 21 with dual long-path cells designed and nonstructed specifically for it by Perkin-Elmer. In the upper part of the figure the sample cell housing has been removed to show the multiple-reflection system which utilizes the Whitel's principle. The two cells, which extend from either side of the spectrometer, are identical and are arranged so that they may be used simultaneously, one in the sample beam and one in the reference beam. In the lower part of the figure the sample cell housing is shown in place; the cell housing is designed to hold vacuum and gauge pressures up to 100 pounds per square inch. A portable gas-handling system which was constructed for the present investigation is shown in both views and is discussed below.

Each cell has a base path of one meter, and can be adjusted to ho traversals or any smaller number divisible by four. The cell housing must be resort in order to adjust the number of traversals. Although electrical heating coils around the cell housing are capable of heating the cells to approximately 100°C, the data included in the present study were obtained with the samples near 30°C.

Very little effort is involved in changing from the dual long-path cells to conventional sport sample cells, or vice versa. In order to use the long-path cells a set of kinematically mounted optics fits into the sampling space between the cells and directs the radiation into the cell and from the cell to the monoton optics. These optics, which are not visible in Fig. 1, can be easily removed when short cells are to be used.

The main body of the gas-handling system shown in Fig. 1 and also the puckless ministure valves were rade of AISI 316 stainless steel which would withstand high pressures and would also allow the possible future use of many gases which are corrective to carbon steel and nonzerrous metals. One of six outlet valves was connected to the sample cell, one to the reference cell, and any of the other four could be connected to cylinders of gas being used. The pressure in the system was recorded by one of four different gauges designed for different pressure ranges. A Wallace and Therman absolute pressure gauge was used in the pressure range between 1 and 50 mm Hg. This gauge is calibrated from 0 to 50 mm Hg and is designed to have an accuracy of one part in you provided the "zero" is properly adjusted. A Molecc' gauge was used for absolute pressures less than 1.5 mm Hg and also served as a check of the readings



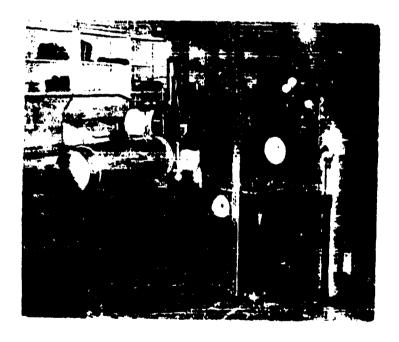


Fig. 1. Two views of the Perkin-Elmer Model 21 Spectrometer with dual long-path cells and gir-heading system

of the Wallace and Tiernan gauge at low pressures. A simple "U"-tube manameter served to measure pressures in the range from 50 mm Hg  $\infty$  one atmosphere. An Ashbroft "Duragauge", which is a Bourdon tube-Tuppe pressure gauge, was used in the higher pressure range. The Ashbroft gauge is also constructed from AIST 316 stainless steel and is calibrated for gauge pressures between 0 and 100 pounds per square inch in steps of one pound per square inch.

Before introducing a sample into the sample cell, the entire gashandling system was evacuated, the reference cell was closed off, and the absorbing gas was added to the system and to the sample cell at the desired pressure. The valve to the sample cell was then cicsed. In order to produce a sample with the same absorber concentration but greater total pressure, mitrogen was added to the sample. Before the mitrogen was introduced, the gas-handling system (excluding the sample cell) was evacuated in order to prevent more absorber from being forced into the cell. Mitrogen was then added to the system to a pressure slightly greater than that in the cell before the valve was opened; more nitrogen was then added to the sample cell until the desired pressure was reached. This procedure was followed in order to reduce the flow of absorber from the cell back into the rest of the system. To produce a sample of still nigher total pressure, more nitrogen was added to the system, without evacuating it; the valve between the main system and the sample cell was then opened and nitrogen added to the desired pressure. This procedure could be repeated several times to measure the total absorption of a series of samples having the same absorber concentration and varying total pressure.

In order to reduce the chance of the absorber diffusing from the sample cell back through the valve, the valve was kept open for the minimum time required to allow the pressure to come to equilibrium. It is probable that most of the absorber which did in some way get out of the cell would remain for several minutes in the 3-foot length of small tubing which connected the cell to the rest of the system. When more nitrogen was added, most of this absorber would be forced made into the cell, and any losses of this type would not be cumulative. It was found that the observed absorption of a sample of for absorber partial pressure with nitrogen added to a high pressure in one step agreed wall with a similar accepted which the introgen was added by several successive steps, thus indicating that leskage of the absorber through the valve was negligible.

When samples containing two gases were being studied, it was essential that the gases be thoroughly mixed to ensure that the proper amount of absorber was contained in the radiation path. One might axpect that as nitrogen was being added to a sample, the absorber would tend to move to the part of the cell away from the entrance; and, until the gases became uniformly mixed, by thermal monion or otherwise, a false absorption reading would be obtained. It was assumed that a uniform gas mixture had been obtained when the

reached a constant value. The mixing problem was investigated for each of the different absorption cells used.

A fan was been built in the retwin-Elmer long-path cells to speed up mixing. Without the fan several minutes are required for mixing of samples at high pressures; however, only a few seconds are required for mixing when the fan is used. The fan blade is connected to a shaft which extends through a rotating vacuum seal to a small electric motor outside the cell. The motor, which is mounted above the sample cell, can be seen in the lower portion of Fig. 1. The necessity for forced mixing probably arises from the fact that the cantilever, which book one of the mirrors inside the cell and is visible in the upper portion of Fig. 1 divides the cell into two chambers. Difficulty from one of these chambers to the other apparently occurs rather slowly.

Figure 2 is a drawing which illustrates the construction of two short cells which were used, one 1.55 cm and the other 6.35 cm long. The synthetic buna-N'C' ring between the window and the cell body serves as the seal, and the windows are held in place by brass end plates joined by threaded brass rods. The grooves for the "O" rings are not deep enough to allow metal-to-window contect; this reduces any local stresses caused by irregularities in the windows or metal which might tend to crack the windows. An "O" ring between the end plate and window serves as a cushion and is directly opposite the "O" ring seal. This makes it possible to hold the windows on firmly without producing appreciable shearing stress. The short cells can be used for vacuum and for gauge pressures up to 60 psig. Grooves were mitted in one endplate which fits into a receptacle provided on the spectrometer directly in front of the entrance slit. Two NaCl windows were put in the reference beam to correct for the attenuation by the cell windows.

It was found that mixing was essentially complete in the short cells within a few seconds. However, a few minutes were required for mixing in a somewhat similar cell different from the type shown in Fig. 2 in that the valve was joined to the main body of the cell by approximately two inches of tubing. The valve was soldered directly to the cells used in the later work. It is probable that the rather long mixing time for the first type of cell was caused by the fact that diffusion through the small-bore tubing to the value took place very slowly.

Studies of the 4260 cm<sup>-1</sup> CO band were made with a Perkin-Elmer Model 99 spectrometer with a LiF prism. A Normat glover served as a radiation source and a thermocouple was used as the detector. The signal was amplified by a Perkin-Elmer Model 107 amplifier and the spectrum displayed by a Leeds and Northrup Speedomax Recordar. A multiple-pass cell with a White-10 option; system and external adjustments was used with the Model 99 spectrometer. The base length of the cell is 1.5 meters and paths as long as 40 meters were used, although paths twice this great can be obtained. The body of the cell is steel, the end plates aluminum, and the windows NaCl. Total pressures are limited to atmospheric and less.

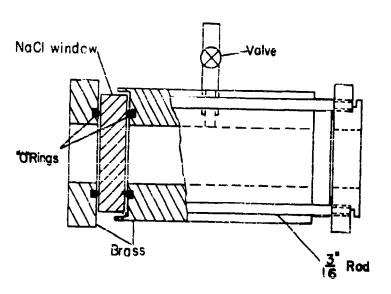


Fig. 2. Design of short absorption cells

Absorption measurements with the multiple-pass certimed in the study of the 4260 cm<sup>-1</sup> CO band indicate that mixing is complete within several seconds after nitrogen is introduced; therefore, forced mixing was not necessary.

## B. MEASUREMENT OF TOTAL ADSORPTION

The total absorption of the samples studied with the Model 21 spectrometer was easily determined by measuring the area under each recorded spectrum with a compensating polar planimeter. The abscissa was linear in frequency v and the ordinate was linear in fractional absorption A(v); therefore the area under the absorption curve was proportional to the total absorption. In some cases slight corrections had to be made for scattered light and for differences in deflection corresponding to zero absorption. Such differences were caused by unequal attenuation in the sample and reference beams by the instrument and by the atmospheric green within the instrument. These effects were easily corrected so that negligible error in the measurement of total absorption was caused by them.

Spectra of the 4260 cm<sup>-1</sup> band obtained with the Model 99 spectromster were replotted before being measured. An "envelope" spectrum was obtained with the cell evacuated, and was superimposed on each sample spectrum. The envelope spectrum served as the zero percent sheorytion curve and the spectrum was replotted on a scale whose ordinate was linear in fractional assorption. The area under the replotted curve was then measured and multiplied by an appropriate "normalizing factor" to determine the total absorption. This process of replotting the spectra has been discussed in detail by HEM<sup>2</sup>.

# C. ERRORS AND ACCURACY

The possible sources of error in a study such as this may be divided into three categories:

- errors in sampling of the gases (in determining the amount of absorbing gas and nitrogen in the cell),
- (11) errors arising from inaccurate recording by the apentrometer due to notes, drift, and nonlinearity in the detection-amplification-recording system;
- (iii) errors in measuring the recorded spectra.

#### 1. Sampling Errors

The major source of error in sampling is the effect of electrotion of the gases on the cell walls; the problems involved are discussed in great detail in Appendix I. Another source of sampling error involves the accuracy with which the pressures of the samples were measured. With the system or four gauges mentioned above it was possible to measure pressures greater than 10 mm Hg to 15 with a somewhat greater error for smaller samples. Improper mixing of the absorber with nitrogen would give rise to error; however, this error is believed to be negligible, as mentioned earlier, since care was taken to ensure sufficient mixing.

Errors in the recorded spectra hight arise from absorption by an impurity in the same spectral region as the absorption band being studied. Small amounts of CO<sub>2</sub> which were present in the H<sub>2</sub>O and CO gases produced alight absorption on the high-frequency side of the 222½ cm<sup>-1</sup> H<sub>2</sub>O band and the 21½ cm<sup>-1</sup> CO band. By recording spectra of CO<sub>2</sub> samples having approximately the same absorption as the CO<sub>2</sub> impurity, it was possible to estimate the CO<sub>2</sub> absorption and to correct the spectra in the region where both the absorbing gas and the CO<sub>2</sub> produce absorption. A similar correction for absorption by H<sub>2</sub>O impurity in the CH<sub>1</sub> was also made. This type or correction was necessary only for samples—large absorber concentration and was easily made to yield an error of 1 to than 15 in the total absorption except in the case of the 1550 cm<sup>-1</sup> Cd<sub>4</sub> band, for which the error is somewhat larger. No absorption by impurities in the region of the CO 426O cm<sup>-1</sup> band was found.

The approximate purities of the gases used in the present study are listed along with the major impurities in Table 1.

# 2. Recording Errors

Errors in recording the spectra erise from the usual noise and from other variations due to faulty instrument operation. These variations can cause a misinterpretation of the apparent location of the zero absorption curve. Errors of this nort are particularly noticeable in spectra with low total absorption and are estimated to be  $\pm i \%$  for total absorption less than 10 cm<sup>-1</sup>. The percentage error decreases for greater absorption and is approximately  $\pm 2\%$  for values of total absorption greater than 10 cm<sup>-1</sup>.

## 5. Errors in Measurement of Records

Repeated measurements indicate that a careful operator can measure the erea under an absorption curvo with a precision of less than x.s.

It is estimated that the over-all error does not exceed \$105 for values of total absorption less than 10  $\rm cm^{-1}$  and is less than 255 if the total absorption is greater than 10  $\rm cm^{-1}$ .

### D. UNITS

In the present study all values of partial pressure, total pressure, and equivalent pressure are expressed in units of mm Hg absolute. Tabulated values of partial pressure of the absorbing gas represent the partial pressure of the gas as it comes from the cylinder without correcting for the purity.

Values of absorber concentration  $\underline{v}$  are expressed in terms of a unit called the atmos cm, and are given by the product of the absorber path length in cm and the partial pressure of the absorbing gas in atmospheres, corrected to standard temperature. The purity of the absorbing gas is taken into account in calculating values of  $\underline{v}$ .

Since values of absorber concentration are expressed units of atmosem, it follows from (6') that the units of band intensity  $f(\mathbf{k}(\mathbf{v}))$  are atmost cm<sup>2</sup>. It is emphasized that values of band intensity are based on values of  $\underline{\mathbf{y}}$  which were corrected to standard temperature. The observed values of  $f(\mathbf{k}(\mathbf{v}))$  will be compared later to results of other workers which are expressed in units of atmost cm<sup>2</sup> sec<sup>-1</sup>. When these units are used the numerical value differs by a factor of  $\underline{\mathbf{c}}$ , the speed or light.

### E. DETERMINATION OF SELF-ERGADENTIS CONFECTIONS

In Section II the concept of equivalent pressure was introduced and was expressed in (14) in terms of a self-broadening coefficient B, which can be expressed as

$$\underline{\mathbf{p}} = \left[\frac{2M_b}{M_b + M_b}\right]^{1/2} \begin{bmatrix} D_{B_1,B_2} \\ D_{B_1,b} \end{bmatrix}^2$$

in terms of the symbols introduced earlier. The experimental determination of B involves a study of the transmission of radiation through cells of different length in which the absorber concentration v is the same. The sample in the short cell, called the "reference well", consists of a pure sample of the absorbing was at pressure p; for this cell pref w pref and the equivalent pressure is given by

The sample in the religion religional the "craph religions at the partial pressure programmed to produce an absorber concentration y equal to that in the reference cell. Initially, the absorption of radiation in the sample cell is considerably less than that in the reference cell, wince the partial pressure and total pressure are less than those in the shorter reference cell. Various amounts of nitrogen are then added to the reference cell corresponding to equivalent pressures.

$$P_{c} = p_{b} + B_{c} \quad \text{(Sample Cell)} \tag{17}$$

where  $p_b$  is the partial pressure of the broadening gas. For the situation is which the absorptions in sample and reference cells are most nearly equal, it is assumed that the values of  $P_c$  in (16) and (17) are equal. The value of the self-broadening coefficient can then be determined from the relation

$$B = \frac{P_0}{V^{ref} - V} \tag{18}$$

where  $p^{ref}$  is the partial pressure of the absorbing gas in the short reference cell and p and  $p_{h,r}$  respectively, are the partial pressures of the absorbing and broadening gases in the longer sample cell when the absorptions in the two cells are equal.

The value of B was determined for all except the 4060 cm<sup>-1</sup> CO band by making use of the double-beam feature of the Model Pl spectrometer, with the 6.55-cm cell in the sample beam and the 1.55-cm cell in the reference beam. The method used to determine B for the 4260 cm<sup>-1</sup> band in described below in the lineussian of the results obtained for this band. The same absorber concentration was added to both cells by making the retic of the partial pressures of absorbing gas in the two cells inversely proportional to the cell lengths. The pressure in the sample cell was then increased in small intervals by adding H<sub>2</sub> and the spectrum was scanned for each sample cell T(sam.) to the transmission of the sample cell T(sam.)

Typical spectral comparisons which were obtained in this manner for the 2224 cm<sup>-1</sup> N<sub>2</sub>O band are shown in part I of Fig. 5 for various values of total pressure in the sample cell. The dashed line at T(sam.)/T(ref.) = 1.00 represents the recorder pen position when the two beans were "balanced": 1.e., the transmissions of both samples were the same. It is seen that the initial of the transmissions decreased with increasing pressure in the sample cell as would be expected, and that same of the spectra occur above

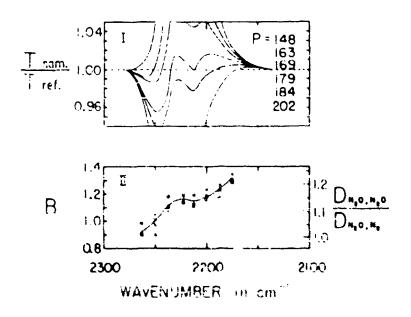


Fig. 3. The self-breadening coefficient 2 for the P224 cm<sup>-1</sup> M<sub>2</sub>O bank. Part 1. Specims obtained with v = 0.30 atmos on in both reference and sample cells, total pressure in reference cell = 164 mm Mg. Total pressure in sample cell for each specimum in indicated. Part II: Plot of 8 and collision cross section ratio. The four different sets of points correspond to the results of the specima shown in Part I as well as to three other exts of samples maring different values of absorber concentration.

the balance line for some frequencies and below the line for others. Therefore the total pressure corresponding to balance was not the same for all frequencies. The pressure corresponding to balance for any . Iven frequency was determined by interpolating octween the two successive curves which occur on opposite sides of the balance line.

Since the value of p, corresponding to equal transmission of the cells was found to vary with frequency, it is apparent then the value of B varies throughout the band. Points corresponding to the experimentally determined value of B were plotted against frequency in Part II of Fig. 5.

The ratio of collision cross sections is shown by the right-hand scale in Part II of Fig. 3. This quantity is included only as a matter of interest and for future comparison. It is related to the more important quantity, B, by Eq. (15).

The present study deals primarily with total absorption rather than with the absorption at certain frequencies. It was therefore desirable to obtain a nominal value of B which could be used for an entire bond. The nominal value was arrived at by taking a "weighted average" value of B over the frequency limits of the band, with extra weight given to the frequency ranges which contribute most of the absorption. The values of B determined for each band, which are listed in Table 2, are Unlieved accurate to 10% as an "effective" value for the entire bands.

The experimental results reported in Sections V - VII indicate that the total absorption never has a dependence on equivalent pressure  $F_0$  which is greater than approximately  $P_0^{0.45}$ . Therefore a relative error of 66 in B would not give rise to a relative error greater than 2 or 36 in the calculation of total absorption from the espirical equations to be presented later. This maximum error would occur only for samples containing the absorbing gas alone; the error would decrease for samples containing  $N_2$  and would vanish in the case of infinite dilution. The nominal values of B given in Table 1 are therefore sufficiently accurate for the present purpose, which is to make it possible to account for the different broadening ability of different games.

The final report on this research program will contain detailed information on the dependence of B with frequency for each of the absorption bands discussed in the present study, as well as for other absorption bands discussed in subsequent reports. Information will also be given which concerns the relative broadening abilities of several inert gases; M2, O2, A, M2, and M3. Singleton<sup>11</sup> of this laboratory, as well as several earlier investigators, has found that unambiguous values of B can be determined and that the determined values are actually independent of pressure for the pressure range covered in this report. This result indicates that the effective collision cross sections of the molecules are ensentially independent of pressure over the ranges of pressure and for gases used in the present study.

It was shown in the HBW report that the broadening sbility of air is very nearly the same a that of N2. Therefore the values of equivalent pressure which are used in the present study would approximate the total pressure in an atmospheric path, provided the partial pressure of the absorbing gas is small in comparison with missopheric pressure.

#### IV. TREATMENT OF DATA

When the study of a given absorption band or region was started, it was found desirable to select some definite slit program for the spectrometer and to take all data with the same program. On the Model 21 spectrometer it is possible to scan the spectrum of a band with the slitwidth constant or to use a slit programming mechanism which automatically adjusts the plits continuously as the spectrum is being scanned. Although the total absorption  $\int A(v) dv$  is independent of alitwidth or of slit program, there are occasions when it is desirable to compare A(v)for some definite frequency for different samples; this would be impossible unless all data are taken with the same slit program. However, since the datum of primary importance in the present work is  $\int A(v)dv$ , it is not necessary to work with narrow slits; it is desirable to work with slits sufficiently wide to give large values of signal-to-noise ratio even for the largest values of total absorption. Another practical reason for the use of wide slits is that the process of planimetering spectral recordings is easier for bands not exhibiting rotational fine structure; for nearly every band studied, it would have been possible to obtain more detailed information concerning band shape had this information been important to the investigation.

Once a program for the slits had been selected, a large number of spectra were recorded. These spectra provided data on fA(v)dv for extremely wide ranges of absorber concentration u, partial pressure p, and total pressure p--including, of course, pure samples of the absorbing gas, for which case p = P. The values of total absorption for various cell lengths, absorber concentrations, equivalent pressure, etc., are erganized in tabular form for use in tests of theoretical predictions; it is hoped that these tabulations will prove useful to theorists.

One type of relationship important to nearly all current theories is the relationship between absorption and pressure. In order to establish this relationship from the data obtained in the present work, a set of logarithmic plots of total absorption fA(v)dv as a function of equivalent pressure  $P_e$  is then made for various values of absorber concentration v. From the alones of the curves, the relationship between fA(v)dv and  $P_e$  can be obtained, usually it is found possible to show that  $fA(v)dv \sim P_e^{\,n}$ , where a usually does not have the value of 0.5 that would be expected on the basis of (9) for a square-root approximation even in although where this approximation might be expected to apply.

Another set of logarithmic plots is then prepared giving  $JA(\nu)d\nu$  as function of absorber concentration  $\nu$  for various values of equivalent

pressure  $P_e$ . From the slopes of the resulting curves, it is possible to determine the relation between  $\int A(\nu) d\nu$  and  $\underline{\nu}$ . It is usually found that  $\int A(\nu) d\nu e^{i\underline{m}}$  where  $\underline{m} \triangleq 0.5$  in cases in which the square-root approximation might be expected to apply.

The next step in attempting to establish experimental relationships between  $fA(\nu)d\nu$  and sample parameters is to make logarithmic plots of  $fA(\nu)d\nu$  as a function of  $\nu P_e^{-2}$ . From the resulting plots it is frequently possible to obtain empirical equations relating total absorption to pertinent sample parameters for certain definite ranges of values of total absorption and equivalent pressure. It is usually possible to obtain empirical equations of the form

$$\int A(\nu)d\nu = c \left[\nu P_e^{a}\right]^b \tag{19}$$

where c, a, and b are empirical constants with b approximately equal to 0.5. Such equations are valid for intermediate values of  $\int A(\nu) d\nu$  and  $P_{\rm e}$  for which strong but not seriously overlapping lines might be expected throughout most parts of the band.

For large values of  $\int A(\nu) d\nu$  it is found that the central portions of the absorption band represent a spectral region of mearly complete absorption. Further increase in  $\underline{\nu}$  or  $P_{\underline{\nu}}$  results merely in growth of  $A(\nu)$  in the vings of the band. For this situation, the total band absorption can be expressed by equations of the form

$$\int A(v)dv = C + D \log (wP_e^{a})$$
 (20)

where C, D, and a are empirical constants.

It is rather difficult to derive satisfactory empirical equations for extremely small values of  $\int A(\nu) d\nu$  and  $P_{\phi}$ . Under certain circumstances  $\int A(\nu) d\nu$  is found to be nearly independent of  $P_{\phi}$ , a result indicating that Doppler broadening is the dominant influence on line width; under other conditions  $\int A(\nu) d\nu$  may vary linearly with  $\underline{\nu}$  as predicted by (8) for small values of  $x = S\nu/2\pi x$ .

Although total absorption is the quantity that is independent of slitwidth, it is sometimes desirable to compare the measured values of  $A(\nu)$  at certain points of interest in absorption bands in order to compare experimental results with predictions based on various theoretical band models. Measured values of  $A(\nu)$  near regions of maximum absorption in the P and R-branches are examples of regions where theoretical predictions can rose there by checked. It should be emphasized again that  $A(\nu)$  is dependent on effective slitwidth; comparisons of  $A(\nu)$  obtained with different executive slitwidths are meaningless.

The treatment of data and the development of empirical relations will be discussed in some detail in the following section—along with a comparison of observed results and theoretical predictions. Many details will be omitted in later Sections.

## V. EXPERIMENTAL RESULTS FOR THE 2224 cm<sup>-1</sup> N<sub>2</sub>O BANL

The  $v_3$  fundamental band of N<sub>2</sub>O which has its center at 2224 cm<sup>-1</sup> occurs in the "atmospheric window" which lies between the strong 2350 cm<sup>-1</sup> CO<sub>2</sub> band and the 1595 cm<sup>-1</sup> H<sub>2</sub>O band. The major portion of the absorption by N<sub>2</sub>O in this spectral region is due to the  $v_3$  fundamental of the common isotope N<sub>2</sub><sup>14</sup>Olb. However, there is absorption by the less common isotopes as well as by at least one weak band  $(v_1+v_2-v_1)$  described by Thompson and Williams<sup>12</sup>. For the sake of simplicity in the present report, all these bands together will be referred to as the 2224 cm<sup>-1</sup> N<sub>2</sub>O band.

Although the concentration of nitrous oxide in the earth's atmosphere near ground level is only approximately three parts in ten million<sup>13</sup>, this band gives rise to a major portion of the atmospheric absorption in the region where it occurs. A knowledge of the absorption of this band is therefore important in the calculation of the transmission of radiation in the "atmospheric window" in which it occurs. The contribution of this band to the heat balance of the earth's atmosphere is negligible since the major portion of the solar radiation occurs at higher frequencies and most of the radiation emitted by the earth's surface occurs at lower frequencies.

More than 200 spectra were obtained by using the double-beam Model 21 spectrometer with absorption cell lengths of 1.55, 6.35, 400, 800, and 1600 cm. Different combinations of No. and No. were used to provide absorber concentrations from 0.00016 to 76.4 atmos cm and total pressures from 1 to 3000 mm Hg. All spectra were obtained with slit widths corresponding to approximately 25 cm<sup>-1</sup> and with the cell temperature man 30°C.

### A. ABSORTION GURVES

Tracings of spectra of some of the samples having the largest absorber concentration and shown in the left-hand portion of Fig. 4. A slight growth of the band as a result of increasing the pressure with Eg is illustrated by these spectra; and the band limits, beyond which there is negligible absorption for even the largest samples, are indicated by the limits of the spectra. A few spectra of samples containing still larger values of absorber concountation were obtained but have not cann included since very little additional information could be obtained from them. It is shown below in the discussion of Figs. 12 and 14 that there is little increase in the total absorption of the band for larger

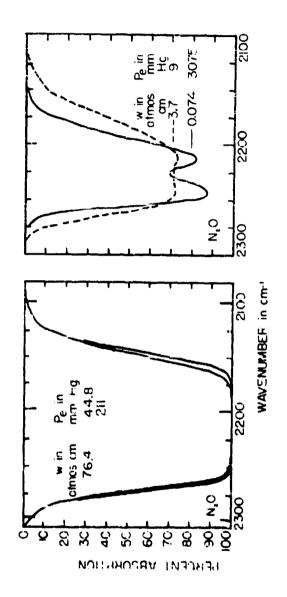


Fig. 4. Spectra of the 2224 cm 1 MgO band.

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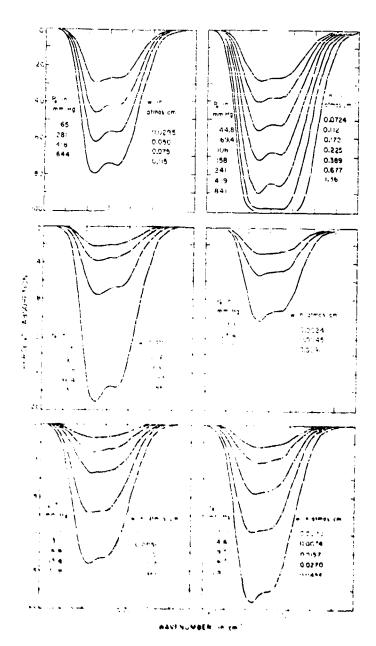


Fig. 8. Spectra of the 222 cark to 0 manage

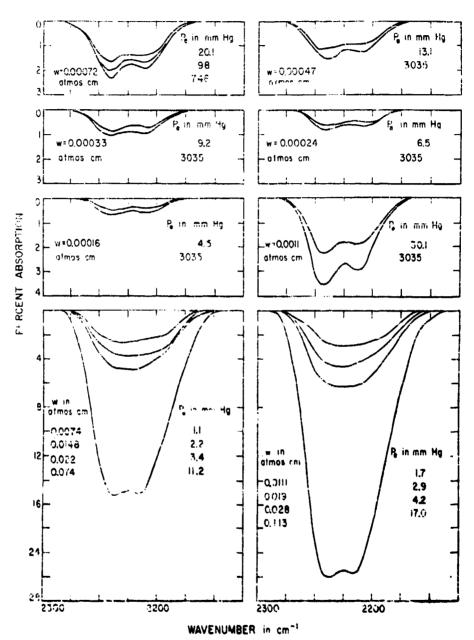


Fig. 6. Spectra of the 222%  $\rm cm^{-1}~N_{\odot}O$  band.

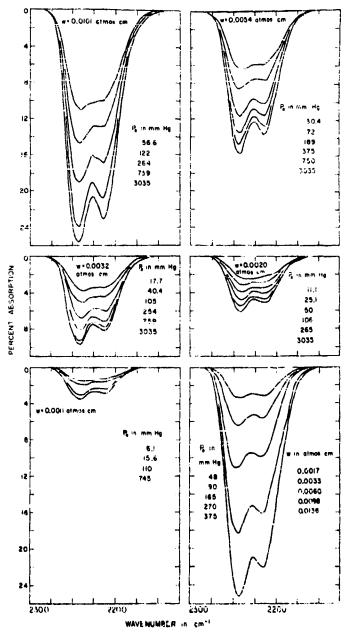


Fig. 7. Spenora of the 2224 cm-1 NoO band.

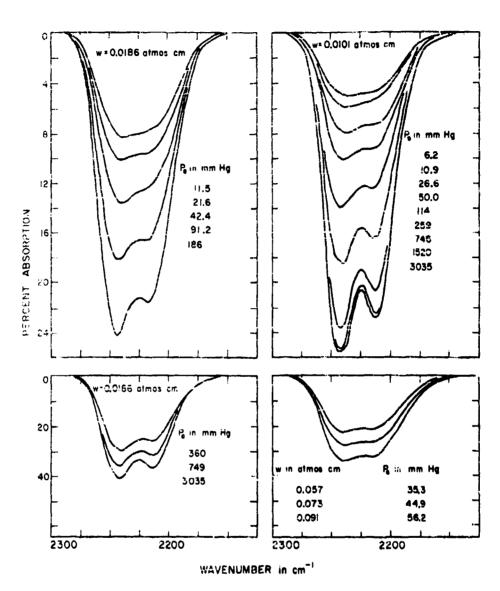


Fig. 8. Spectra of the 2224 cm<sup>-1</sup> N<sub>2</sub>O band.

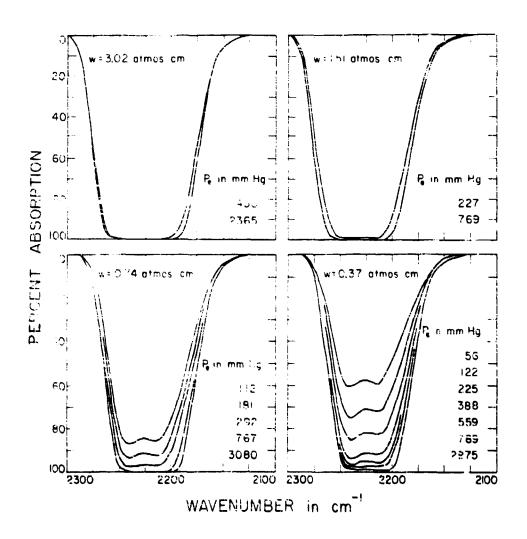


Fig. 9. Spentin of the 2224 cm<sup>-1</sup> N<sub>2</sub>O band.

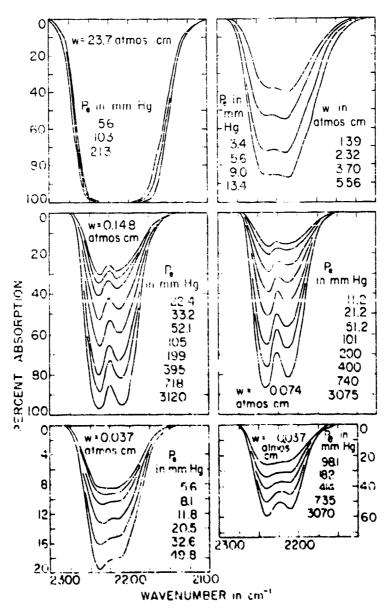


Fig. 10. Spectra of the 222% cm<sup>-1</sup> M<sub>2</sub>O band.

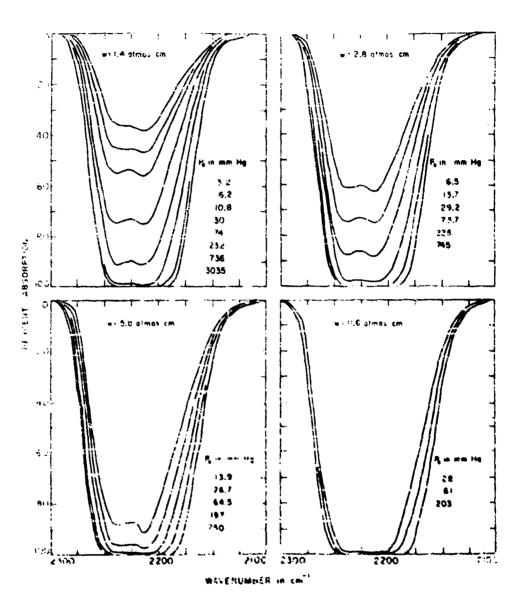


Fig. 4: Spectra of the Down at Nov mans.

Sections of many of the spectra have been omitted to prevent over-crowding of the figures, but as much of each spectrum as could be presented clearly was included. The percent transmission of the samples in the spectral regions where segments were omitted can apply to estimated by interpolation between adjacent absorption curves.

All of the experimental results obtained for the 2224 cm<sup>-1</sup>  $\rm H_2O$  band are summarized in Table 3. The absorber concentration  $\underline{v}$  and equivalent pressure  $\rm P_0$  for each sample are shown. The equivalent pressure  $\rm P_0$  was calculated from the total pressure  $\rm P$  and the  $\rm H_2O$  partial pressure  $\rm P$  by the simple equation

in accordance with Eq. (14) and Table 2.

Talums of the total absorption as well as the observed percent absorption at 2040 cm<sup>-1</sup> and 2213 cm<sup>-1</sup> are tabulated. These two frequencies correspond to the maximum absorption in the R- and P-branches, respectively. Remarks conserving the sampling procedure are included in the right-hand column of the table.

# B. TOTAL ABSORPTION AS A PUNCTION OF ABBORPUS CONCENTRATION UID SCITYALFUS PRINTING

May of the results summarized in Table I are plotten to a liquint their sends in Fig. 12 with much absorption above as a function of the equivalent pressure Po. Such curve corresponds to a given absorber concentration y each above the effect of increasing the equivalent pressure by adding No. The "path" refers to the leasth of the absorption cell. The points marked a which occur at the lower scale of surves A through P represent results obtained for samples of No alone. The other points on a given curve give results for samples containing the same amount of No., but with different amounts of No added. The pulses on the lower cash of curves Q. N., and S represent results for samples onde from previously mixed No. and No.; on these curves where are no points which represent total absorption due to No.0 alone.

The point of Survey D and E corresponding to high equivalent prescures are emitted from the figure for purposes of visitly. If curves D and E had been extensed to higher pressures they would be issued to be within 1% of the value of total absorption expected from an extrapolation to larger absolute except to the based on curves F. C and E.

there is a present therefore of the evenes with ingreeningly high presence, indication a decrease in the depositance of

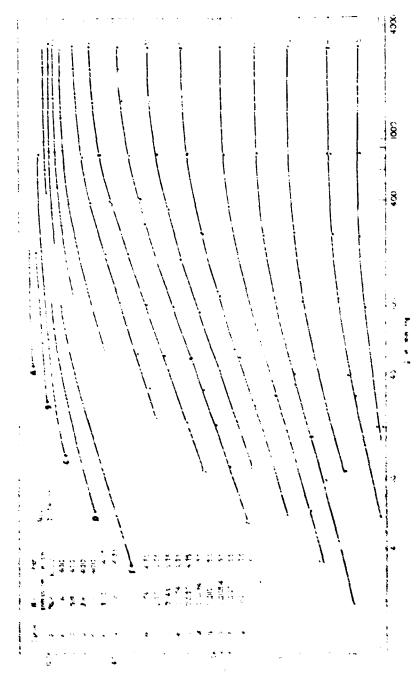


Fig. 12. The Coll Shoppition of the 222 cm<sup>2</sup> M<sub>2</sub>O bind matter a statement pressure for Laples laving command with more of Losofter confentation.

total absorption on equivalent pressure. This "saturation phenomenon" can be explained on the basis of the change in line shape with pressure. As the pressure increases, the half-width or a spectral line increases and the absorption coefficient  $k(\nu)$  increases in the wings and decreases near the center of the line. For pressures sufficiently high that the line half-width is approximately equal to the spacing d between the lines, the fine structure of a band diminishes and further increases in pressure have but little effect on total absorption. Goody and Wormell have shown that the half-width  $\alpha=d$  for  $N_20$  bands at a pressure of approximately 4100 mm Hg (5.4 atmospheres), where  $\alpha$  and d are necessarily averaged over the entire band. One would therefore not expect the total absorption to change appreciably with pressure in the higher-pressure ranges used in the present study.

It is seen from Fig. 12 that the value of  $P_0$  above which the total absorption increases only slightly with increasing pressure is greater for larger values of absorber concentration than for smaller values. In the case of large  $\psi$ , the decrease of  $k(\psi)$  near the line centers which occurs with increasing pressure does not appreciably decrease the absorption  $A'(\psi)$  at these frequencies since  $k(\psi)w$  is still smalleightly large that the absorption remains essentially complete at the line centers. However, in the case of small  $\psi$ , a lower pressure  $Y_0$  is sufficient to reduce  $k(\psi)w$  near the line centers to a value such that the decrease in absorption near the line centers with increasing pressure is approximately the same as the increase in absorption in the wings of the line. There is therefore little or no net increase in total absorption with increasing pressure.

Curves J through 0 have approximately the same slope for values of pressure below 200 mm Hg. The slopes of any of the curvesin this region have a maximum value of approximately 0.38 and an average of approximately 0.37. Thus it is possible to express the total absorption as some function of the absorber concentration we times  $P_0^{-0.37}$  for values w and  $P_0^{-0.37}$  represented by these curves.

That the variation of the total absorption with changes in  $\psi$  and  $P_{\psi}$  becomes vary slight for large values of  $\psi$  and  $P_{\psi}$  is illustrated by the crowding of the curves in the upper right-hand corner of Fig. 12. This result is due to the fact that  $k(\psi) v \gg 1$  for most frequencies within the hand, and hence the total absorption as given by Eq. (6), changes very slowly as  $\psi$  and  $P_{\psi}$  change. Inspection of the spectra of samples with large  $\psi$  and  $P_{\psi}$  such as those in Figs. 10 and 11, which are represented by the points in the upper right portion of Fig. 12, reveals that absorption is complete near the band center and that the greath of the band with increasing  $\psi$  and  $P_{\psi}$  is elight and occurs in the viage.

In Fig. 1) are five curves relating the total absorption to the equivalent pressure. The curves of Fig. 13 are different from those of Fig. 12 in that those in Fig. 1) do not correspond to constant values of absorber connentration. Curves  $A_1$  B, and C correspond to semples of

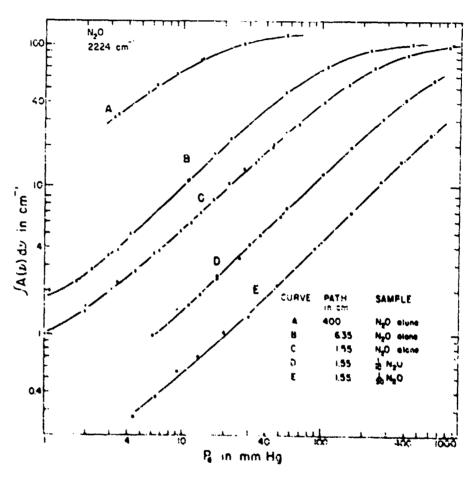


Fig. 15. The total absorption of the 222k cm<sup>-1</sup> F<sub>2</sub>O band versus equivalent pressure for samples whose values of absorber concentration are proportional to the equivalent pressure.

 $H_2O$  alone in the absorption cell, and the absorber concentration y is therefore proportional to  $T_a$ . The ratio of y to  $P_a$  for the three course is different because of the different cell path lengths. Curve D corresponds to samples composed of a mixture of one part  $H_2O$  to size  $H_2O$  to  $H_2O$  to and  $H_2O$  to an  $H_2O$  to  $H_2O$ 

The slopes of curves A, B, and C are seen to decrease, as in the case of the curves of Fig. 12, for large values of  $\mathbf{v}$  and  $\mathbf{P}_a$ . This "saturation", of course, occurs for smaller pressures for the longer paths since the values of  $\mathbf{v}$  are greater; the maximum values of  $\mathbf{v}$  are not sufficiently large to produce saturation in curves D and B.

It is noted that curves B, C, D, and E of Fig. 13 have regions which are approximately linear over a rather wide range of values of equivalent pressure greater than 10 mm Eg. The linear positions of these curves have slopes of approximately 0.91, which is considerably greater than the slope of the linear regions of the curves in Fig. 12. The increase in slope is, of course, due to the fact that w increases with F<sub>0</sub> in the curves of Fig. 13, whereas w is constant for each curve of Fig. 15. The slope of 0.91 indicates that, for the values of w and F<sub>0</sub> covered by the linear portion of these curves,

$$\int A(v)dv + c v^B P_0^{\ D} \qquad (22)$$

where n+n=0.91 and c is a constant. It was shown in the discussion of Fig. 12 that the total absorption is proportional to  $P_0^{0.57}$  over a considerable range of values of y and  $P_0$ . This result indicates that the constant n defined immediately above equals 0.57, and n therefore equals approximately 0.5h. It will be shown in the discussion of Fig. 1h that the total approximately is, in fact, approximately proportional to  $y^{0.5}$  over the same range of values of y and  $P_0$  as are represented by the linear portions of the curves of Fig. 12 and 1).

It is noted that the slopes of curves N. C. D and E of Pig. 13 decrease toward lawer values of P<sub>0</sub>, for P<sub>1</sub> less than approximately 10 mm Eg. This decrease is along toward lower presource cannot be explained on the basis of a band commisting of limes having the lorests manage gives by Eq. 1. The decrease in slope is attributed to Boppler broadening of the lines. Similar devictions from the espected behavior of a band composed of lorests lines were observed by Goody and Vormalita a study of the 116/ end 1200 cm<sup>-2</sup> H<sub>2</sub>U bands and were attributed to

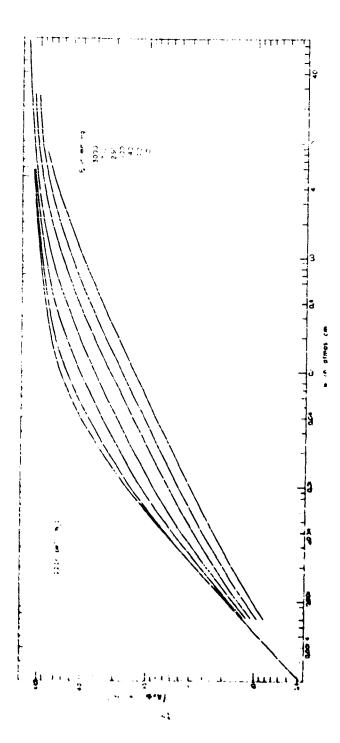


Fig. 14. The total absorption of the 2024 cm<sup>-1</sup> NgO band vermus absorber concentration

Doppler broadening. Benefict<sup>15</sup> et al have who shown that the contribution of Doppler broadening to the absorption or a line or band becomes important at pressures less than approximately 10 mm Hg, causing the absorption to be larger than expected from the Lorentz theory. These results are confirmed in the present study.

There are also indications of a decrease in the slopes of a few of the curves in Fig. 12 at values of  $P_{\rm e}$  less than approximately 10 mm Hg. However, these curves include only a few points at lower pressures, and any explanation of the effects of Doppler broadening based on these curves alone would hardly be conclusive.

In order to illustrate the dependence of total absorption upon absorber concentration w, values were taken from Figs. 12 and 15 and plotted on a logarithmic scale in Fig. 14. Each curve corresponds to a constant value of equivalent pressure indicated in order from top to bottee. From the curves of Fig. 14 it is possible to determine the total absorption of any sample for which the values of absorper rememberation and equivalent pressure fall within the range covered. For samples having values of equivalent pressure which lie between values corresponding to the curves, it is possible to interpolate. The points were emitted in order that readings from the curves night be made more readily. Each curve was down over the range of absorber concentrations covered by the data.

### C. MAID INTERSITY

The lower portion of the uppermost curve in Fig. 12 which corresponds to an equivalent pressure of 1000 mm Hg is seen to be linear with a slope of approximately one, thus indicating a linear dependence on v. The linear dependence on v can be explained on the basis of Eq. (0) for samples such that k(v)v <<1 for all frequencies. Under this condition Eq. (0) can be written as

$$\int h(v)dv + v \int h(v)dv . \qquad (61)$$

where kiv dv is the band intensity.

The condition  $h(v)v\gg 1$  is attained by keeping v very small with  $P_v$  sufficiently large so that h(v) is small, even mean the centers of the spectral lines. The condition of large  $P_v$  and small v are serisfied for the linear region of the top curve in P(v). In. From  $E_v$  (5') the linear dependence on v is obvious, and it is approximate that the value of the quantity f(v) can be determined graphically from the absorption curves of amples represented by the linear portion of the curve.

At the intercept of the uppermost curve of Fig. 14 with the ordinate axis,  $\int A(v)dv = 0.5$ ; cm<sup>-1</sup> and  $v = 2 \times 10^{-4}$  atmost cm. Substitution of these values into Eq. (6') yields a value of

$$\int k(v)dv = 1850 \text{ atmos}^{-1} \text{ cm}^{-2}$$
 (25)

for the band intensity of the 2224 cm -1 NoO band.

The value of  $\int k(v)dv$  given by Eq. (23) can be compared to values of this quantity which here been published by other workers. The results of three other workers are as  $fc^{\dagger}lo^{*}s$ .

ÚDR <b>erve</b> r	k(+)4+	in atmos 1 cm-2
Thorndike and Vellelö	1867	
Calloman, McKean and Thompson17	1617	
Eggers and Crawford18	1650	

In order to find the value of  $\int b(w) dw$  by the method employed in the present study, the approximation that k(v)v << 1 for all v must be valid, a requirement which is satisfied by small values of v and values of  $P_{v}$  sufficiently large so that k(v) is small, even near the centers of the lines. Under these conditions the lines or the band are said to be "completely broadened" and the total openrythm is independent of presence.

Without data which include sufficiently wide ranges of y and  $P_g$  it is difficult to be sure that the conditions of small y and large  $P_g$  are sufficiently well estimated so that the approximation  $k(v)v\ll 1$  is valid. Rome question has arises concerning the results of Thorndita and Wells, which were mentioned above, with regard to this approximation  $k^2$ : Economy. In the present study sufficient data are available which indicate that the approximation is indeed valid.

The left-hand portions of the upper three curves, which correspond to writer of  $P_0 \geq 250$  as Eq. converge. Therefore one conclude that for the value of  $v = 2 \times 10^{-6}$  across on used to calculate  $\int k(v) dv$ , the value of  $P_0 = 1000$  as Eq used was sufficiently large to produce complete broadcateg.

The requirement that y be small is apparently satisfied since the lower portion of the upper curve of Fig. 15 to linear with a slope of very smarly unity. It therefore seems that the accuracy of the value given by Eq. (21) depends more on the accuracy with which the ejectra were measured and the accuracy to which y was determined rather than on the ralidity of the approximation that  $R(y)v\ll 1$ .

The spectra which correspond to the portion of the curve in Fig. 4 from which the reading was taken were recorded with the ordinate scale expanded by a factor of ten. Three spectra were obtained for each sample and the average was used. The short absorption cell (1.55 cm) was used in order to produce small values of absorber concentration without using extremely small partial pressures of N20. In order to improve the accuracy of determining w, the N20 was admitted to the cell in the form of a mixture of N20 and N2. This procedure man it possible to deal with pressures sufficiently large that they could be measured with greater accuracy. Consideration of the total absorption measurements indicates that the value of A (v) dv given by (23) is accurate to \$10\$.

### D. EMPIRICAL EQUATIONS: LIMITATIONS

The curves of Fig. 14 are seen to converge at the lowest and highest values of v, but the five lower curves are nearly linear and parallel over a wide range of v near the centers of the curves. The slope of the linear portions of the curves was found to so approximately 0.53, thus indicating that the total absorption is proportional to  $v^{0.55}$ . The range of values of v and  $P_v$  for which fA(v)dv is proportional to  $v^{0.55}$  is approximately the same as the range for which fA(v)dv was found to be proportional to  $P_v^{0.55}$  in the discussion of Fig. 12. For values of v and  $P_v$  which are represented by the linear portions of the curves of Figs. 12 and 15. these results indicate Unit

$$\int A(v)dv = e^{-V^{0.55}} P_{e}^{0.57} + e^{-[vP_{e}^{0.7}]^{-0.55}}, \qquad (24)$$

where c is a constant. The result in Eq. (24) is seen to be in good agreement with Eq. (22), which is based on Fig. 15. It is recalled that the curves of Fig. 15 are based on data obtained from samples of H<sub>2</sub>O alone and samples of H<sub>2</sub>O and H<sub>2</sub> which were mixed before being introduced to the relimitation from samples formed by adding H<sub>2</sub> to the absorption cell after the H<sub>2</sub>O had been introduced. The good agreement between the results of Fig. 15 and those of Figs. 12 and 15 therefore indicates that very little error in sampling was introduced in the process of adding H<sub>2</sub> to the samples.

From Figs. 12, 1), and 14 is in seen that a very complex function would be required to relate the total absorption to y and  $P_y$  for all values of these parameters. Several experimentary have succeeded in using simple expressions, but these were only valid for limited values of y and  $F_y$ , as in the case of Eq. (24).

In the LBM report, two different types of equations were used. A "weak fit" equation of the form

$$\int A(v)dv = c v^{1/2} P_e^k$$
 (25)

was used for small values of total absorption, and a "strong fit" of the form

$$\int A(v)dv = C + D \log v + K \log P_{e}$$
 (26)

was used for larger values of total abscription above a certain value called a determining value.

Valuer for the constants c, k, C, D, and K as well as the determining value were chosen for each band to give the best fit with the data. Equations (25) and (26) were intended only as empirical equations with the constants having no particular physical or theoretical significance. Since the data presented in the present report cover much wider ranges of both  $\underline{\mathbf{y}}$  and  $P_{\underline{\mathbf{y}}}$  than those covered by HEW, it is of interest to use these new data to determine the regions for which equations of the form given by (25) and (26) are valid. It is of equal interest to demonstrate the short-comings of such empirical equations and to point out the danger involved in uning then to extrapolate to values of  $\underline{\mathbf{y}}$  and  $P_{\underline{\mathbf{c}}}$  not included by the data from which they were derived.

Equation (24) indicates that the total shortion is a function of the product  $vP_0^{-0.7}$ , at least for certain values of v and  $P_0$ . In order to check the usefulness of a function of  $vP_0^{-0.7}$  for all the data obtitued and to illustrate its deficiencies, the curves of Fig. 15 have been drawn. Each curve corresponds to a given value of  $P_0$  as indicated. The points shown in Fig. 15 do not represent actual sumples, but correspond to values taken from the curves of Figs. 12 and 1%. Since each curve corresponds to a given value of equivalent pressure it has the same shape as the corresponding curve in Fig. 14, but the relative positions of the curves are shifted along the abscissa. Some of the curves were and extended over the full range since several curves would be so not included, all the points were included, and indicate shere the curves would occur.

If the total absorption were a function were a function of only the product  $\Psi_{\rm c}^{0.7}$  for all values of w and  $F_{\rm c}$ , the curves of Fig. 15 would all coincide. It is not surprising that the curve corresponding to  $P_{\rm c} \approx 5000$  mm Hg occurs well to the right of the others since the fine offsets of the band is small smoothed out at this high pressure and the effects of increasing the pressure are small as indicated in Fig. 12.

It is interesting to note that the current corresponding to  $P_a \leq 250$  mm Hg occur relatively close together for values of  $vP_a^{0.7}$  greater than approximately 0.5 (steve cm) (mm Rg)<sup>0.7</sup>, a result indicating that over

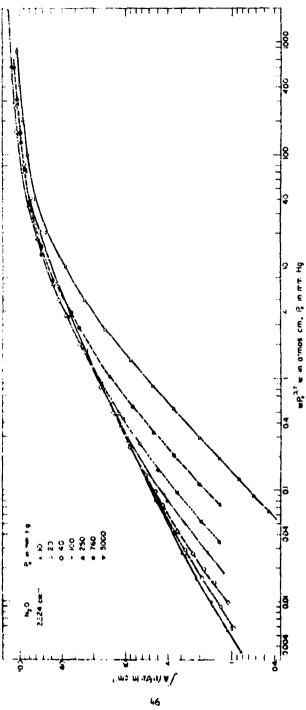


Fig. 15. The total absorption of the 2224  $\rm cm^{-1}~M_2O~tend$  versus ve.  $^{0.7}$  .

this region of values of w and  $P_e$  the total absorption could be approximated by a function of  $wP_e^{-0.7}$ . The curve corresponding to 760 mm Hg also converges with the lower pressure curves for values of  $wP_e^{-0.7}$  greater than approximately j, but this curve falls definitely below the others for smaller values of  $wP_e^{-0.7}$ .

The lower portions of the curves diverge, with the curves for higher pressures and lower absorber concentration falling below the curves for lower pressure and higher values of absorber concentration. This result is expected on the basis of the lower curves of Fig. 12, for which the value of absorber concentration is small and the increase in total absorption with increasing pressure is slight.

The points corresponding to all the values of F<sub>e</sub> except 5000 am Hg occur close together in the rather rist upper portion of the curves of Fig. 15. It should be noted that the closeness of these points does not necessarily indicate that the values of the exponent (0.7) to which P<sub>e</sub> is raised has been determined accurately for the values of w and P<sub>e</sub> represented by these points. Since a change in the exponent would cause only a relative horizontal shifting of the curves of Fig. 15, the value of the exponent could be changed considerably without appreciably altering the ordinates of the rather flat portions of the curves.

The upper portion of the curve corresponding to 10 mm Hg crosses the curves of higher pressure (except 5000 mm Hg) and falls below them. This indicates that for these points the exponent of  $P_0$  should be somewhat greater than 0.7. Thus, one concludes that 0.7 is low for some values of  $\nu$  and  $P_0$  and too large for others, but it is a nominal value which can be used satisfactorily over rather wide ranges without introducing serious error.

From Fig. 15, it is a simple matter to determine graphically some simple empirical functions of the product  $vP_e^{0.7}$ , and to determine the values of  $\underline{v}$  and  $P_e$  for which each function is valid.

The middle portions of these curves which are nearly linear can be represented by the following equation:

for 
$$\int A(v)dv = 19.0 \quad v \left[ P_{o}^{0.7} \right]^{0.55}$$
 and 
$$10 < \int A(v)dv < 15 \text{ cm}^{-1},$$
 (27) 
$$10 < P_{o} < 250 \text{ mm } E_{d} .$$

Equation (27) is the same as  $R_2$ . (2h) except the countant c of (2h) has been evaluated from the current or Fig. 15 and substituted into the latter equation.

For larger values of total absorption the following equation, which is similar to the strong fit (26) of HEM, has been derived.

$$\int A(v)dv = 15 + 40 \log (vP_e^{C.7}),$$
 for 
$$45 < \int A(v)dv < 120 \text{ cm}^{-1},$$
 (28) and 
$$10 < P_e < 760 \text{ am Hg}.$$

Values of total absorption calculated by use of Eqs. (27) and (28) are believed to be accurate to approximately ± 10%.

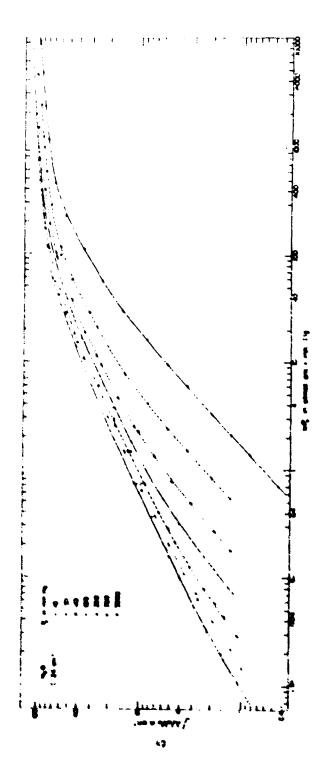
Several investigators have used functions of  $vP_0^{\pm}$  to represent data which included values of y and  $P_0$  covering such smaller ranges than in the present study. Therefore, the limitations of the use of such functions were not illustrated so vividly as in the present study. It should be emphasized that empirical equations involving a function of  $vP_0^{\pm}$  cannot be used for small values of  $v/P_0$  or for high pressures. It is, therefore, suggested that caution should be used when applying the equations listed in the EDM report to samples that are smaller than those from which the equations were derived.

### R. COMPARISON OF MERITAR WITH THE PREDICTIONS HASED ON ELBASSER BAND MODEL

It is recalled from Scutics II that for certain values of v and  $P_{\rm c}$  the strong-1'm approximation of the Eleaseer band model is valid and the total observation is a function of  $w_{\rm p}^{\rm c}$ . Superpr. it has been shown in the previous section that a function of  $v_{\rm p}^{\rm c}$ . If it the 2224 cm<sup>-1</sup>  $E_{\rm p}^{\rm c}$  data over rather wide ranges of values of the parameters. It is therefore of interest to show the relationship between the total absorption of the 2724 cm<sup>-1</sup>  $E_{\rm p}^{\rm c}$  band and the product  $v_{\rm p}^{\rm c}$ .

In order to show this relationship graphically, the logarithm of the total absorption was plotted in Fig. 16 against the logarithm of wP<sub>0</sub> for the different pressures indicated. The points shown in Fig. 16 do not correspond to samples actually measured, but to values taken from the curves of Figs. 12 and 15. By this providers it was possible to obtain sets of points which correspond to a given value of P<sub>0</sub> as was done for Figs. 15 and 15. Since each curve corresponds to a given value of equivalent pressure, it has the same shape as the corresponding curves in Figs. 15 and 15, but the relative positions of the turves are phifted along the abscisse.

The different curves corresponding to values of  $P_0 \leq 0.02$  as  $M_0$  occur very close together in the region corresponding to  $M_0$  greater



Pig. 15. The total absorption of the 222%  $m^{-1}$  H2C than regress of  $\epsilon$ 

than approximately 300 atmos cm x cm Hg. This result indicates that it is possible to relate the total absorption to the variable  $vP_{\mu}$  for the values of w and  $P_{c}$  which are involved. However, this result is accessively inconclusive, where it was shown in Fig. 19 that the total absorption of samples involving the same values of w and  $P_{c}$  could also be related to the variable  $vP_{c}^{(0)}/l$ . In the discussion of Fig. 15 it was pointed out that for these samples the absorption is nearly complete, and that the total absorption of the band increases only slightly with increasing w or  $P_{c}$ . The total absorption can therefore be satisfactorily related to a function of  $vP_{c}^{(0)}$ , and the value of the exponent a can be varied considerably.

The curves are separated considerably for values of  $\mathbf{v}$  and  $\mathbf{P_c}$  for which the absorption is still increasing with either  $\mathbf{v}$  or  $\mathbf{P_c}$ . This result, of course, indicates that the total absorption cannot be said. factorily related to the single variable  $\mathbf{vF_c}$ . However, it is of interest to look in particular at the regions of one curves of Fig. 16 for which the strong-line approximation would be valid if the absorption band being studied could in fact be represented by an Eleasser band.

The regions of the curves which are to be investigated can be determined from curves and equations given by Plass<sup>9</sup> for the region of validity of the strong-line approximation for an Elsanger band. Before the present data, which are given in terms of the parameters  $\mathbf{v}$  and  $\mathbf{P}_{c}$ , can be compared to the theoretical predictions, it is necessary to relate  $\mathbf{v}$  and  $\mathbf{P}_{c}$  to the dimensionless quantities  $\mathbf{x}$  and  $\boldsymbol{\beta}$  which serve as the variables in the theoretical equations.

As indicated in Section II, the quantity x is given by

$$x = \frac{8v}{2\pi i}; (29)$$

the parameter \$ is given by

$$b = \frac{2\pi z}{d} \tag{30}$$

where d is the frequency separation of adjacent absorption lines. For a given band,  $\beta$  is proportional to  $\alpha$ , which in turn is proportional to  $P_{e}$  (Eq. 14); therefore it is necessary to find the constant relating the two quantities  $\beta$  and  $P_{e}$ . The value of the proportionality constant can be determined by using values of  $\alpha$  and d for an NgO band, taken from an article by Goody and Wormelli's mentioned previously. The value used for d was 0.04 cm<sup>-1</sup>; and  $\alpha$  was negligible to be proportional to  $P_{e}$  and equal to 0.15 cm<sup>-1</sup> at a pressure of one atmosphere. Therefore it can be shown that  $\beta=1$  corresponds to  $P_{e}=1$ 

$$\beta = 2e/600 \tag{31}$$

where Pe is expressed in ma Ra.

by uning Eq. (fr; and applying the conditions given by these or the validity of the strong-line approximation of an Eleaser band, it is possible to determine regions of the curves of Fig. 16 over which one would expect the total absorption to be a function of the variable vP<sub>n</sub>, accurate to 25%.

For example, for  $P_{e} \leq 250$  mm Hg, the total absorption to a function of  $vP_{e}$  accurate to 15% for values of  $\int A(v)dv$  greater than approximately 60 cm<sup>-1</sup>, and the relative error should decrease as the total absorption increases. At the value of  $vP_{e}$  for which the curve of Fig. 16 corresponding to 250 mm Hg gives a value of  $\int A(v)dv$  equal to 60 cm<sup>-1</sup>, the curve corresponding to 10 mm Hg reads approximately  $VP_{e}$  cm<sup>-1</sup>. Thus it is seen that the total absorption cannot be represented, with an accuracy of 5%, as a function of the single variable  $vP_{e}$  for the values of  $VP_{e}$  and  $VP_{e}$  predicted by the Elsasser band model. The error is seen to be more nearly 20% than 5%. Similar deviations from the predicted strong-line dependence were found for other values of  $P_{e}$ .

As just indicated, the dependence of the total absorption of this band on  $\underline{v}$  and  $P_{e}$  cannot be described by a simple Elsasser band, and can be described by a nample empirical equation only over limited ranges of  $P_{e}$  and  $\underline{v}$ . However, the total absorption could possibly be described in terms of a "random Elsasser model" made up of two or more Elsasser bands of different intensities. Such an analysis of the 9.6 $\mu$  except band has been made by Place9.

From the previous discussion it can be concluded that the total absorption of the 222½ cm $^{-1}$  N<sub>2</sub>O bend is less dependent upon  $F_{\rm c}$  than upon  $\Psi$  for essentially all values of  $\Psi$  and  $P_{\rm c}$ . This is in contrast to an Elsasser band for which the strong-line approximation is valid over a considerable range of values of  $\Psi$  and  $P_{\rm c}$ , and the total absorption has the same dependence on both  $\Psi$  and  $P_{\rm c}$ .

The low dependence upon pressure as compared with an Elsasser band can be explained, quelitatively at least, from the fact that the spectral limit are not all of equal intensity 2 as assumed for an Elsasser band. Much of the band where the lines are weak, for which (8) is approximately valid, and where the effect of increasing pressure is small while the effect of increasing absorber concentration is large. Unless all the lines of the band are strong, the effect of increasing absorber concentration, which is the result observed in the present study.

### F. FURTHER TESTS OF ELSASSER MODEL: ABSORPTION AT R-MAX. AND P-MAX.

Since it is probable that the absorption by weak lines in the wirgs of the band contributes to the deviation of the total absorption form a strong-line dependence, it is of interest to investigate the absorption  $A(\nu)$  at the frequencies of maximum absorption in the R- and F-branches, where the strong-line approximation might be expected to apply. These two frequencies will hereafter be referred to as R-max, and P-max, respectively.

The major portion of the absorption observed at these frequencies arises from the spectral lines whose centers occur within the spectral interval passed by the spectrometer. If one could assume that all the lines in this interval were evenly spaced and had the same half-width and line strength, one would expect that for a certain range of values or w and P<sub>0</sub>, the absorption would show the strong-line dependence producted by an Elsasser model. However, it was mentioned above that the spectral region under investigation contains weak bands as well as the strong fundamental band. Therefore a portion of the absorption observed at frequencies of peak absorption is due to weak lines whose centers occur within the spectral interval of the spectrometer. The effect of these weak lines would be to cause the observed absorption to deviate somewhat from the titoug-line dependence.

There is also a contribution to the absorption by the wings of spectral lines whose centers occur outside the spectral interval passed by the spectrometer. However, as was discussed in Section II in connection with the streng-line approximation, the absorption by the wings of a line has a strong-line ispendence. Therefore, the absorption by lines whose centers are outside the spectral interval passed by the spectrometer would not be expected to cause the observed absorption to deviate from the strong-line dependence.

The values of the absorption A(2240 cm<sup>-1</sup>) and A(2215 cm<sup>-1</sup>), the frequencies corresponding to R-max, and N-max, respectively, have been tabulated for each spectrum in Table 2. The frequency of the regions of maximum absorption shifted very slightly as the shape of the band changed for different values of y and P<sub>p</sub>, but the difference between the maximum absorption and that at the two frequencies listed is very slight, since the slope of the absorption curves is small near these frequencies. Some of the absorption curves, such as those in Fig. 5, do not have two distinct absorption maxima, but for such spectra the value of percent absorption was measured at the two frequencies designated.

The values of absorption  $A(2240 \text{ cm}^{-1})$  of several of the samples are plotted against the equivalent pressure in Fig. 17. Each curve corresponds to a given value of y as indicated, with "x's" corresponding to camples of N<sub>2</sub>O alone and the dots corresponding to anyles having N<sub>2</sub> satisfy to the  $R_2$ O. The curves have the sum general features as those in Fig. 12,

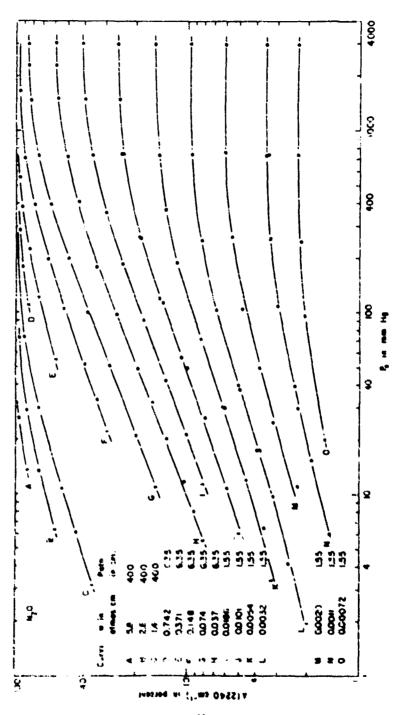


Fig. 17. The percent absorption at 22.0 on? 'vertus equi w'est presture for (taples maxing ton-stat. "Thestration.

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In which the total absorption was plotted against  $F_e$ . The slopes of the curves decrease at high veloce of equivalent pressure and the curves converge at 100% ebsorption for large values of  $\underline{v}$  and  $P_e$ . Several of the curves have portions for 10 km Hg <  $P_e$  < 200 km Hg which are nearly linear and approximately parallel. The slope of these portions of the curves is approximately 0.41, i licating that the absorption is propertional to  $P_e^{-0.41}$ . It is noted that the  $P_e^{-0.41}$  dependence is significantly less than the  $P_e^{-0.5}$  dependence which is associated with the well-known square-root approximation.

There seems to be a slight doctense in the slope of curves K and L at equivalent pressures less than six or eight as Hg; this decrease in slope probably corresponds to additional absorption resulting from Duppler broadening.

The absorption  $A(2240~cm^{-1})$  at H-branch maximum was plotted in Fig. 18 against  $P_{\rm e}$  for the samples for which total absorption is plotted in Fig. 15. These curve differ from those in Fig. 17 in that each curve in Fig. 15 does not correspond to a given value of w, but w increases in proportion to  $P_{\rm e}$ . In each sample w is proportionat to  $P_{\rm e}$ , but the ratio  $w/P_{\rm e}$  is smaller for the sample represented by each successive curve. Curves  $B_{\rm e}$ ,  $D_{\rm e}$ , and  $E_{\rm e}$  are seen to have marrly linear portions of slopes approximately 0.92, with decreasing slope for larger values of  $P_{\rm e}$  for which overlapping of the lines become significant. The linear portions of the curves suggest that the absorption is proportional to w  $P_{\rm e}^{\, B}$  where the run of the exponents m and m is approximately 0.92. This result is to be compared with the predicted square-root region over which the absorption is proportional to  $w^{0.5}P_{\rm e}^{\, 0.5}$ . The decrease is slope of curves  $B_{\rm e}$   $C_{\rm e}$   $D_{\rm e}$  and  $E_{\rm e}$   $D_{\rm e}$  and  $E_{\rm e}$   $D_{\rm e$ 

As stated earlier, the absorption can is represented by the single variable way for values of w and P for which the strong-line approximation is valid. In order to show the dependence of the observed absorption on the variable way, values of absorption A(2240 cm<sup>-1</sup>) were taken from the curves of Figs. 17 and 18 and plotted as a function of way in Fig. 19, where each curve corresponds to a given value of equivalent reserve. The features of these curves are similar to times of Fig. 26 in which the ordinate represents total absorption rether than percent absorption at a fixed frequency. Validity of the strong-line approximation would, of course, be represented in Fig. 19 by a coincidence of the curves corresponding to the different pressures. As in the case of Fig. 25, which is smiller except that total absorption is plotted as the ordinate, must of the curves corresponding to higher pressures occur to the right of and below the curves corresponding to lower pressures tend to converge for large values of w, a result indicating that for these values of w and P, the total absorption child be expressed by a attrong-line approximation.

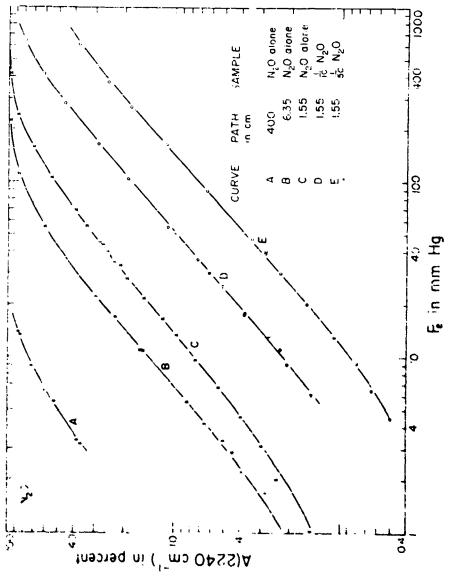


Fig. 18. The percent absorption of 22t0 cm² versus contralent pressure for samples where values of absorber concentration are proportional to the equivalent pressure.

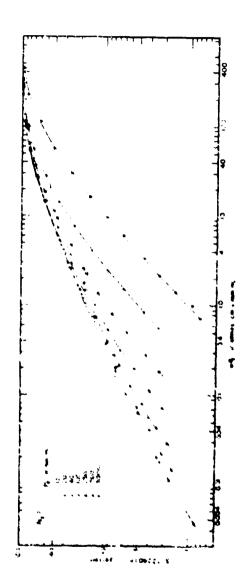


Fig. 19. The percent sixospison at 22-C mm2 versus are

The three curves corresponding to values of  $P_e$  equal to 10, 20, and 40 we Hg contain portions which the nearly straight with a slope whose value is approximately 0.50. This result corresponds to the well-known square-root dependence on  $\mathbf{v}$ ; however, it is apparent from these curves that the dependence on  $P_e$  is less than square-root since the curves corresponding to higher pressures occur to the right of those corresponding to lower pressures. It is also recalled from the discussion of Fig. 17 that the maximum dependence on  $P_e$  is only approximately  $P_e^{-0.61}$ .

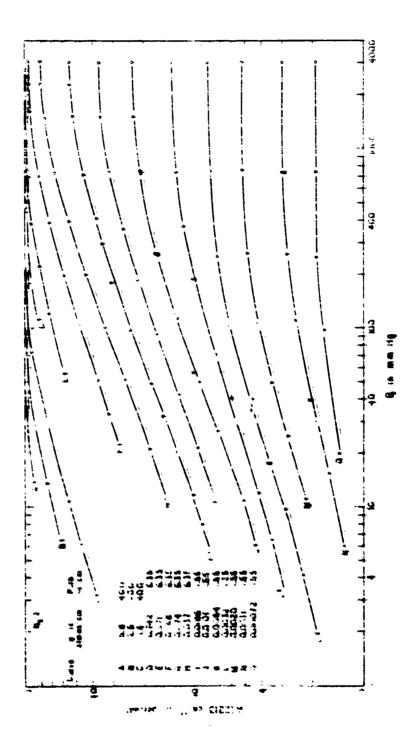
It is noted that the  $v^{0.50}P_n^{0.51}$  dependence which is found from Pigs. 17 and 19 is consistent with the result of Pig. 18, which indicated that the absorption was proportional to  $v^m$   $P_n^{-n}$  where n+n=0.92.

Curves statlar to those in Figs. 17, 18, and 19 were drawn for the absorption A(2217 cm<sup>-1</sup>), the frequency corresponding to maximum absorption in the F-branch. The features of the curves of Fig. 20 are similar to the corresponding curves of Fig. 17; the slope of the straight protions of the curves of Fig. 20 is approximately 0.36, which is slightly less than the value 0.41 created in Fig. 17. This result indicates that the maximum dependence on prescure is slightly less for the peak absorption in the F-branch than in the r-branch.

The curves of rig. 22 are also similar to the curves of the corresponding Fig. 15 which represents the A-tranch absorption. The slopes of the straight portions of the curves in Figs. 15 and 21 are essentially the same, 0.92. Similarly, the features of the curves of Fig. 22 are like those of Fig. 10; however, the slope of the ctraight portions of the curves of Fig. 22 seems to be slightly greater (0.5c instead of 0.50), although this small difference is hardly significant.

Figures 19 and 22 indicate that the absorption can be represented by the strong-line approximation only for numit values of  $P_0$  and very large values of v. For smaller values of v such that the absorption is proportional to  $\overline{v}^{0.5}$  the dependence on  $P_0$  is significantly less, and the strong-line approximation is v- longer valid. It is of interest to convert the curves of Figs. 19 and 22 with theoretical results based on an ileaseer band in the same manner as was done with the curves of Fig. 16. It was possible to make such a comparison by assuming the relationship between 0 and  $P_0$  that is given by Eq. (31) and applying the conditions given by Place? for the validity of the strong-line approximation.

For example, if  $\Gamma_o \leq 150$  cm Eq. the absorption of an Essesser band can be expressed as a function of  $vP_o$  within 95% for values of absorption greater than approximately  $v(\phi)$ . In comparing this with the curves of Fig. 19, it is found that the absorption  $h(2200 \text{ m}^{-1})$  equals approximately 95% at  $P_o = 250$  cm Eq for the same value of  $vP_o$  which gives a value of 50% absorption on the 10 cm Eq curve. Thus there is an observed relative uniformly  $v_o = 1500$  cm equals on the 10 cm Eq curve.



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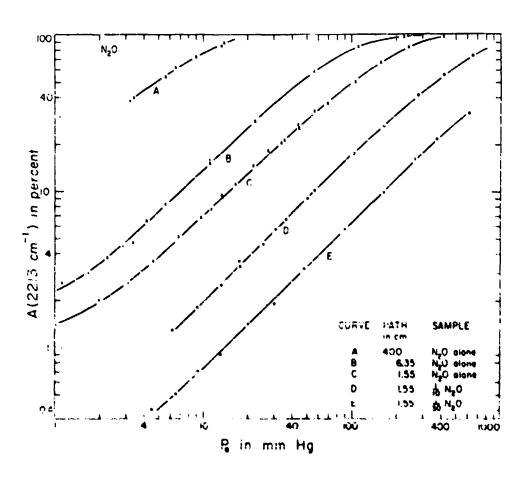
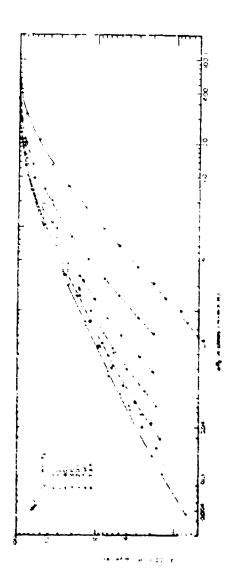


Fig. 21. The percent absorption at wait can't versus equivalent pressure for samples whose values of shauther concentration are projectional to the equivalent pressure



Mg. 22. The percent alsorption at 2213 cmr versus afe

the Elsasser model. Similar readings from the curves of Fig. 22 which represent  $A(2213~{\rm cm}^{-1})$  indicate an even greater deviation from the predicted values.

If  $P_e$  is restricted to lower values, such as 100 mm Hg or less, Plass has shown that theory based on the Eleasser model predicts that the strong-line approximation is accurate to  $\pm 5\%$  for absorption greater than approximately 20%. Inspection of Fig. 19 shows that there is approximately 20% relative difference between the leadings of the two curves corresponding to 10 mm Hg and 100 mm Hg in the region where  $A(22\text{ho cm}^{-1})$  equals approximately 20%.

On the basis of these comparisons it can be concluded that the absorption at frequencies of maximum absorption in the P- and R-branches cannot be accurately represented by a strong-line approximation. It seems improbable that an error in the value of the constant relating  $P_{\rm g}$  to  $\beta$  in (51) could account for the deviation of the observed results from the predicted behavior. It can be shown that if the constant in Eq. (51) were in error by as much as a factor of two, the data would still show deviation from the strong-line approximation for smaller values of  $P_{\rm g}$  than those predicted.

It is of interest to compare the results of Figs. 19 and 22, which deal with absorption A(v) at certain frequencies, with those of Fig. 16 which deals with total absorption A(v)dv; the curves representing absorption approach a very definite upper limit of 1005. But the curves representing total absorption tend to indicate a small increase with increase with a continuing growth in the wings of the band. The curves which correspond to the lower values of present in all three of these figures all contain portions which are nearly straight with slopes of approximately 0.5. Thus, the total absorption as well as percent of absorption increases approximately as the square root of absorber concentration for vertain ranges of values of v and F<sub>a</sub>. Untilar comparison of Figs. 12, 13, and 20 indicates that the maximum dependence of a not greatly different for total absorption on F<sub>a</sub> was found to be approximately P<sub>a</sub> 0.37, while a dependence of P<sub>a</sub> 0.31 and P<sub>a</sub> was conserved for the absorption at R-max. Twe pectively.

It was suggested earlier that the absorption by weak lines in the vings of the bead might be one of the factors accounting for the fact that the total absorption and not show a square-root dependence on P<sub>a</sub>. However, it is seen that the absorption at R-max, and P-max, which is unaffected by absorption in the wings of the band, showed only a slightly greater dependence on P<sub>a</sub> than did the total absorption. It is nessible, therefore, that the failure of either of these quantities, paramate atsorption or total absorption to exhibit strong-line properties is included hands which are superimposed in the strong fundamental hand.

By considering the difference between the growth of weak lines and the growth of strong lines with increasing pressure, it is probable to give a qualitative explanation of the greatly different shapes of the absorption curves shown in the right-hand portion of Fig. 4.

It has both shown that the absorption at R-max, and P-max, has a rather large dependence on P<sub>0</sub> over fairly wide ranges of values of y and P<sub>0</sub>. It is therefore possible for a sample of low absorber concentration and high pressure to produce greater absorption than a sample of much greater absorber concentration but low pressure. This is the case for the spectra in Fig. 5; the sample having the smaller value of absorber concentration and higher pressure shows greater absorption at R-max, and P-max, than does the sample of larger absorber concentration and lower pressure. However, the lines in the wings of the band are much weaker and grow less rapidly with increasing pressure. The absorption in the wings of the band is therefore less for the sample having the smaller value of absorber concentration, in space of its higher pressure.

## VI. THE FURNISHMAL AND FIRST OVERTORS. BARRE OF CARRON MONORIDE

## A. THE 2145 cm-1 CO MARD

Rigeotte<sup>ly</sup> identified CO as a permanent atmospheric constituent from solar spectra obtained at Columbus, Chio; and several investigators, including Manusol, have since made studies of the CO communt in the atmosphere and found a concentration of approximately 0.1 atmosphere con par air mass. One air mass is an atmospheric path containing the same amount of air as a vertical path through the entire atmosphere.

The fundamental band of CO cours near the v<sub>2</sub> B<sub>2</sub>O long in the atmospheric window which lies between the strong \$550 cm<sup>-1</sup> CO, hand and the 1995 cm<sup>-1</sup> B<sub>2</sub>O band. The major portion of the absorption by CO in this spectral region is due to Colombia vibration-rotation band of the cumum tutiops Charle. Other isotopic bands occur in this spectral region, as do some difference made which are very weak for samples at room temperature. In the present report all tasses absorption bands together are referred to as the 21½ cm<sup>-1</sup> CO band.

Approximately 180 spectra of the CC hand were obtained in the present study by using the Arch's-beam spectrameter along with adsorption colle of lengths 4.35, 6.35, and 400 cm. Different combinations of CO and My were used to provide values of absorber concentration from 9.6 x 10<sup>-6</sup> to \$5.6 whose cm and total pressures from approximately 3 to 3000 cm Mg.

Tractage of a large number of absorption curres are shown in Figs. 23-29, where they appear in groups. The values of absorber concentration y and equivalent pressure  $P_q$  of the sample curresponding to each

**17.** 

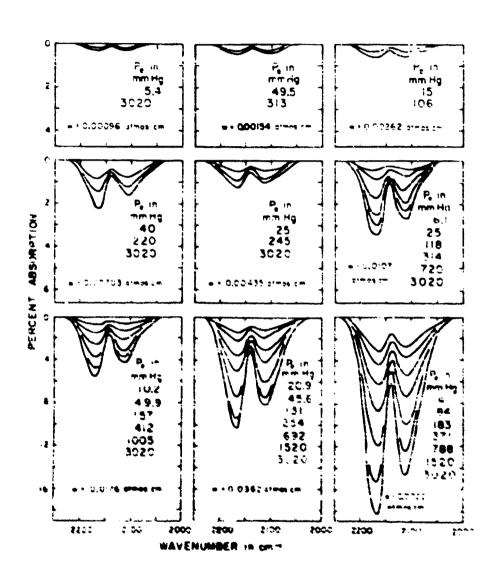
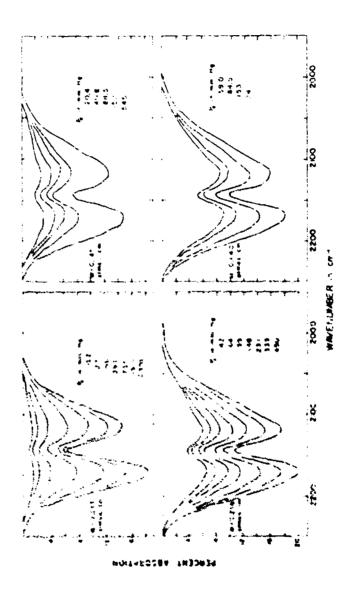


Fig. 75. Specime of the 2143 cm<sup>-1</sup> CO bend

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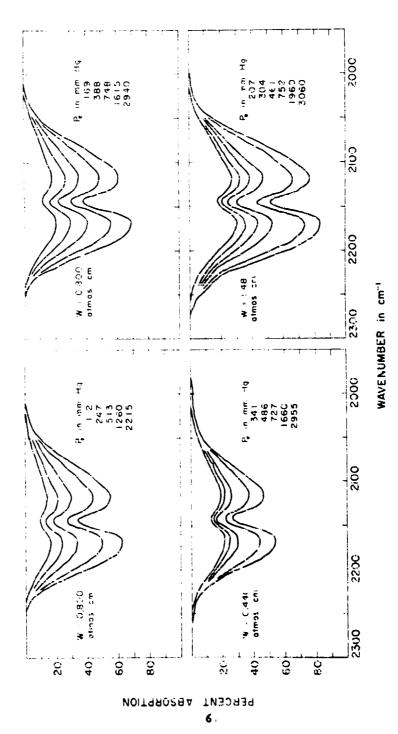
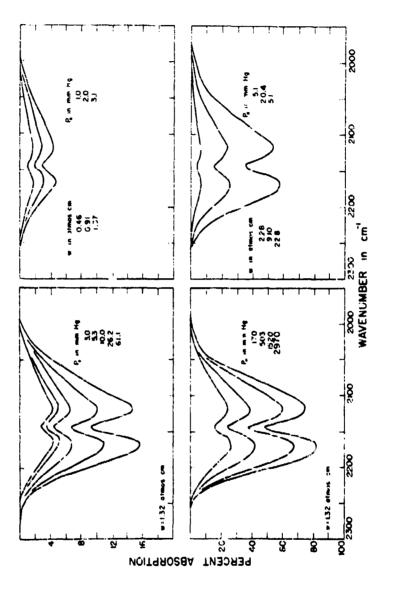


Fig. 25. Speates of the 2143 can. 30 band



Pig. 27. Spectre of the 2145 cm<sup>-1</sup> to band

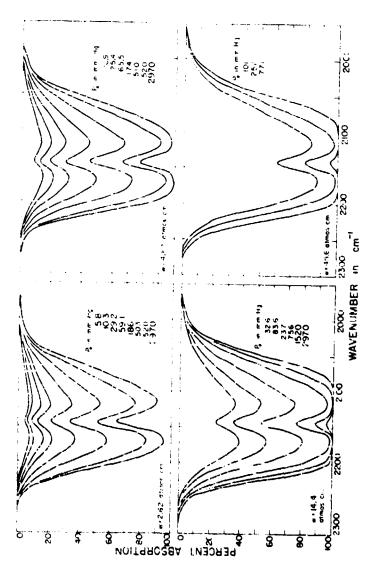


Fig. 28. Spectra of the 2143 ontl CO bana

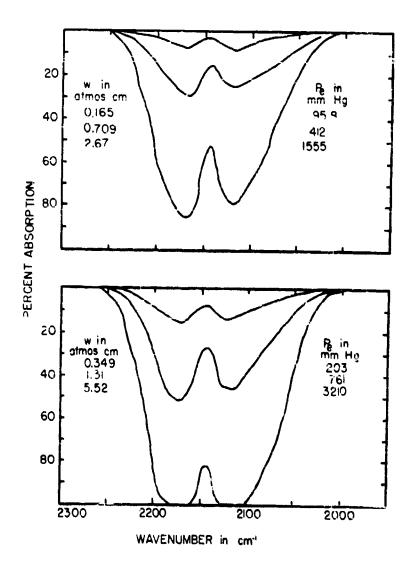


Fig. 29. Spectra of the 2145 cm-1 CO band.

spectrum appear with each group. All the spectra were obtained with the same spectrometer slit width, which corresponded to approximately 25 cm<sup>-1</sup>.

All of the data for the 2145 cm<sup>-1</sup> CO bond included in the present report are collected in Table 4. The absorber concentration y and equivalent pressure  $P_{e}$  for each sample are shown. The value of the equivalent pressure  $P_{e}$  was calculated from the total pressure P and the CO partial pressure P by the equation

$$P_{\rm e} = P + 0.02p,$$
 (32)

in accordnance with Eq. (14) and Table 1.

Included in Table 4 are the values of total band absorption  $\Lambda(v)$ dv along with the observed percent of absorptions  $\Lambda(2116~{\rm cm}^{-1})$  and  $\Lambda(2168~{\rm cm}^{-1})$  corresponding in good approximation to the absorption maxima in the P- and R-branches of the band, respectively.

Many of the total absorption data listed in Table 4 are plotted in Fig. 30 with the total absorption presented as a function of the equivalent pressure. Each curve corresponds to a given value of absorber concentration and shows the effect of increasing the total pressure by adding M2. Curves F, G, H, and I correspond to values of absorber concentration which occur in the same range as the values corresponding to curves B, C, D, and E. These two sets of curves represent samples which were obtained by using two absorption cells of quite different length and with quite different values of CO martial pressure. Since these two sets of curves represent approximately the same range of values of absorber concentration; it is possible to compare the total absorption of a sample in the 6.35-on cell with that of a sample in the 400-cm cell which has the same, or nearly the same, values of w and P2. Such comparisons can be seen to indicate that the results are consistent to suproximately 136.

The general features of the curver of Fig. 30 are the same as those of the corresponding curves for the 222½ cm<sup>-1</sup>  $B_2O$  band which were shown in Fig. 12. The variation of the total absorption with changes in w and  $P_g$  becomes very slight for large values of w and  $P_g$ . This result is illustrated by the crowding of the curves in the unper right-hand corner, and is due to the fact that  $k(v)v \gg 1$  for most frequencies within the band, and the total absorption changes slwly.

There is a general decrease of the slope of the curves towards higher pressures as a result of the change in the shape of the spectral lines which was described in connection with Fig. 12. Rowever, the decrease in slope is less pronounced for the CO curves in Fig. 30 than for the  $R_{2}$ O curves in Fig. 12; this result can be explained by the fact

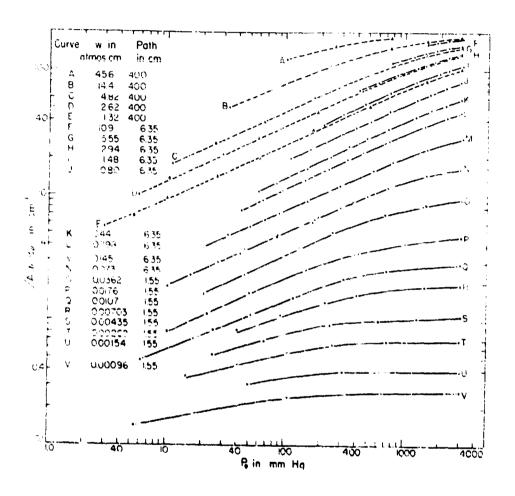


Fig. 50. The total absorption of the Plus own CO band versus equivalent pressure for samples having constant values of absorber concentration

that the quantity  $p=0m^2/d$  is less for 00 than for N<sub>2</sub>0 at the same pressure. By using an average value of  $d=5.8~\rm cm^{-1}$  and  $\alpha=0.08~\rm cm^{-1}$  reported by Shaw<sup>1</sup>9 at a P<sub>e</sub> of 760 mm Hg, it is easily shown that  $\alpha=1$  corresponds to N<sub>2</sub>0 equal to approximately 7750 mm Hg. It is recalled that for N<sub>2</sub>0,  $\beta=1$  for P<sub>e</sub> equal to approximately 680 mm Hg. The dependence of total absorption on P<sub>e</sub> therefore starts decreasing at a smaller value of P<sub>e</sub> for the N<sub>2</sub>0 band than for the C0 band. Therefore the structure of the C0 band would be less "smeare" than that of the N<sub>2</sub>0 band at the maximum values of P<sub>e</sub> used in the present study.

Several of the curves of Fig. 30 contain portions which are nearly straight and parallel to each other. The slope of the straight portions of these curves is approximately 0.44. This result indicates that the total absorption is proportional to  $\Gamma_0^{-0.44}$  for samples whose values of  $\Gamma_0$  and  $\Psi$  are within the range represented by the straight portions of the curves.

In Fig. 31 are three curves relating the total absorption to the equivalent pressure. The curves of Fig. 31 are different from those of Fig. 30 in that those in Fig. 31 do not correspond to constant values of absorber concentration. The curves of Fig. 31 correspond to samples of CO alone in the absorption cell, and the absorber concentration  $\underline{\mathbf{v}}$  is therefore proportional to  $\mathbf{P}_{\mathbf{n}}$ .

The features of Fig. 31 are similar to the corresponding Fig. 13 for the 2224 cm<sup>-1</sup> B<sub>2</sub>0 band. There is a decrease in the slope of curve A at low pressures as a result of Doppler broadening, and at large values of y and  $P_0$  as absorption because nearly complete. The curves of Fig. 31 have straight portions with a slope of approximately 0.98. For values of y and  $P_0$  represented by the straight portions of the curves, the total absorption is therefore proportional to  $y^{m_0}$  where  $m_0 + n = 0.98$ . It was shown previously in the discussion of Fig. 30 that the total absorption was proportional to  $P_0^{0.58}$ . Therefore n = 0.58, and  $m_0 = 0.98 = 0.98 = 0.98$ . It is accordingly to be expected that the total absorption is proportional to  $y^{0.58}$ . This result is confirmed below in the discussion of Fig. 35, where the ranges of values of y and  $P_0$  for which this simple relationship is valid will be determined.

In order to determine the dependence of total absorption upon absorber concentration w, values were taken from Fig. 50 and 31 and plotted in Fig. 52, in which each curve corresponds to a constant value of equivalent pressure as indicated in order from top to bottom.

The lower portion of the uppermost curve, which corresponds to an equivalent pressure of 3000 mm Eg, is linear with a slope of approximately 1. From this portion of the curve it is possible to determine directly the value of the band intensity  $\int k(\nu) d\nu$  in the case manner as was done for the 2224 cm<sup>-1</sup> EgO band.

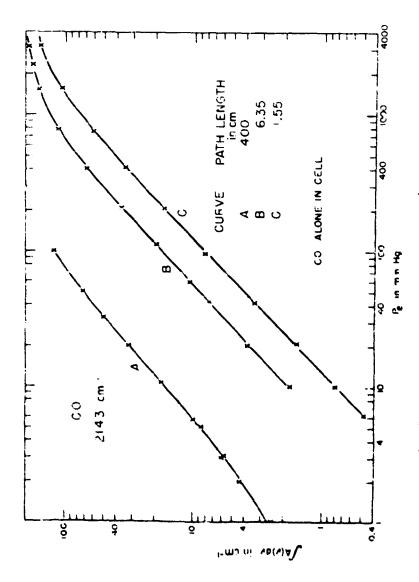
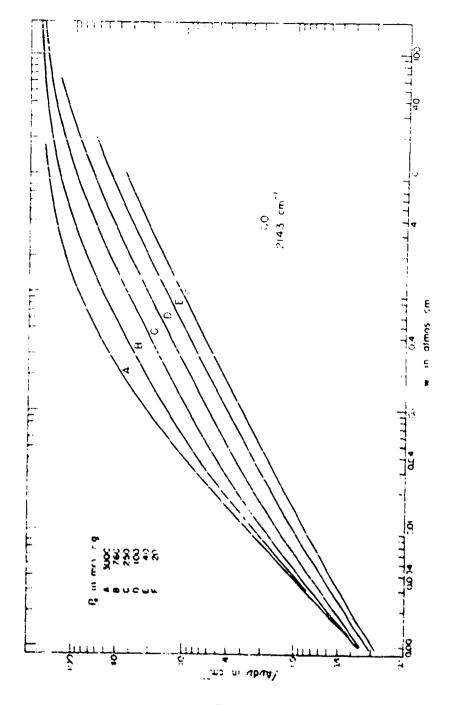


Fig. 31. The total absorption of the 2145 cm<sup>-1</sup> fo band were as equivalent prisure for samples whose we use of absorper circentration are proportional to the hydra's of pressure



Pig. Sci. The ties absorption of the 2145 meril 20 panal versus absorber concentration.

The value determined for the 2143 cm-1 00 band is

$$\int k(\nu) d\nu = 260 \text{ atmos}^{-1} \text{ cm}^{-2} , \qquad (55)$$

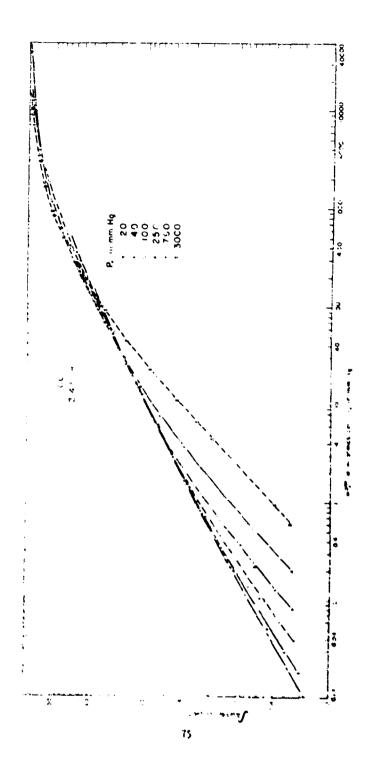
This value, which is believed to be accurate to within \$1.5, can be compared to the value of 237 atmos<sup>-1</sup> cm<sup>-2</sup> reported by Penner and Weber<sup>-1</sup>, and apparently not corrected to STP. If the values of v in Fig. 32 had not been corrected to standard temperature, a value of  $\int f(v) dv = 235$  atmos<sup>-1</sup> cm<sup>-2</sup> would have been obtained. The difference between this value and that of Penner and Weber is within the limits of experimental error.

From the curve of Fig. 32 it is possible to determine the total absorption of any sample whose values of apsorper concentrative and equivalent pressure full within the range covered. Economy, it is possible to derive empirical equations which relate total absorption to the parameters y and P<sub>2</sub> and make it possible to estimate total absorption without the use of the curves of Fig. 32.

The curves of Fig. 32 converge at the lowest and highest values of  $\psi$ , but the five lower curves are meanly linear and parallel over a wide range of  $\psi$  near the centers of the curves. The slope of the linear portions of the curves is approximately 0.55, thus indicating that the total absorption is proportional to  $\psi^{0.55}$ . The range of values of  $\psi$  and  $P_0$  for which  $fA(\psi)d\psi$  is proportional to  $\psi^{0.55}$  is approximately the same as the range for which  $fA(\psi)d\psi$  was found to be proportional to  $P_0^{0.55}$ . In the discussion of Fig. 31 it was suggested that the total absorption was proportional to  $\psi^{0.56}$   $P_0^{0.55}$ . The  $\psi^{0.56}$  is approximately the absorption was proportional to  $\psi^{0.56}$   $P_0^{0.55}$ . The  $\psi^{0.56}$  is a consistent of  $\psi$  which were derived from two different sets of curves or the exponent only to samples of CO alone, while the curves of Fig. 32 are based on samples of CO alone as well as easiles of CO atoms with  $H_0$ .

Since, for certain values of y and  $P_0$ , the total absorption is proportional to  $v^{0.55}$   $P_0^{0.54}$   $= (vP_0^{0.5})^{0.55}$ , it is of interest to investigate the range of values of y and  $P_0$  for which the total absorption might be expressed as a function of  $vP_0^{0.5}$ . In order to check the possibility of using such a function and to illustrate the deviations of the data from it, the curves of Fig. 33 were plotted with the total absorption upding the variable  $vP_0^{0.5}$ . Each curve corresponds to a given value of  $P_0$  as indicated. The points such do not represent actual samples but correspond to values taken from the curves of Figs. 39 and 31.

If the total absorption were a function of the product  $\Psi P_0^{-0.8}$  for all values of  $\Psi$  and  $P_0$ , the curves of Fig. 33 would all coincide. It is seen that there is a inther wide region over which the curves corresponding to  $P_0 \leq 260$  mm Rg are nearly coincident. From the curves it is easy to show that



71g. N5. The total absorption of the 7143  $\rm cm^{-1}$  10 band wreass  $v_{\rm e} \phi_{\rm e} \theta_{\rm e}$ 

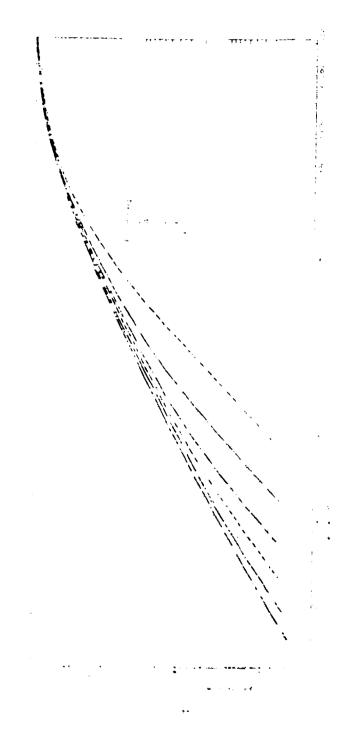
$$\int A(\nu) d\nu = 2.75 \; (\nu P_e^{0.8})^{0.55}, \qquad (3b)$$
 for 
$$4 < \int A(\nu) d\nu < 10 \; \text{cm}^{-1} \; \text{if } 20 < P_e < 250 \; \text{ma} \; \text{Hg}$$
 and for 
$$10 < \int A(\nu) d\nu < 10 \; \text{cm}^{-1} \; \text{if } 20 < P_e < 760 \; \text{m Hg} \; .$$

In the lower left-hand portion of Fig. 35 the curves corresponding to high pressures fall to the right of the curves corresponding to lower pressures. This result is consistent with the results of the corresponding Fig. 15 for the H<sub>2</sub>O band and indicates that a function of vP<sub>0</sub> gives too much weight to the equivalent pressure. The data corresponding to the curves in the lower left portion of the figure could therefore he more accurately represented by a function of vP<sub>0</sub> where a is less than 0.8. In the upper right portion of Fig. 35 the curves corresponding to high pressures fall to the left of, or above, the curves corresponding to low pressures. Some of the curves corresponding to lower pressures do not extend to the right-hand portion of the figure; however, it is apparent that if they were extrapolated they would fall below the curves corresponding to higher pressures. The data represented by the right-hand portion of the figure could thus be more accurately represented by a function of vP<sub>0</sub>, where a is greater than 0.5.

Thus, the total absorption can be represented by a function of wP<sub>0</sub>. Thus, the total absorption can be represented by a function of wP<sub>0</sub>. The certain values of w and P<sub>0</sub>, but sum a function gives too named to a function of w. For larger values of w such a function gives too little weight to P<sub>0</sub>. This latter result would be expected on the basis of the requirement: for the strong-line approximation, which gives the total absorption as a function of wP<sub>0</sub> and is valid for large w and small P<sub>0</sub>.

In order to show the relationship between the total energities and the wariable w2, these parameters were platted in Fig. 34 for the different values of the pressures indicated. Over the major portion of Fig. 34 the curves corresponding to high pressures lie to the right of the curves corresponding to lover pressures. This result is, of course, to be expected on the basis of Fig. 33. However, in the upper right-hand nortion of Fig. 34 the curves lond to converge. The curves corresponding to the lover pressures, 20, 40, and 100 am Hg, do not extend no far to the right no do the curves corresponding to the higher pressures. Therefore, one can only estimate where the love-pressure curves would occur relative to the high-pressure curves in the right-hand pertion of the figure, but it would appear that if the love-pressure curves curves extrapolated, they would meanly converge with the nigh-pressure curves.

Complete convergence of the curves in Fig. 3h would indicate validity of the strong-line approximation. On the basis of the curve experimentally obtained, it was be concluded that the strong-line experimentally obtained, in white of you large that the excreption of the basi is meanly complete and the deprendence of total absorption on both y and F<sub>p</sub> is small.



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If one is willing to allow as much as  $\pm 10$ -15% error, it is possible to derive from the curves of Fig. 34 some rather simple equations which relate total absorption to one variable  $\Psi_0$  for values of  $\int A_1 \nu / 4 \nu$  not included by Eq. (34). The ampirical equations which have been derived, and the region of validity for each is as follows:

$$\int A(v)dv = 10h \ i \ 61 \ \log(vP_0),$$
 for 
$$120 < \int A(v)dv < 200 \ cm^{-2}$$
 and 
$$20 < P_0 < 5000 \ cm \ Rg. \tag{56}$$

It is emphasized that the total absorption of a given semple can be determined more accurately from the curves of Fig. 32 than from the empirical equations, which are included as that it might be possible to make determinations without the use of the curves. The empirical Eqs. (34), (34), and (36) are seen to be useful over rather wide ranges of values of y and  $F_0$ ; and they are believed to give values of  $f_0(y)$  accurate to within 110f except for values of y and  $F_0$  near the limits of the region of validity, where  $f_0(y)$  we given by the equations may be in error by as much as tips.

It is noted that Eqs. (35) and (36) have the time expected on the besix of a strong-line approximation, i.e., the total absorption is a function of the product  $w_0$ ; however, estitler of those equations nor Eq. (34) corresponds to the square-root relation, which absold apply in the special case of the strong-line approximation with no everlapping of the spectral lines. For the equare-root approximation the total absorption is proportional to  $v^{0.5}$  and  $\Gamma_0^{0.5}$ , but the maximum dependence on  $P_0$  which was observed was  $P_0^{0.5h}$ . The same qualitative explanation of the invalidity of the square-root relation given in the discussion of the E<sub>0</sub>0 band also applies to the C0 bank.

The characteries of R-man, and P-man, have been investigated in a manner similar to the mathet used for the Ryu bead. The values of observed absorption A(2168 cm<sup>-1</sup>) at R-man, for a major portion of the spectra are plotted in Fig. 35 equinat the equivalent pressure. Each nurve represents a given value of absorber consentration as indicated. It is noted that several of the curves have portions which are mostly straight with slope equal to approximately 0.56, indicating that the

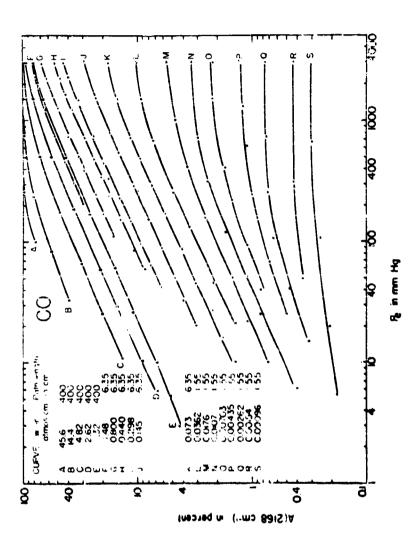


Fig. 35. The percent absorption at 2168 cm-1 versus equivalent pressure.

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absorption is proportional to  $P_c^{-0.48}$ , which is nearly equal to the  $P_c^{-0.5}$  dependence associated with the square-root approximation. The other features of the curves of Fig. 55 are similar to those of Figs. 17 and 20 which were described previously.

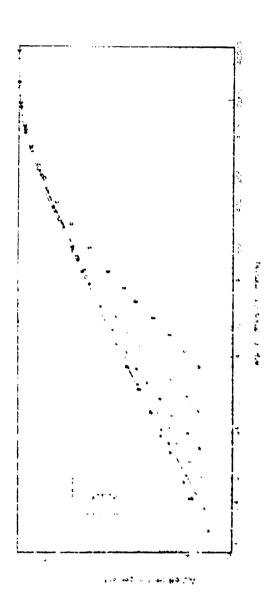
Values of percent absorption were taken from the curves of Fig. 35 and plotted in Fig. 36 with the variable  $vP_e$  as the abscissa. Each curve represents a given value of equivalent pressure, and validity of the streng-line approximation would be illustrated by a coincidence of the different curves. Fortions of some of the curves were emitted from the figure to avoid overcrowding; however, points were included which indicate where the curves would occur if they had been included. It is noted that the four upper curves corresponding to the lower pressures are nearly coincident over a wide range of values of the variable  $vP_e$ . These four curves also contain portions which are nearly straight with a slope of approximately 0.51, indicating a  $v^{0.52}$  dependence.

The results of these measurements can be compared to the absorption  $A(2115~cm^{-1})$  at P-max. In Fig. 37 are plotted the values of  $A(2115~cm^{-1})$  against equivalent pressure  $P_{\rm q}$ , where each curve represents a given value or absorber concentration. The gross features of the curve of Fig. 37 are, of course, sinitus to those of Fig. 35 which represent the absorption at R-max. The straight portions of the curves of Fig. 37 have a close of approximately 0.45, which is slightly less than the value of 0.45 in Fig. 35.

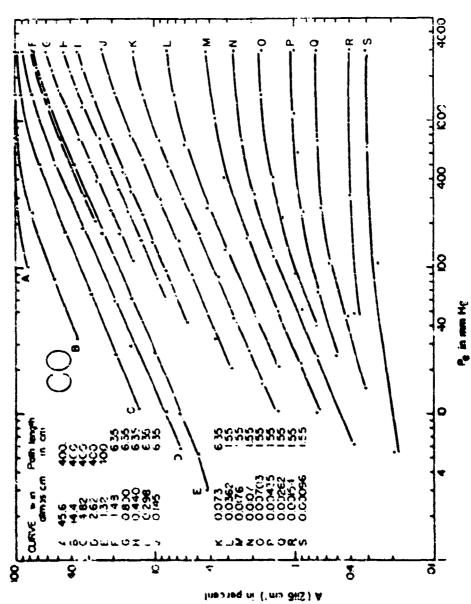
Values were taken from Fig. 37 and plotted in Fig. 38 with wP<sub>e</sub> as the sheciese. For the largest values of vP<sub>e</sub> the curves converge and approach 100% absorption. The straight pertions of the curves have a slope of approximately 0.52 as compared with 0.51 for the curves of Fig. 36.

Thus, for ceruin values of v and  $P_c$ , the absorption at P-max, and R-max. for the 21% cm<sup>-1</sup> CU hand above v dependence on these parameters which is only slightly different from the aguary-root dependence for which the absorption is proportional to  $v^{0.5}$   $P_c$  . As in the case of the absorption at R-max, and P-max, of the 27% -m<sup>-1</sup> MgO band, the maximum dependence on  $P_c$  is somewhat less the  $P_c$  0.5 dependence associated with the square-root approximation. However, the deviation, from the square-root dependence is considerably less for the CO bank than for the MgO band. It was pointed out in the discussion of the MgO band that the absence of a square-root dependence on pressure could be attributed to the west sines which occur along with the strong lines of the strong fundamental band. It is well known from high-resolution spectra than there are fewer weak limits in the CO bank than in the MgO band. It is therefore not currenter; that the deviation of the CO absorption from the square-root dependence is less than for the MgO absorption

It is further moted that the deviation from the square-root dependears of the abscrition at P-max, is greater than at R-max, for the CO band.



Pier Mr. The perient absorption at 235 cm<sup>-1</sup> vernus



Mg. 37. The percent absorption at 2115 cart versus equivalent pressure

Although the difference in dependence on  $P_e$  is small ( $P_e^{C,45}$  compared with  $P_e^{O,46}$ ), it is believed to be significant. Further evidence to the difference can be seen by comparing the straight portions of the four upper curves of Fig. 36 with the corresponding curves of Fig. 38. The curves of Fig. 36 occur considerably closer together than do those of Fig. 38; the difference in spacing indicates that the strong-line approximation is more nearly valid for R-max, than for P-max. This result can undoubtedly be explained by the presence of rather prominent lines of the isotopic band  $C^{1,2}$  old near P-max; the lines of this isotopic band are much weaker near R-max, and therefore give rise to less deviation from the square root dependence. The difference in the relative strengths of the isotopic lines in the two different spectral intervals is well known from nigh-resolution spectra and is a result of the shift of the band center of the  $C^{1,3}$  old band relative to the center of the band of the more common isotope.

Comparison of the straight politions of the four upper curves of Fig. 3h, which relates  $fA(\nu)d\nu$  to  $\nu P_{\sigma}$ , with the corresponding curves of Figs. 36 and 36 also indicates that the square root approximation is more valid for the absorption at  $F_{\rm cmax}$ , and  $F_{\rm cmax}$ , than for the total absorption. This result can undoubtedly be explained by the fact that the was times near the vings of the hand contribute to the total absorption by not to the absorption at  $F_{\rm cmax}$ , and  $F_{\rm cmax}$ .

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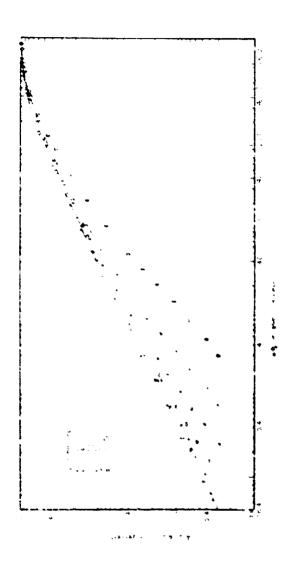
## B. THE 4260 cm 1 CO BAITD

The CO overtone band at \$260 cm<sup>-1</sup> was investigated by using a louble-pass, single-beam Model 99 Ferbin-Elmer spectroseter with a LiF prize. A serm-1 glover served as the source and a thermocouple as the detector. External adjustments on the multiple-traversal cell which was used, ands it possible to record spectra of each sample at \$, 6, ab and \$2 trave-rate -orresponding to path language of \$25, 1232, 2445, and \$675 cm, respectively. The spectra were replotted on a linear scale of percent absorption were under sech spectrum.

Fractings of several of the replotted spectra are shown in Figs. 39 and  $\delta\theta_3$  the spectral slits/dth of the spectromater was between 15 and 20  $cm^{-1}$  .

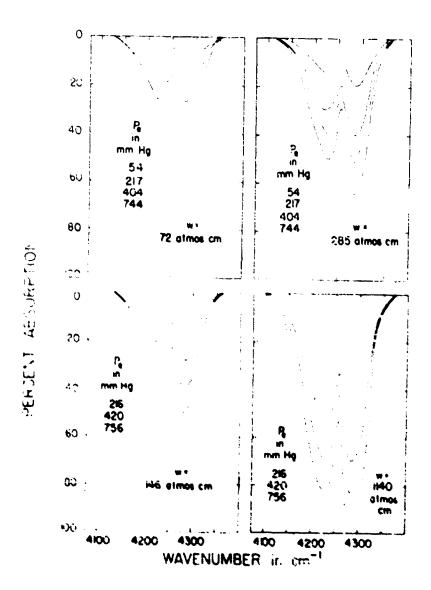
Table & lists the absorption parameters and the total absorption for all of the spectra obtained. The call was at room temperature, which varied by approximately  $k^{\rm QC}$  during the course of the experiment.

In order to determine the value of  $P_0$  for each sample it was first messessive to determine the value of the self-broadening coefficient B in Eq. (1k). Since the spectrumeter used for this portion of the study was not a double-beau type instrument, it was not possible to determine the value of D by the method described previously for the 2225 cm  $^{-1}$   $T_0O$  and 2145 cm  $^{-1}$  CO bands. The method of determine D for this band made



Mg. 13. The persent absorption at 211 carl versus who

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Fact the opening the active on a CO band

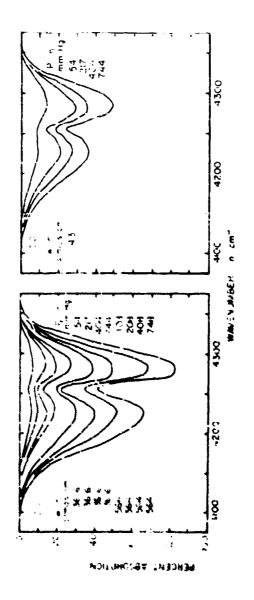


Fig. 36. Spectos of the 8:60 cs. 10 bend

use of the fact that several different samples which were studied nontained approximate'y the same absorber concentration at different path lengths. For example, an absorber concentration of approximately 145 atmos on was obtained with path lengths of 625, 1232, and 2445 on by using different partial pressures of CO in the absorption cell. Spectra of each of these samples were obtained at different values of total pressure which were produced by adding Io. The different samples, which contained the same absorber concentration but different partial pressures of CO, mosssarily required different amounts of \$2 in order to produce the same total absorption. By comparing the different values of CO partial pressure and H2 partial pressure required to produce the case total absorption by samples having the same absorber concentration, it was possible to determine the value of B according to Eq. (15). Approximately 20 separate calculations, which involved a large portion of the data obtained for the \$260 cm<sup>-1</sup> CO bank, were made. These calculations yielded a value of B = 1.00, which is believed accurate to 166. In accordance with Eq. (14) the value of P. was determined by

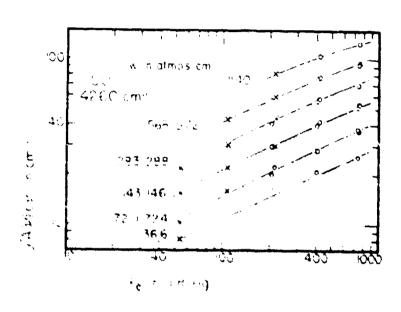
$$P_0 = P + 0.08 \, p_0$$
 (57)

where P is the total pressure and p is the partial pressure of C).

The value of B=1.0% is somewhat larger than the value of 1.00 determined for the fundamental CO band. The two values differ by slightly less than the maximum experimental error; however, it would not be surprising to find a significant difference in B for different bands of the same gas since variations were found within the 2224 cm<sup>-1</sup>  $M_2O$  cml lime 2143 cm<sup>-1</sup> CO bands. The method used to determine the value of B for the 2340 cm<sup>-1</sup> CO band did not make it possible to intest small variations of B within the band.

The data for the 1260 cm<sup>-1</sup> CO hand are shown in Fig. 11 with total absorption plotted as a function of equivalent pressure. The top and better curves represent single values of absorber conscatration. Such of the other curves correspond to an approximately constant value of absorber conscatration; the values indicated at the left of each curve indicate the maximum and maximum values of absorber rescentration represented by that curve. The a supersent complete of CO alone, while the "O"s represent complete composed of both CO and Ep.

Figure hi relates the same parameters on does Fig. 30 for the fundamental CO bank. However, Fig. hi does not represent searly as vide a reaso of values of w and P<sub>a</sub> as does Fig. 30. Therefore, some of the deflatts features of Fig. 30 are not present in Fig. hi, such as the decrease in slope at small values of P<sub>a</sub> as a result of Deppler broadening and at large values of P<sub>a</sub> as a securit of overlapping of the spectral lines. The crowding of the curves in the appear right-hand sormer size does not occur in Fig. 41, since the largest values of w and P<sub>a</sub> were not



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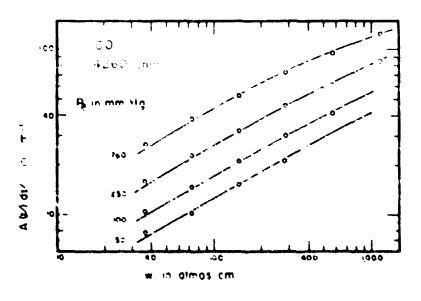
sufficiently large to produce complete absorption over the band.

In order to show the dependence of total absorption on  $\underline{w}$ , values were taken from the curves of Fig. 41 and plotted in Fig. 42 for the four different values of  $P_0$  indicated. This figure is similar to Fig. 32, which was drawn for the fundamental 30 band: but as in the case of Fig. 41, this figure does not contain all of the features of Fig. 32 because of the much smaller range of values of  $\underline{w}$  and  $P_0$  which are represented.

By comparing the points in Fig. 42 with the curves of Fig. 32, it was found that at a given equivalent pressure the total absurption of the 2145 cm 1 band was very nearly the same as the total absorption of the \$200 cm - band at the same equivalent pressure but with the value of absorber concentration 150 times as great. In fact, each of the curves or Fig. 42 was drawn by tracing the curve of Fig. 32 corresponding to the same value of Po but with the value of w different by a factor of 150. It is noted that the experimental points fit aircely the traced curves. This direct tracing of the curves with the value of the abscissa different by a constant factor was possible, of course, because of the fact that the abacterse are ingertibute. The curve in Fig. 42 which corresponds to an equivalent pressure of 50 mm Hz was not traced from Fig. 32, as were the other three curves, sluce Fig. 32 does not contain a curve corresponding to this pressure. The curve corresponding to 50 m Eg in Fig. 43 was therefore drawn from another similar curve shich corresponds to 50 cm Mg for the 21h5 cm- band but is not included in this report.

The total absorption of the 4260 cm<sup>-1</sup> hand for samples in the "linear region" (6') cannot be determined from the surves of Fig. 42 since the largest values of P<sub>0</sub> were not sufficiently large, nor was the samilest value of y sufficiently small. Sowever, Fenner and Veber<sup>21</sup> investigated the total absorption of both bands in the linear region and found that the value of 8(v)Mv for the fundamental band was approximately the times that of the overtoon band. The forms channel is the linear region in therefore the case as that of the fundaments, with the value of the fundaments, with the value of the fundaments 1/14% as great.

On factor of the to within the cimits of experimental error of the factor of the, which was observed for a different range of values y and for discretible two bands are similar with regard to line stage. The against, and relative time attracted within the band, it is probably true that the total absorption of the weater overtone band can be determined from the cases, of Fig. 5, which device with the fundamental band, by using the same cause of Fig. and a value of which is 1/150 times that of the cample whose wise absorption is being described. Equations (th), (15), and (16) could probably also be applied satisfacturily by which is 1/10 in the equation. Startween appears.



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## VIII. HER JACO, 1500, AND 1000 com By PANTAG

It has been known for quite some time that CH<sub>a</sub> copies in rather large quantities in the atmospheres of the planets Jupiter, Saturn, Uranus, and Reptune. Not until 1948 was CH<sub>a</sub> discovered by Migestia to be a permanent constituent of the earth's atmosphere; the average abundance has since been found to be approximately to parts per million by volume.

Approximately 190 absorption ujectra of CD<sub>2</sub> were obtained in the spectral region from 5400 - 700 cm<sup>-1</sup>. Absorption path lengths of 6.35, 400, and 1600 cm were used to obtain values of absorber concentration from 0.015 to 188 atmos cm. Values of total pressure were varied from approximately 5 to 3000 cm. Fig.

Virtually all of the accorption occur, in two different intervals, one between \$400 and \$200 emml, and the other between 1750 and 1160 cmml. Tracings of many of the spectra of the high-frequency region are presented in Figs. 43-46 and spectra of the low frequency region are shown in Figs. 47-51. Portions of some of the spectra were emitted from the figures to proved occurrenting, and to was accessary to correct season portions of sume of the absorption curves for overlapping absorption by  $00_2$  and  $8_20_2$  which occurred as impurities in the CM, and in the air in the optical math of the spectrumeter. For some of the samples containing very large values of absorber concentration, the fin, staorgiller extended into the region of the 2350 cm-1 CO, band, and because of the CO, impurity in the CH, it was necessary to correct for CO, absorption between 2450 and 2200 cm-1. It was also necessary to correct for somrious  $\rm H_2O$  absorption in the region between 1750 and 1500 cm<sup>-1</sup> for samples of large v. The effective slit width varied throughout each absorption curve with an average value of approximately 25 cm-1 in the high-frequency region and 10 cm-1 in the low-frequency region.

For purposes of analysis the 1750 - 1100 cm<sup>-1</sup> region was divided into 2 regions of 1555 cm<sup>-1</sup>. For indentification purposes the high-frequency region is referred to as the 1550 cm<sup>-1</sup> band and the low-frequence region as the 1506 cm<sup>-1</sup> band, since these too frequencies correspond to the centers of the two stronger bands in the region. Similarly the spectral region from 5400 - 2200 cm<sup>-1</sup> in referred to as the 3000 cm<sup>-1</sup> band. Since the region referred to as the 1550 cm<sup>-1</sup> band occurs in the region of the water vapor absorption, the accuracy of the traced spectra and the values of  $\int A(v) dv$  for this band are semawhat less than for the other bands.

The CH<sub>1</sub> data corrected for spurious absorption are given in Table 7, which includes values of fA(v) for all three hands for the listed values of v and  $P_{e}$ . The value of the self-breedening coefficient B was found to be 1.30 for the Julius manifold and 1.38 for the 1900 cm<sup>-1</sup> band. I was not possible to obtain an accounter measurement of B for the 1550 cm<sup>-1</sup>.

It has been known for quite some time that CH<sub>0</sub> occurs in rather large quantities in the atmospheres of the planets Jupiter, Saturn, Uranus, and Neptune. Not until 1948 was CH<sub>0</sub> discovered by Migeotte<sup>22</sup> to be a permanent constituent of the earth's atmosphere; the average abundance has since been found to be approximately 1.6 parts per million by volume.

Approximately 190 absorption ejectra of CH<sub>b</sub> were obtained in the spectral region from 3400 - 700 cm<sup>-1</sup>. Absorption path lengths of 6.35, 400, and 1500 cm were used to obtain values of absorber concentration from 0.015 to 188 atmos cm. Values of total pressure were varied from approximately 5 to 3000 mm Hg.

Virtually all of the absorption occurs in two different intervals, one between 5400 and 2200 cm-1, and the other between 1750 and 1160 cm-1. Tracings of many of the spectra of the high-frequency region are presented in Figs. 43-46 and spectra of the low frequency region are shown in Figs. 47-51. Portions of some of the spectra were omitted from the figures to provent everowiling, and it was necessary to correct small portions of some of the absorption curves for overlapping absorption by  $00_2$  and  $00_2$ , which occurred as impurities in the CH<sub>0</sub> and in the air in the optical path of the spectrometer. For some of the samples containing very large values of absorber concentration, the CEL absorption extended into the region of the 2350 cm-1 CO2 band, and because of the CO2 impurity in the CH4 it was necessary to correct for CO2 absorption between 2450 and 2200 cm-1. It was also necessary to correct for sparious H<sub>2</sub>O absorption in the region between 1750 and 1500 cm<sup>-1</sup> for samples of large v. The effective slit width varied throughout each absorption curve with an average value of approximately 25 cm-1 in the high-frequency region and 10 cm-1 in the low-frequency region.

For purposes of analysis the 1750 - 1100 cm<sup>-1</sup> region was divided into 2 regions at 1535 cm<sup>-1</sup>. For indentification purposes the high-frequency region is referred to as the 1550 cm<sup>-1</sup> band and the low-frequency region as the 1500 cm<sup>-1</sup> band, since these two frequencies correspond to the centers of the two stronger bands in the region. Similarly the spectral region from 5400 - 2700 cm<sup>-1</sup> is referred to as the 3020 cm<sup>-1</sup> band. Since the region referred to as the 1550 cm<sup>-1</sup> band occurs in the region of the vater vapor absorption, the accuracy of the traced spectra and the values of  $\int A(\nu) d\nu$  for this band are somewhat less than for the other bands.

The CH, data corrected for spurious absorption are given in Table 5, which includes values of  $\int A(v) dv$  for all three bands for the listed values of v and  $P_e$ . The value of the self-broadening coefficient B was found to be 1.30 for the July man band and 1.38 for the 1500 cm<sup>-1</sup> band. It was not possible to obtain an accounts measurement of B for the 1550 cm<sup>-1</sup>

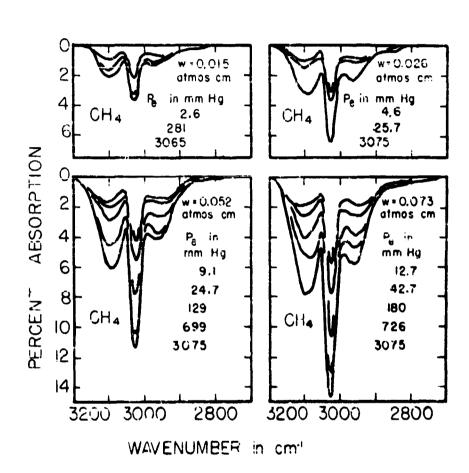


Fig. 45. Spectra of the 3020 cmr4 CM; hand

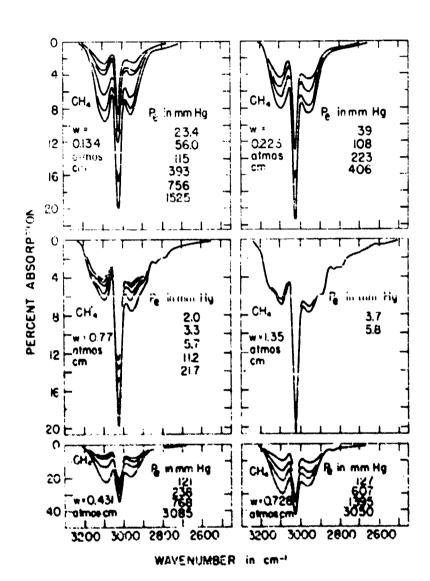


Fig. 44. Species of the 3000 cm-1 CEC band

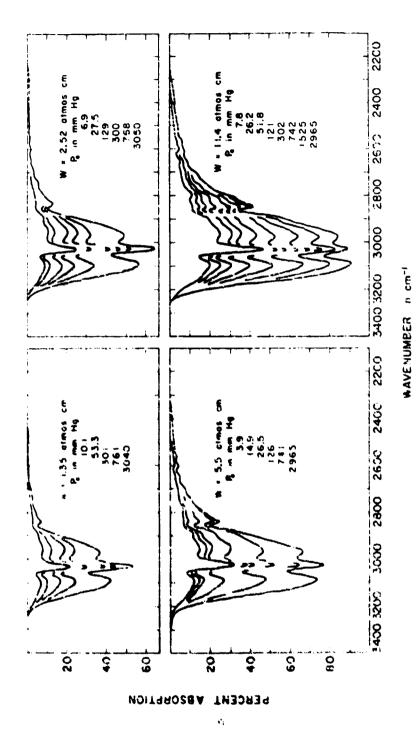
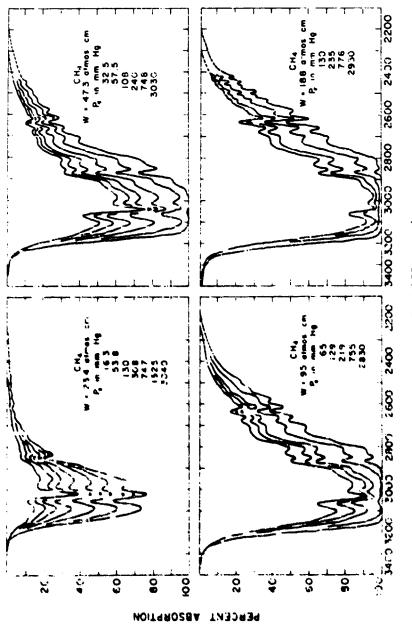
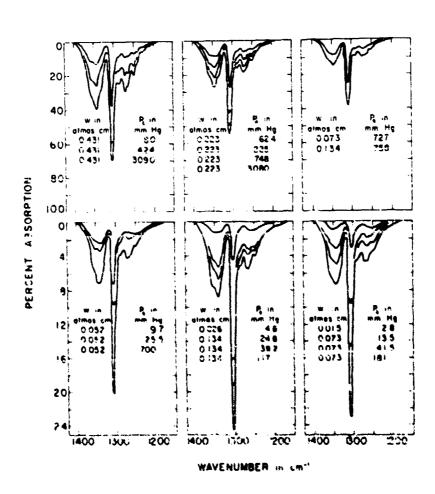


Fig. 4.1. Springs of the CO arms Mr. Sand



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WAVENUNBER in crit



Pin. 87. Spentin of the 1506 out? The band

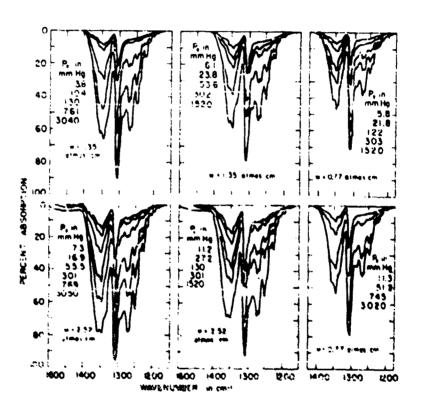
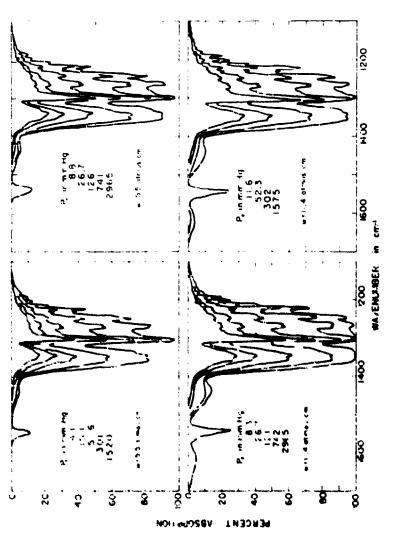


Fig. 30. Spectra of the 1305 cm<sup>-1</sup> CH<sub>4</sub> band



Piles 49. Spectors of Se 170 or 1 and 1930 or 5

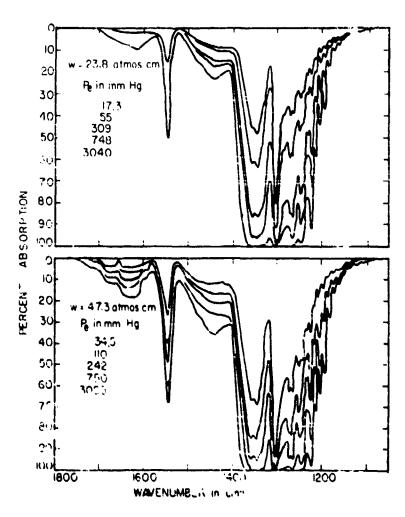


Fig. 40. Sprinting, the their manageria is  $^{\circ}$  Tily bands

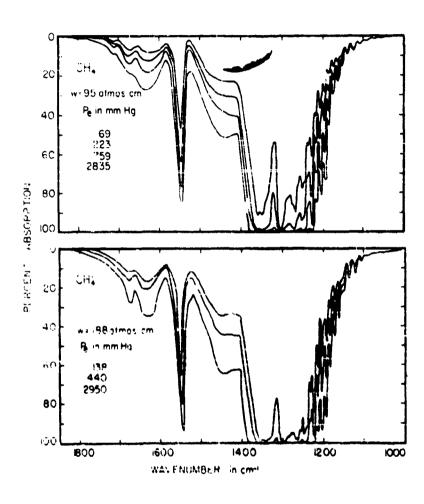


Fig. (1. Spectra I the iffe of and 1.50  $\rm dm^{-1}$   $\rm dH_{\odot}$  ) and s.

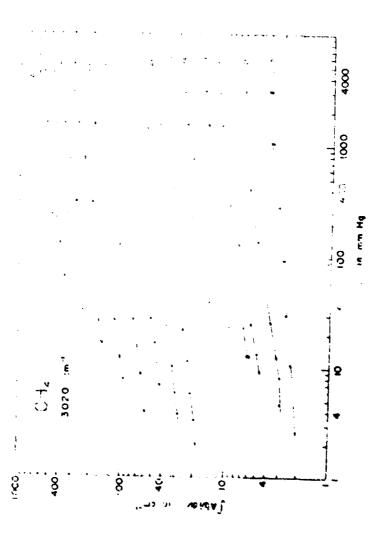
band because of the over-lapping water vapor absorption; therefore, the same value of B, 1.35, was used for this band as for the 1306 or the land. Since two different values of B were used in (14), two different values of  $P_{\rm e}$  appear in Table 5 for each sample.

The data on the CH<sub>h</sub> absorption are presented in a manner similar to that in which the 2224 cm-1 N<sub>2</sub>O and 2145 cm-1 CO data were presented in previous sections. Many of the 3020 cm-1 tata are plotted in Fig. 52 with fA(v)dv versus P<sub>D</sub>, where each curve corrusponds to a constant value of w. The features of the curves of Fig. 52 are similar to those of the curresponding curver for the 2224 cm<sup>-1</sup> N<sub>2</sub>O band in Fig. 12 and the 2143 cm<sup>-1</sup> CO band in Fig. 3C. However, two features of the curves of Fig. 52 should be noted. The maximum slope of any of the curves of Fig. 52 should be noted. The maximum slope of any of the curves of Fig. 52 is approximately 0.23; this indicates that the maximum dependence on P<sub>E</sub> is P<sub>E</sub>O-23, which is considerably less than for the N<sub>2</sub>O and CO bands. It is also noted that the slopes of the curves corresponding to values of v between 0.11 and 11.4 atmos-cm increase with increasing values of equivalent pressure to approximately 75 mm Hg.

It is recalled that the slopes of some of the corresponding curves for  $N_2O$  and CO were found to increase with increasing equivalent pressure up to approximately 10 mm kg. This phenomenon was attributed to the effect or the Dippler width of the absorption lines. However, it seems unlikely that an increase in the clope to pressures as high as 75 mm Eg could be a result of the imppler width of the times. It seems, therefore, that another explanation of the increasing slope of the curves of Fig. 52 is needed. A possible qualitative explanation which takes into account the rather unique structure of this band follows.

A high-resolution spectrum of the just on "I CH, band shows that it is composed of several strong lines which occur in groups, each group being separated by approximately 10-12 cm<sup>-1</sup>. Also occurring in the spectrum are many seeker lines throughout the band. Except for samples with low pressure and very small values of absorber concentration, each group of strong lines could be treated as a single line for purposes of estimating the absorption. The band could then be approximated by a band model consisting of an Eleasser-type band of strong lines spaced 10-12 cm<sup>-1</sup> spart, and superimposed on it conther band consisting of much weaker lines spaced with the proceed such closer together.

For relatively low pressures one might expect that there would be a rather wide nince of values of w over which the contribution to the absorption by the strong lines would have approximately a square-root dependence. The lotal absorption by these lines would therefore increase approximately as the square-most of the equivalent pressure. On the other and, the contribution to the absorption by the weak lines would have sittle or no dependence on equivalent pressure. At low pressure a sajor position of the chaorytion would arise from the weak lines; and the total absorption of the band would therefore have little dependence on pressure. However, as the pressure is increased, the relative



Pig. 52. The total control of the 3000 man

contribution of the pressure-dependent strong lines would become greater; the total absorption of the band would therefore exhibit the observed increasing dependence on pressure.

Without investigating the absorption in a highly resolved spectrum, it would be difficult to prove that the above explanation is valid. Towever, it can be said that it is a possible explanation of r dependence on pressure not observed in most absorption bands studied, and is based on a band model which is similar to the actual band.

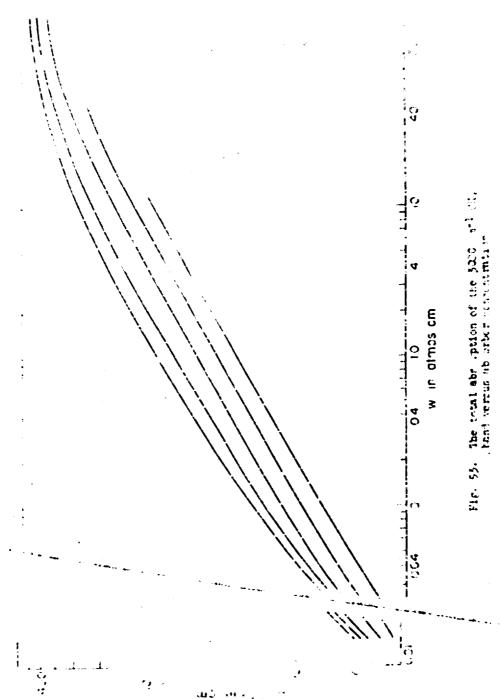
Values of  $\int A(v)dv$  were taken from the curve of Fig. 52 and plotted in Fig. 55 against absorber concentration v. Each curve corresponds to the constant value of  $P_{e}$  as shown. By interpolation it is possible to determine the total absorption of a sample of CR<sub>0</sub> having values of v and  $P_{e}$  within the range included by the curves of Fig. 53. The features of the curves are similar to those of the curves are figures for the 4.0 and CO bands. The nearly linear portions of the curves were found have a spea of approximately 0.55, indicating that the total absorptions is  $v = \frac{1}{2} \frac{$ 

'be determination of the intendity ares to enticiausty pitty from the newconsecon involves "ently low so that the m, items matten sucwier three con-Tuesty in very small i-marti-· womilest ditions the . values of w and line . term Sign absorption is ween to incipa--To as developed apparent that the band intemptly cour. MISWAY 12 a direct reading from the JUNU mm Mg cur. the corresponding figures for the \$50 and Co and a computating the curve to smaller values of y, beening is a we remain alone of the extrapolate fortion of the curve about he the value of the band intensi" - 3020 cm-1 CM, band was estimated to be riven by

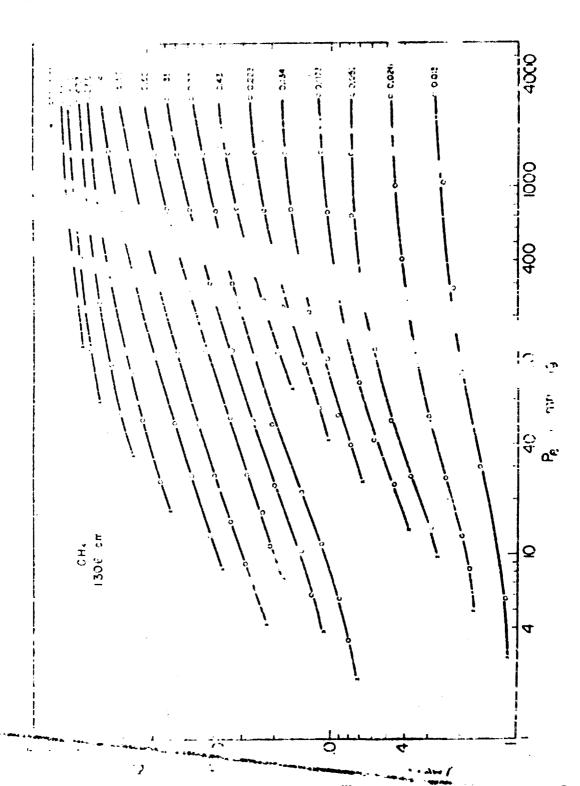
$$\int k(v)dv = 320 \text{ atmrs}^{-1} \cos^{-2} ST2 . \tag{38}$$

This value, which is believed to be accurate to within 215%, is approximately 12% last than the value of 360 states of car2 reported by Welab, Packler, and Dung43.

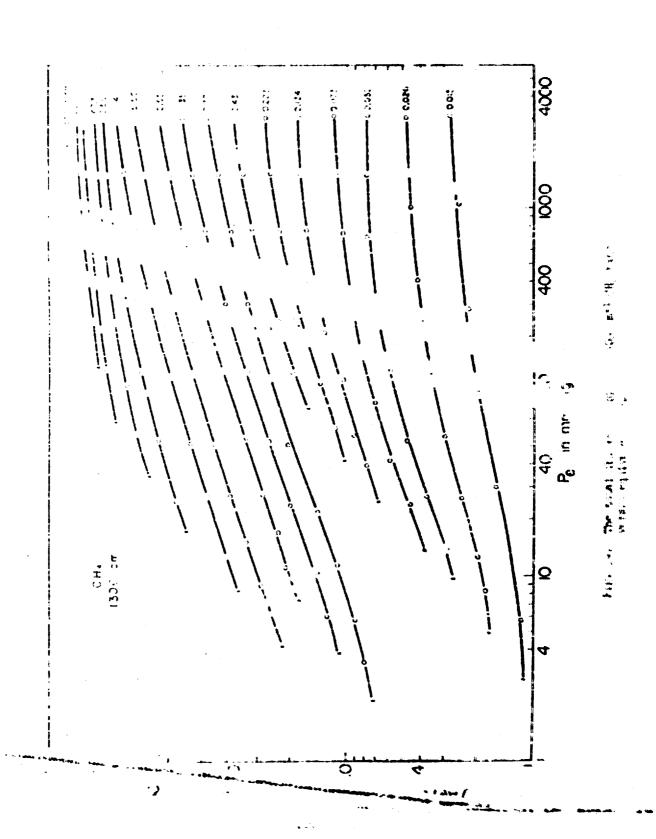
Values of  $\int A(v)dv$  for the 1305 cm<sup>-1</sup> CM, hand are plotted against  $P_0$  in Pig. 54. Several of the curves of the figure contain portions oring slopes of approximately 0.30, indicating a maximum dependence of  $P_0$  for our  $P_0$  for the first thin the maximum dependence on  $P_0$  for our Maximum terms of the containing of the conta



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oren: for values of equiva to the court of the supervent to the supervent in the majories of the section of the Pig. 52 for the 3020 cm<sup>-1</sup> band. Right-resolution of the community bands show that their band structures are product at the simple band model which was used to describe the growth of the band would not be expected to securify for the 1500 cm<sup>-1</sup> band. would therefore not necessarily expect the two limits on arms with the creasing pressure in the same manner.

Values of  $\int A(v)dv$  were obtained from the ... even of Fig. 5% and plotted egainst absorber concentration to Pig. 59 for the different volues of Po indicated. Rather large portions of the curves were found to have the or approximately 0.48, indicating that the total absorption varies essentially as the square root of the atsurcer contentration for the values of y and P, represented.

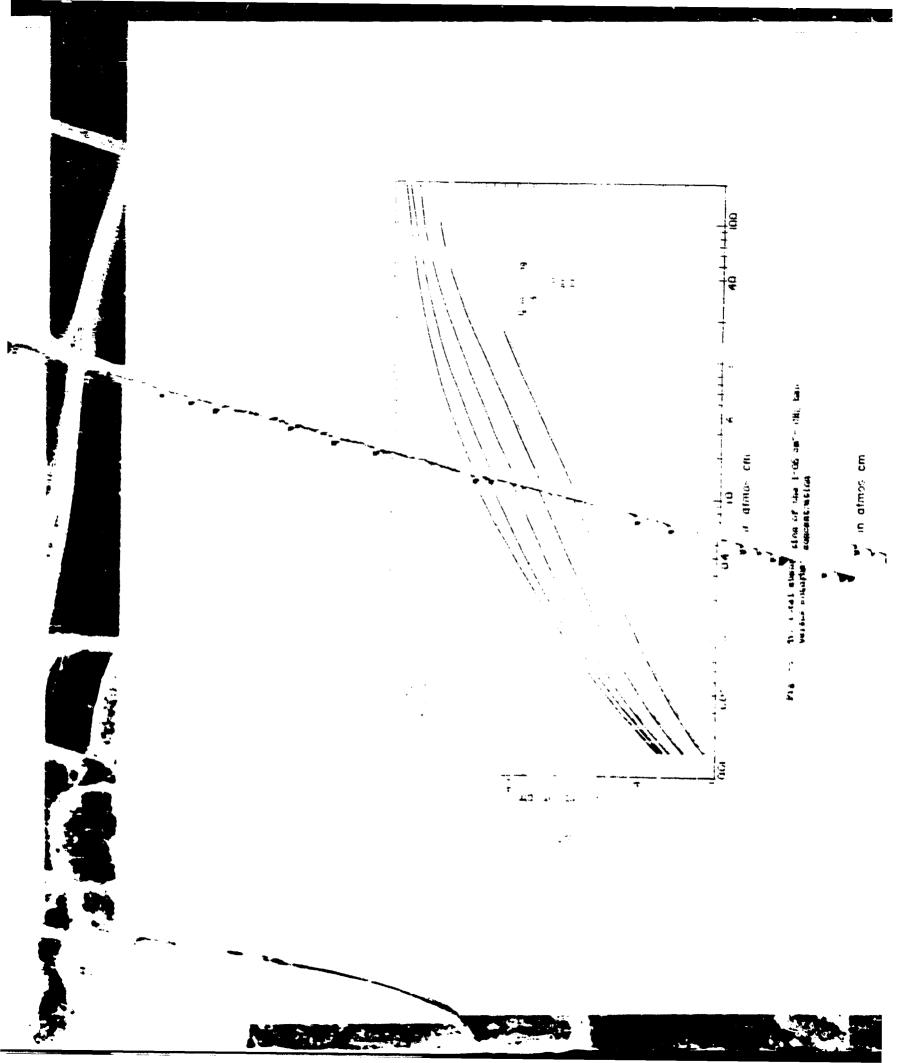
By extrapolating the corres of Fig. 55 to smaller values of w, the band intensity of the 1305 mail CH, band was found to be given by

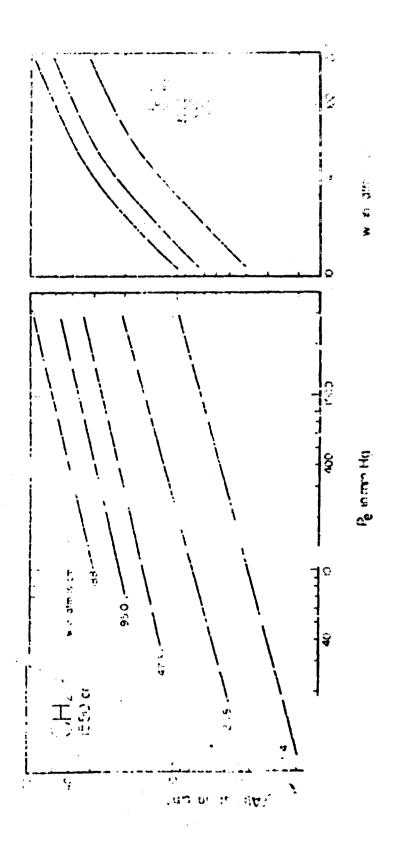
> $\kappa(v)dv = 105 \text{ ntmos}^{-1} \text{ cm}^{-2}$ ini

Tais value, which is be neved to the secretary east than 1154, as approximately 10% greater than the wine of 157 atm == cm=2 reports by Welsh and Sandiford24.

The absorption by the .500 cm-1 tand overlaps the absorption by the " cm-1 band for camples of large absorber concentration. The frequency Thosen as the frequency dividing the two bands represents an between the P- and Q-branches of the we band at region referred to in this report as the 1306 cm. of the absolute and consentially unaffected competed with tron the volume. This is believed to be to the transfer to the william by the bands in the region where the bands over the negligible for the region. values of absorber or entraises value are regress, is by use often of the curves from vo to the sand intensity was determined

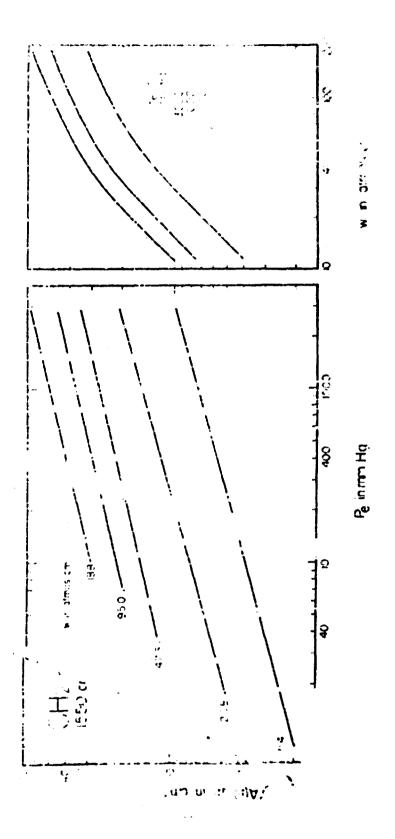
The data for the 1950 cm 1 from band are platted in Mg. 56. The total absorption a street against Pe in the left-rend portion for the values of windfoot . and values were taken from these curves to extain the curves in . right-and portion of the figure. As stated proviously, the dat for this must were itsited to eccules having large values of absorber continuition, and the recording of the spectra was less accurate than to the other heads because of the overlapping absorption by water wap. The curves of Fig. 36 are proposity a curete to 211% for values of the elementary greater than 50-40 as it; the the error for the same in samples might be no great as 215-20%.





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g. 35. the total atsorption of the 1855 cm<sup>12</sup> Ch. tena

The slopes of the curves in the left-hand portion of Fig. 56 indicate a maximum dependence of approximately  $P_e^{0.27}$  for the values of v and  $P_e$  represented. This result is approximately the same as that observed for the other two  $CH_b$  bands. The curves in the right-hand portion of the figure do not exhibit regions having aloyes of approximately 0.5, as do the corresponding curves for the other bands, since the largest values of absorber concentration employed were not sarficiently great for the lines in this weak band to exhibit "strong-line" properties.

Because of the absorption by  $\rm H_2O$  and the limited amount of data obtained for the 1550 cm<sup>-1</sup> band, it was not possible to determine the band intensity with very high accuracy. The value of  $\rm jA(\nu)d\nu$  indicated by the ordinate of Fig. 56 of course does not include absorption by the F-branch of the voband, since only the absorption at frequencies greater than 15%5 cm<sup>-1</sup> is included. Therefore, any value of absolute intensity determined by extrapolating the curves of Fig. 56 would be too low. However, by extrapolating these curves and by answers that the contribution of the F-branch in apper fixately the sums as the R-branch it can be estimated that the absolute intensity of the  $\nu_0$  bend is between 2 and 3 atmos<sup>-1</sup> cm<sup>-2</sup>. To the knowledge of the authors, the intensity of this weak band low not seen seasoned previously.

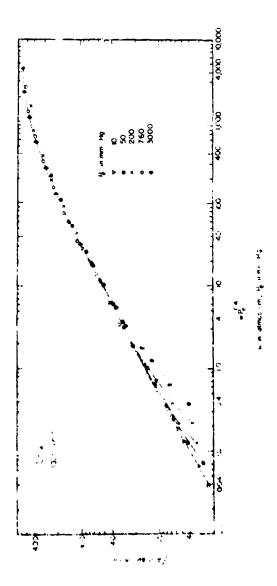
In order to obtain more accurate values for the interaction of the CE, bands, higher values or equivalent pressure about a be used on a are data should be obtained for simples having small values of specified or contention. In the case of the 1550 cm<sup>-1</sup> had, the CE, st. should be dried and a spectrometer which could be evacuated or thoroughly dried should be used. This extra care was not taken since the determination of band intensities was only of recondary importance in the present study.

In order to obtain ampirical equations to relate the total absorption of the 3020 cm<sup>-1</sup> tend to the parameters v and  $P_{\rm H}$ , the curves of Fig. 57 were drawn with the variable  $vP_{\rm H}^{\rm H, 0}$  as the shedisms. The value 0.5 was arrived at from the slopes of the linear portions of the curves of Figs. 52 and 55 to the serie manner as was done for the 2225 cm<sup>-1</sup>  $E_{\rm H}^{\rm H}$  band. The points corresponding to the lifterent values of pressure in the 57 occur very close together over a large region of values of A(v) are the v-distance over a large region of values of A(v) are the v-distance corresponding v- the different pressure were tacketed. S. a. Form 57 the following equations were derived for the A(v) cm<sup>-1</sup> till band.

$$\int A(v)dv = 15.5 \left[ M_0 u.h. \right]^{C_{11}}$$

$$15 < \int A(v)dv < 250 ce^{c\lambda}$$

$$10 < P_1 < 160 ma mg.$$
(40)



57. The total absorption of the  $_{\rm JCSO~cm^{-1}~CH_{11}}$  band versus wPe

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$$\int A(v)dv = -375 + 272 \log v P_e^{0.4}$$
 for 
$$250 < \int A(v)dv < 500 \text{ cm}^{-1}$$
 and 
$$10 < P_e < 3000 \text{ mm Hg}. \tag{41}$$

Values of  $\int A(v)dv$  determined by use of the above equations are probably accurate to +10%, except for values of  $\int A(v)dv < 40$  cm<sup>-1</sup>, below which the error might be as great as ±15%.

Figure 58 shows the total absorption of the 1506 cm<sup>-1</sup> CH<sub>h</sub> band as ordinate and wPe<sup>0.0</sup> as the abscicsa. The curves are not as close as in portions of Fig. 57. It is possible that a somewhat different value of the exponent of Pe could be used to cause the curves more nearly to coincide; towever, the various curves in Fig. 58 dc not form the relatively elmin pattern occurring in the corresponding figures discussed earlier; for example, a portion of the curve corresponding to 10 mm Hg lies between the curves corresponding to 50 mm Hg and 200 mm Hg. It is apparent that an increase in the value of the exponent of Pe would tend to cause the two curves corresponding to 760 and 50 mm Hg to occur those rogether; but on the other hand, the curves corresponding to .0 mm Hg would be further separated. The value 0.6 was used as a nominal value for the exponent.

Since the curves of Fig. 98 are not as close together as the corresponding curves for other bands, it is probable that any empirical equations derived from them would be less accurate than the equations derived for the other bands. The following equations which were obtained for the 1306 cm<sup>-1</sup> CH<sub> $\mu$ </sub> band are believed to be capable of yielding correct values of total absorption to  $\mu_1$ thin 1156 for the conditions specified:

for 
$$\int A(v)dv = 7.5 \left[ vP_0 0.6 \right]^{0.45}$$
 and 
$$20 < \int A(v)dv < 150 \text{ cm}^{-1}$$
 and 
$$10 < P_0 < 760 \text{ on } 26 \text{ .}$$
 (42)

$$\int A(v)dv = -190 + 117 \log w v_w^{0.6}$$
 for 
$$150 < \int A(v)dv < 500 \text{ cm}^{-1}$$
 and 
$$10 < 9 < 160 \text{ ma My}.$$
 (45)

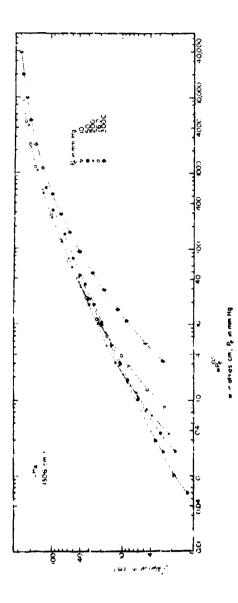


Fig. 56. The total absorption of the 1506  $z\pi^{-1}$   $zR_{\mu}$  band versus vPeO.6

## VIII. SUMMARY

The work povered in the present report has dealt with the infrared absorption of the minor atmospheric gases nitrous oxide, carbon monoxide, and methane. Samples consisting of these gases alone and in binary mixtures with mitrogen have been investigated over wide ranges of absorber concentration w and total pressure P. From the observed absorption for each band, it has been possible to present curves showing total absorption  $\int A(v)dv$  for various values of absorber concentration as a function of equivalent pressure Pe, which is a parameter that includes the total pressure P and a small additional term proportional to the partial pressure of the absorbing gas; the magnitude of the additional term is also dependent on the value of a "self-broadening coefficient," defined in the report and measured experimentally for each absorption band. 'other set of curves for each band shows the total absorption for various equivalent pressures as a function of absorber concentration. The curves provide a basis for the prediction of total absorption for any values of w and Pe included in the wide range of these parameters covered in the study. Extensive tables included at the end of the report provide additional absorption data for use by other investigators who may be interested in atmospheric transmission or in more general problems of accorption.

The experimental results are compared with various theoretical predictions. For values of v and  $P_0$  for which the "strong-line approximation" predicts the dependence  $\int \overline{A}(v) \ dv \sim \sqrt{vP_0}$ , it has been found that  $\int \overline{A}(v) \ dv$  is indeed roughly proportional to  $\sqrt{v}$  but not to  $\sqrt{V_0}$ . For sufficiently large values of P, overlapping of lines causes total bend absorption to approach a "saturation value"; for very low values of P, the total absorption for a given absorber concentration also approaches a limiting value for which the Doppler broadening of lines, independent of pressure, predominates.

It has been possible to obtain empirical equations relating total absorption to  $\underline{\nu}$  and  $P_e$  for certain limited values of total absorption  $\int A(\nu) d\nu$  and equivalent pressure  $P_e$ . The limits of validity for such equations are discussed. The empirical equations are summarized in Table (.

By studying total absorption at small values of absorber concentration w and sufficiently large values of equivalent pressure  $P_{\theta}$ , it has been possible to obtain values for the band intensity  $\{z(v)dv\}$  for the various bands studied. Present results have been compared with the values reported by other investigators. The values obtained for band intensities are summerized in Table 8.

A second set of comparisons of the present results with existing theories has involved utudies of the absorption  $A(\nu)$  in spectral regions of maximum absorption in the P- and R-branches of various bands. In most cases it was found that  $A(\nu) \sim \sqrt{\nu}$  over certain ranges of  $\nu$  and  $\nu$  however, it was reach that the proportionality  $A(\nu) \sim \sqrt{\nu}$  was rarely

observed. The possible effect of weak lines adjacent to strong ones is discussed as one cause of the fathure of the "strong line approximation": weak lines are always present in the wings of absorption bands, but sometimes are not present near the regions of maximum absorption in the P- and R-branches.

It is hoped that the present results can be used as a basis for the development of adequate general theories of absorption in the infrared spectra of gases as well as in more immediately "practical" studies of the atmosphere.

Table 1. Materials

Gas	Supplier and Grade	Purity,	Impurities		
<b>#2</b> 0	Mathieson	99	Chiefly #2, traces of CO <sub>2</sub>		
CO	Mathieson C. P.	97	Chiefly CO2		
CEL,	Mathieson C. P.	99.0	Traces of Calle, No.		
<b>4</b> 5	Linde High Purity Dry	99.99	Max. of 15 grains of $E_0$ 0 per 1000 cu. ft.		

Table 2. Self-Broadening Coefficients for Minor Atmospheric Cases

Bood	Self-Brondening Coefficient 30
255# cm_, H <sup>5</sup> 0	1.12 2 0.07
2145 cm <sup>-1</sup> CO	1.02 ± 0.06
1260 cm-1 co	1.08 + 0.06
5020 cm <sup>-1</sup> 0M4	1.30 ± 0.08
1306 cm <sup>-1</sup> CE <sub>6</sub>	1.38 ± 0.08
1550 ca <sup>-1</sup> CE.	1.38 a 0.08

\*infined in Eqs. (1)) and (16), where Eq is the reference gas.

Table 5. Data for the 2224 cm<sup>-2</sup> SgO Band

The Pig. Ro. corresponds to the figure is which the spectrum of each sample is shown. Absence of Pig. No. indicates spectrum is not shown. Uncertainties of values of  $A(\nu)$  and  $fA(\nu)$  are discussed in text.

Sample.	Fig.	V (etmos ca)	Pe	JA(v)4v (cm <sup>-1</sup> )	A(2240)	A(2213)	Remairs
Ro.	No.	Intmos ca.	(= Hg)	(CE)			<del></del>
			Path Le	ngth = 1.5	5 cm		
1	5	0.0016	1.0	1,00	1.5	1.4	Samples 1-21:
2	55555555	0.0033	2.0	1.56	2.4	2.5	Ego alone in
3	Ś	0.0051	3.1	2.03	3.0	~ K	rel:
2 5 4	Š.	0.0074	4.6	2.75	4.0	5. č	
5	Ś	0.0110	6.8	3.79	5.5		
5 6 7 <b>E</b>	ś	0.0157	9.7	.21	7.5	<i>F</i> .	•
7	έ′	0.0217	13.4	6.93	9.7	9.5	
ř	ź	0.0270	16.7	8.49	17.0	11.2	
9	5 8 8	0 0352	21.8	10.9	15.3	14.6	
10	2	0.0456	28.1	13.6	19.2	18.2	
11	Ö	0.057	35.3	16.1	23.3	21.3	
12	8	0.073	<del>9</del>	20.2	28.1	26.t	
13	8	0.091	56.2	24.7	34.C	33.9	
14		0.0547	33.8	15.3	21.3	20.6	
15	5	C.0724	44.0	19.6	27.5	25.8	
16	5	0.112	69.4	28.2	39.5	<u> </u>	
17	5	0.172	106	40.2	55.0	52.1	
18	5	0.225	158	53.2	71.0	68.0	
19	5	0.389	241	69.3	ôŷ.1	85.0	
20	÷	0.677	419	85.2	96.0	98.3	
21	N555555	1.36	âŭî	96.8	99.0	100 0	·
23	8	0.0186	11.5	5.93	8.5	7.7	10.5 m Eg ef
25	8	0.0186	21.6	7.21	10.2	9.6	320
24	ã	0.0186	42.4	9.32	13.5	12.3	Samples 23-30:
25	8 8	0.0186	91.2	12.5	13.5 13.2	16.6	in added to
25	ê	0.0186	186	15.8	24.2	21.6	samples 22
2!	ā	o ก <u>เก</u> ิร์	*60	19. J	30.0	20.2	
28	ĕ	0.0186	749	22.0	36.0	31.8	
23	•	C.0186	1520	24.4	₹0.0	35. 9	
50	6	2.0185	3035	24.9		25, 19	
-		9.0105	יכל של	2	41.1	, i.e.	
<u> </u>	8	0.0101	6.2	3.62	5.1	4.4	5.5 ma Eg or
52	ō	0.0101	10.9	4.21	6.0	5.1	<b>%</b> 20
33	8	0.0101	26.ú	5.60	ê c	7.3	•
54	8 8 8	0.0101	50.0	6.06	30	وُ.ق	
55 55 57 58 57 58 59	8	0.0101	114	9.11	14.0	12.5	
36	8	0.0101	259	4.6	19.5 23.3	16.5	Samples 32 32:
77	ă	0.0101	746	14.4	23.2	20.6	In added to
<b>5</b> 8	8	0.0161	1520	15.4	25.4	22.6	maple 31
•20	8	0.0101	3035	15.7	25.€	22.9	
17	J	0.1202	,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,	A7-1	23.0	<b>44.7</b>	

Table 5. Data for the 2224 cm-1 NgO Band

The Fig. No. corresponds to the figure in which the spectrum of each sample is shown. Absence of Fig. No. indicates spectrum is not shown. Uncertainties of values of  $A(\nu)$  and  $fA(\nu)d\nu$  are discussed in text.

Semple	Fig.	V	Pe_	JA(v)dv	¥(5510)	A(2213)	Romarks
No.	No.	(atmos cm)	( <u>= Rg</u> )	(cm-1)			
			Path Le	ngth = 1.5	5 cm		
_							
1	555555555888	0.0016	1,0	1.00	1.5	1.4	Samples 1-21:
ė	5	0.0033	2.0	1.56	2.4	2.0	N20 alone in
5 4	5	0.0051	3.1	2.03	3.0	2.5	rell
4	5	0.0074	4.6	2.75	4.0	3.6	
5 6 7 &	5	0.0110	6.8	3.79	5.5	5.3	
6	5	0.0157	9.7	.21	7.5	6.8	
7	5	0.0217	15.4	6.93	9.7	9.5	
	5	0.0270	16.7	8.49	11.8	11.2	
9	5	0.0352	21.8	10.9	15.3	14.6	
10	5	0.0456	28.1	13.8	19.2	18.2	
11	8	0.057	35.3	16.1	23.3	21.5	
12	8	0.073	44.9	20.2	28.1	26.6	
1,3	8	0.093	56.2	24.7	34.0	33.0	
14		0.0547	33.8	15.3	21.3	20.6	
15	5	C.0724	44.0	19.6	27.5	25.8	
16	5 5 5 5 5 5	0.112	69.4	28.2	39.5	37.1	
17	5	0.172	106	40.2	55.0	2.1	
18	Ś	0.225	158	53,2	71.0	68.0	
19	Ś	0.389	241	69.3	69.1	85.0	
50	É	0.677	419	85.2	98.u	98.3	
21	ś	1.36	841	96.8	99.0	100 0	
22	8	0.0196					
	8	0.0185	11.5	5.23	8.5	7.7	10.3 m Mg of
25	9	0.0186	21.6	7.21	10.2	<b>3.6</b>	M20
5#	8 8	0.0186	42.4	9.31	15.5	19.3	<b>Semples 23-30</b> :
25	ŏ	0.0186	91.2	14.2	<u> </u>	16.6	No added to
26	9	0.0186	186	15.8	34.8	<b>8</b> 1.6	sembles 95
21	6	0.01 <b>6</b> 6	460	19. 1	٥.٠٠	80.K	
26	8	0.01 <b>8</b> 6	749	22.0	56.0	32.8	
29	_	0.0186	1520	24. 1	<u>₩0.0</u>	34. 9	
30	5	0.0186	3035	24.9	÷	36,8	
31	8 8 8	0.0101	6.2	5.62	5.1	4.4	5.6 m Mg of
5¥	ð	0.9101	10.9	4.21	6.0	7.1	N <sub>2</sub> O
33	8	0.0101	26. ú	5.60	8 3	7.3	•
بنو	8	0.0101	50.0	6.86	34.4	ý. 3	
52 55 55 56 57 59	8	0.0101	114	9.11	14.5	12.5	
36	8	0.0101	259	11.6	15.6	16.5	Bankler 39 39
37	4	0.0101	745	14.4	23.8	20.6	No added to
18	8	0.0101	1520	15.4	25.4	22. ń	nemple \$1
*^	8	0.0101	3035	15.7	25.6	22.9	

Table 5. (Continued)

mab y s	Fig.	Y	Pe	LV(A)GA	A(2240)	1(5513)	Remerks
no.	No.	(atmos cm)	( <b>=</b> 'g)	(cm <sup>-1</sup> )			DC:0248
40	7	0.0101	56.6	7.43	11,1	10.2	Sample 40;
41	7	0.0101	122	9.44	14.7	13.6	56 mm of mix
42	7	0.0101	26 <sup>4</sup>	11.9	19.0	16.9	ture of 1/10
45	7	0.0101	759	14.4	24.1	20.8	No0 and 9/10
44	7	0.0101	3035	15.5	25.8	23.0	No. Sam. 41-
	•	******				-2	44: #2 added
							to sample 40
45		0.005h	5.4	0.40			a
46		0.0054		2.32	3.3	3.2	Sample 45: 3
		0.0054	6.6	2.59	3.6	3.5	ma Hg of H20
47		0.0054	12.1	3.14	4.6	4.2	clone. Samp.
₩9		0.0054	50.4	1.2h	6.3	5.9	46-48: No add
							to sample 45.
49	7	0.0054	30. <b>4</b>	4.18	6.2	5.7	Sample 49: 30
50	7	o.mek	72.0	5.40	8.5		am Hg of 1/10
51	7	0.0054	189	7.02	$\mathbf{u} \in$	10.5	N. mixture.
52	7	0.0054	575	8.00	4 1.5	11.5	tracles (0-5)
53	7	0.0054	750	8.6	14.7	۲۰۰۰	So whaten to
54	?	2.10054	5035	9.05	15.7	12.6	####10 %9.
55		0.60%	2.1	1.44	2.1	1.9	Sample 55:
56		0.0032	4.2	1.73	2.0	2.3	1.75 m le ci
57		6.0032	9.9	2.09	5.2	2.9	No. Service
ýĠ		0.0032	17.7	3.61	4.0	3. ;	56-5. an
<u> </u>		0.0032	35.0	3.21	5.2	6.6	ad' id to sage
73		٠.٠٠	22.5	7.64	,. <b>.</b>	4.0	ple 55.
							P44 33.
60	1	v.w <b>y2</b>	11.7	2.55	3.8	3.5	Emple 60:
Ü	7	0.0032	40.4	3.21	5.1	4.5	17.5 m de es
62	7	0.0033	105	4.20	6.8	6.2	1/10 Be str.
65	¥	ପ୍ର, ମହକୃଷ୍ଟ	254	*.96	ð. •	7.3	ture. Complet
ė.	(	0.0012	755	5.40	9.5	7.7	61-651 📭
65	7	ھو⇔.ہ	<b>5</b> 035	5.57	9.7	8.5	added to sea-
							ple 60.
66	7	0.0020	11.1	1.59	2.5	2.2	Sample 66:
67	7	0.0020	25.1	2.07	3.2	2.8	11.1 mm Mg of
48	Ť	0.0020	40.C	2.64	1.8	5, 5	1/10 No mis
69	7	0.0000	106	9.147	1.9	1.2	ture. Ample
70	ż	0.3620	265	. 24	5.5	1.8	67-72: %
71	•	0.0000	770	1, 69	6.1	5.2	04404 to man-

Table 3. (Continued)

No.	Fig.	(atmos cm)	P. (ma Eg)	∫A(v)dv (cm <sup>-1</sup> )	A(2240)	A(2213)	Royal se
73	7	.تــــ0.00	6.1	0.98	1.5	1.3	Sample 75: 6.0
74	Ż	0.0011	15.6	1.22	1.9	1.6	mm Hg of 1/10
75	•	0.0011	39.9	1.54	2.4	2.1	No mixture.
76	7	0.0011	110	1.75	3.0	2.5	Samples 74-79
77	•	0.0011	259	1.90	3.5	2.8	No added to
78	7	0.0011	745	2.03	3.5	3.0	sample 73.
79	•	0.0011	3035	2.03	3.4	2.8	
80	6	0.0011	50.1	1.52	2,3	1.9	Sample 80:
91		0.0011	746	1.95	3.4	2.9	30.0 mm Hg of
82	6	0.0011	3035	2,00	3.5	3.0	1/50 N <sub>2</sub> 0 mix- ture. Sample 61-82: N <sub>2</sub> added to sam- ple 80
83	6	0.00078	20.1	1.03	1,6	1.4	Sample 83: 20
34	6	0.00072	98.0	1.21	2.1	1.7	HE of 1/50
85		0.00072	250	1.30	2,2	1.9	MyO mixture.
87 87	6	0.00072 0.00072	746 3035	1.30	2.3 2.3	1.9	Samples 84-87 No added to sample 83.
66	ó	0.00047	13.1	0.70	1.1	0.9	Sample 88: 13
89	6	0.00047	3035	0.86	1.5	1.5	mm Hg of 1/50 MgO mixture. Sample 89: Mg added to samp 88.
90	6	0.00033	9.2	0.55	0.8	0.7	Sample 90: 9.2
91	6	0.00033	3035	0.62	0.9	o.8	mm H <sub>0</sub> of 1/50 N <sub>2</sub> 0 mixture. Sample 91: N <sub>2</sub> added to sam- ple 90.
<b>92</b> 93	6 5	0.00024 0.00024	6.5 3035	0.37 0.35	0.7 c.e	0.5	Sample 92: 6.1 mm Hg of 1/50 NgO mixture. Sample 93. Ng added to sam- ple 92.

Table 3. (Continued)

			- 4020	). ( <del>•••</del>			
Sampla No.	Fig.	(atmos cm)	F <sub>e</sub> (zee Eg)	$\int_{(CH^{-1})}^{A(v)dv}$	A(2240)	A(2213)	Resm. r.s
94	6	0.00016	4.5		^ ^ -	0.4	Semple 94: 4.5
25	6	0.00016	3035	0.27 0.35	0.5 0.6	0.5	mm Hg of 1/50 HgO mixture. Hample 95: Hg added to sam- ple 94.
96	7	0.00017	<b>48.</b> 0	2.18	3.4	3.2	Samples 96-
97	7	0.0033	90.0	3.93	6.4	5.7	101: 1/50 <b>%</b> 20
98	7	0.0060	165	6.96	11.2	9.9	mixture alone.
99	7	ს.0098	270	11.33	18.5	16.1	
100	7	0.0136	<b>3</b> 75	15.4	25.0	22.0	
101		0.0223	618	23.7	37.5	33.5	
102	5	0.0016	9.1	1.50	2.1	1.8	Samples 102-
103	555555555	0.0024	13.3	1.87	2.7	2.5	112: 1/10 H <sub>2</sub> 0
104	5	0.0032	17.8	2.45	3. <b>8</b>	3.3	mixture alone.
105	5	0.0045	25.8	3.5k	5.1	b., 5	
106	5	0.0065	36.4	4.82	7.3	6.6	
107	5	0.0091	50.6	6.46	10.0	9.0	
108	5	C.0185	103	12.76	19.1	17.5	
109	5	0.0295	165	19.50	26.9	26.3	
110	5	0,0500	581	30.70	45.4	41.6	
111	5	0.0750	418	41.90	61.6	56.5	
112	5	0.1150	<b>Cyty</b>	57.00	78.8	75.0	
			Path L	ength = 6	.35 ca		
113	10	0.0371	5.6	5.4	8.5	8.35	Sample 113: 5.0
114	10	0.0371	8.1	7.0	9.3	8.97	Es of Hoo.
115	10	0.0371	11.8	7.9	10.6	10.1	Samples 114-
116	10	0.0371	20.5	9.6	13.4	12.55	125: Ep added
117	10	0.0571	52.0	11.4	16.2	15.5	to seamle 113.
118	10	47,000	49.8	13.7	19.6	18.0	
119	10	0.0371	98.1	17.6	26.2	23.8	
120	10	0.0371	185	24.2	34.2	5 <u>1</u> .1	
151		0.0371	508	26.2	41.3	35.0	
155		0.0333	\$ 5	6.3	8.3	8.23	
125		0.0571	49.4	13.9	19.5	17.9	
124	10	0.0371	414	27.9	45.4	27.6	
125	10	0.0371	735	32.3	51.0	45.0	
195		0.0571	1550	35.9	58.0	52.4	
127		0.0771	2300	57.9	59.5	54.5	
108	70	0.0571	3070	57.9	59.5	54.9	

Table 5. (Continued)

ample	Fig.	7	P <sub>r</sub>	/A(v)dv	A(2240)	A(2213)	
ilo.	No.	(atmos cm)	(ma Hg)	(cm-1)	5	<u> </u>	Romarka
129	10	0.074	11.2	11.5	15.7	15.3	Sample 129:
130	10	0.074	21.2	14.5	20.0	18.3	10.0 <b>=</b> itg of
131	10	0.074	51.2	19.8	28.7	25.7	E_C. Sample
132	10	0.074	101	26.0	39.0	34.2	130-135: 15
133	10	0.074	200	33.2	49.7	43.5	added to same
134	10	0.074	400	12.7	65.5	55.5	ple 129.
135	10	0.074	740	19.	75.2	66.5	•
136		0.074	11.2	ц.7	15.9	15.5	Sample 156:
137		0.074	101	25.8	36.7	34.0	10.0 mm Hg of
138		0.074	741	48.8	74.5	65.7	NoO. Sample
139		0.074	1485	54.0	83.8	76.0	157-141: 16
140		0.074	2295		<b>8</b> 6.0	80.5	
141	10			55.5		90.5	added to same
141	10	0.074	3075	56.3	85.7	81.2	ple 136
142	10	0.148	22.4	22.4	29.6	28.1	Sample 142:
143	10	0.148	33.2	24.5	3Å.O	32.0	20.0 mm Mg of
144	10	0.148	52.1	26,8	39.7	37.2	Mot. Samles
145	10	0.148	105	57.2	53.3	48.0	143-151: M
146	10	0.148	199	45.5	65.9	59.3	added to san-
147	10	5,148	395	56.2	79.1	72.0	ple 142.
148	10	0.148	718	62.0	90.1	83.0	<b>Jan 1 10</b>
149	10	0.148	1655	68.5	96.1	92.5	
150		0.146	2315	70.5			
	10			70.5	95.5	94.0	
151	10	0.148	3120	71.0	96.5	95.0	
152	9	0.37	<b>56</b>	46.9	60.1	59.3	Septe 152:
155	9	0.37	1.22	97.6	74.7	72.2	and light of Red.
. < 4		0.37	225	67.2	85.1	82.7	Samples 155-
157	9	0.57	388	75.2	93.8	91.9	160: 15 added
156	á	0. 1	559	76.1	97.2	96.7	to semple 15
157	ý.	ं <del>शं</del>	789	ô0.2	97.4	97.	
158	,	0.3	1210	82.8	97.9	98.1	
159		0.37	1985	و. و و . رو	90.U	90.5	
160 160		0.37	207	84.5	96.2	98.9	
700	9	0.31	4477	04.7	œ.∠	yo.y	
161	9	0.74	112	70.	86.3	85.0	Sample 161:
162	9	0.7	131	77.6	95.1	91.9	100 am Mg of
363	9	C.7	5,65	97.7	<b>36.6</b>	76.5	MgO. Samples
164		0.74	479	₩Ċ.8	98.4	<b>98.</b> •	168-167: No
165	9	0.74	767	90.5	22.0	29.4	edded to com-
166	-	0.74	1530	52.2	29.5	99.6	ple 161.
167	9	0.74	3000	93.5	çş.3	99.8	

Table 3. (Continue)

Sample No.	Fig.	(atmos cm)	P. (mm. Hg)	(cm-1)	A(2240)	A(2213)	Remarks:
					98.7	00.1	Sample 168:
168	9	1 1	227	91.		99.1	
169	_	1.51	365	91.8 97.8	99.5	99.8	201 mm Eg of NoO. Samples
170	9	1.51	769	98.8	99.9 100	99.9 100	169-173: Ka
171		1.51	1545	98.9	100	100	added to sem-
172		1.51 1.51	2360	100	100	100	ple 168
173		1.51	3105	100	100	100	bre 100
174	9	3.02	456	102.1	100	100	Sample 174:
175	•	3.02	611	103.3	100	100	405 mm Hg of
176		3,02	877	106.5	100	100	NoO, Saugles
177		5.02	1570	105.6	100	100	175-178: No
178	9	3.02	2365	105.2	100	100	added to sam-
-,-		•					ple 174.
179	6	0.007*	1.1	2.0	2.5	2.6	Sagles 179-
180	6	0.0111	ì	و.۶	2.8	3.0	106: %0 alone.
181	6	0.0148	2.2	2.8	3.9	3.8	• • • • • • • • • • • • • • • • • • • •
182	6	0.0192	2.5	3.5	4.5	j. 🖫	
18:	á	0.0225	5, 4	3.8	5.1	4.7	
184	6	0.0270	4.2	ă.9	6.6	6.5	
185	6	0.0742	11.2	11.7	16.0	15.8	
186	6	0.172	17.0	17.5	22.6	32 1	
			Pach L	ength = bo	0 ca		
197	11	1.4	3.≥	32.7	36.4	<b>38.</b> 4	Pample 187: 2.9
186	11	1.4	6.2	38.7	45.2	46.8	m Mg of MgO.
189	n	1.4	10.8	46.4	54.9	<b>%</b> .≎	Samples 188-
190	u	1.4	50.0	60.7	74.5	74.5	194: No added
191	11	3.4	74.0	<b>80</b> .0	91.5	91.5	to sample 167.
192	17	1.4	575	œ.ä	<b>:8</b> .4	98.€	
193	41	1.6	736	98.8	90.5	100	
104	11	1.4	3035	101.2	700	100	
495	11	2.8	4.1	<b>77.</b>	60.4	62.2	desple 195:
196	$\mathbf{n}$	2.8	13.1	64.1	73.9	75.0	5.8 mm Mg of
131	17	2.8	29.2	10.3	11.4	01.1	igu. d <b>em</b> les
700	11	2.8	73.7	22.2	27.3	.77.€	196 201. 🔩
199	77	2.5	220	101.4	29.8	100	added to see-
3/0	11	2.8	745	106.3	100	700	ple 195.
507		2.8	3050	108.2	100	100	

Table 3. (Continued)

Sample.	Fig.		P	JA(v)dv	A(2240)	4/0037	
No.	No.	(atamos cm.)	(mma hg)	(cm-1)	A(2240)	A(2213)	Remarks
202	13.	5.8	13.9	79.9	87 -	90.0	Sample 202:
203	11	5.8	26.7	91.6	96.⊥	97.0	12.4 mm Hg of
204	ü	5.8	64.5	100.0	99.0	99.6	NoO. Samples
205	11	5.8	197	103.5	99.6	700	203-206: No
206	n	5.8	750	113.4	100	100	added to sam- ple 202.
207	11	11.6	28.0	102.0	99.0	99.5	Sample 207:
208	11	11.6	61	108.8	100	100	25 mm Hg of
209	n	11.6	203	115.7	100	100	N20. Samples 205-209: N2 added to sem- ple 207.
<b>21</b> 0	10	23.7	56	115.6	100	100	Sample 210:
517	10	23.7	103	120.1	100	100	50.3 mm Hg of
\$13	10	23.7	21)	125.3	100	100	N <sub>2</sub> 0. Samples 211-212: N <sub>2</sub> added to sam- ple 210.
215	10	1.39	5.4	32.7	37.7	40.0	Samples 213-
574	10	عز .2	5.5	45.8	52.7	54.5	216: <b>5</b> 20
216 216	10 10	3.70 5.56	9.0 13.4	62.5 78.5	72.2 85.4	73.4 86.4	alone.
			Path Le	ereth = 80	<u>ده</u>		
il.		18.8	22.4	103.2	100	100	Sample 217:
578		18.8	12.1	108.7	100	100	20.0 = Eg of
219		18. <b>8</b>	105	115.2	100	100	Noo. Samples
220 221		18.8 .8.3	476	120.2	ĵúć	100	218-221: N
		70.5	742	121.8	100	100	ple 217.
			Path Len	veth = 1600	) ca		
كخذ	•	16.4	44.6	125	100	100	Remple 222:
٠.٠	=	76.	168	129	100	170	MO.0 mm lig ut
le <sup>i,</sup>	4	16.4	SII	LEY)	100	100	NoO. Samples 225-224: No suited to sam- ple 222.

Table 4. Data for the 2145 cm<sup>-1</sup> CO Band

The Fig. No. corresponds to the figure in which the spectrum of each sample is shown. Absence of Fig. No. indicates spectrum is not shown. Uncertainties of values of  $A(\nu)$  and  $JA(\nu)d\nu$  are discussed in the text.

Sample	Fig.	v	Pe	JA(V)av	A(2168)	A(2116)	<b>D</b>
No.	No.	(atmos cm)	(mm Hg)	(cm-1)	<u> </u>	- \$	Remarks
			Path Le	ngth = 1.5	5 cm		
					<del></del>		
1	23	0.00096	5.4	0.14	0.18	0.20	Sample 1: 5.4
2		0.00096	107	0.24	0.25	0.25	mm Hg of mix-
3		0.00096	670	0.26	0.30	0.30	
14	23	0.00096	3020	0.27	0.32	0.32	CO and 9/10 N <sub>2</sub> Samples 2-4: N <sub>2</sub> added to sample 1.
5	23	0.00154	49.5	0.31	0.36	0.37	Samples 5-7:
5 6	-/	0.00154	313	0.39	0.43	0.40	
7	23	0.00154	3040	0.39	0.43	0.40	
•	-,	,	,	****			CO and 9/10 No added.
8	23	0.00262	15	0.36	0.35	0.31	Sample 8: 15
9	-	0.00262	40.7	0.45	0.47	0.11	men Hg of mix-
10	23	0.00262	106	0.56	0.65	0.56	
$\mathbf{n}$		0.00262	256	0.64	0.72	0.65	
12		0.00262	738	0.68	0.74	0.69	Camples 9-13:
13		0.00262	3020	0.69	0.76	0.70	No added to
							sample 8.
14	23	0.00435	25	0.53	0.50	0.50	Sample 14: 25
15	-	0.00435	68	0.68	0.70	0.68	Hg of mixture
16	23	0.00435	245	0.91	1.0	0.85	or 1/10 CO and
17		0.00435	619	0.99	1.1	0.95	9/10 No. Sam-
18		0.00435	1315	1.05	1.2	1.0	ples 15-19: No
19	දිර	0.00435	<b>₹020</b>	1.09	1.5	1.1	added to sampl
20	25	0.00703	40	0.84	0.85	0.7	Sample 20: 40
57	-/	0.00707	93	1.10	1.1	1.0	ma Hg of mix-
22	23	0.00705	220	1.40	1.5	1.2	ture or 1/10 0
25	-,	0.00703	410	1.60	1.7	1.4	and 9/10 No.
24		0.00703	1520	1.87	2.2	1.7	Samples 21-25:
25	23	0.00703	3020	1.91	2.2	1.7	No edded to
-,			,,,,,	/-		** 1	sample 20.

Table 4. (Continued)

emple	Fig.	/ ¥ \	, Fe	JA(v)dv	Y(5128)	A(2116)	Remarks
No.	No.	(atmos cm)	(mm Hg)	(em-1)			
26	25	0.0107	6.1	0.48	0.40	0.38	Sample 26: 6,0
27	23	0.0107	25.0	0.84	0.85	0.70	mm Hg of CO
28	-	0.0107	52.0	1.11	1.10	0.88	alone. Sumple
29	23	0.0107	118	1.51	1.70	1.40	27-32: No adde
30	23	0.0107	314	2.14	2.45	1.90	to sample 26.
31	23	0.01.7	720	2.50	2.95	2.35	•
32	23	0.0107	3020	2.83	3.40	2.70	
33	23	0.0176	10.6	0.81	0.73	0.72	Sample 33: 10.
34	23	0.0176	22.2	1.12	1.1	1.0	mm Hg of CO
35	23	0.0176	49.9	1.65	1.5	1.3	alone. Sample
36	23	0.0176	157	2.55	2.6	2.2	34-39: No adde
37	23	0.0176	412	3.48	3.7	3.2	to sample 33.
38	23	0.0176	1005	4.14	4.7	3.7	-
39	23	0.0176	3020	4.67	5.6	4.3	
40	23	0.0362	20.9	1.57	1.4	1.2	Sample 40: 20.
41	23	0.0362	45.6	2.32	2.2	1.9	man Hg of CO
145	23	0.0362	131	3.79	3.5	3.1	alone. Sample
43	23	0.0362	254	5.10	4.9	4.2	41-46: No adde
71,71	23	0. 2362	692	7.09	7.4	6.1	to sample 40.
45	23	0.0362	1520	8.43	9.3	7.4	
46	23	0.0362	3020	9.33	10.3	8.1	
47	23	0.0722	41.8	3.39	3.2	2.7	Sample 47: 41.
)+8	23	0.0722	85	4.70	4.3	3.7	mana Hig of CO
49	23	0.0722	189	6.75	6.2	5.3	alone. Sample
50	23	0. <i>0</i> 722	371	9.05	8.7	7.3	46-53: No adde
51	23	0.0722	788	11.97	11.9	9.9	to sample 47.
52	23	0.0722	1520	14.65	15.4	12.5	
53	23	0.0722	3020	16.76	18.4	14.6	
54	29	0.165	95.8	8.3	7.3	6.4	Samples 54-59:
55	29	0.349	203	17.6	15.2	15.8	CO alone.
56	29	0.709	412	34.4	29.3	26.1	
57	29	1.31	761	60.7	51.7	46.1	
58	29	2.67	1555	108	84.5	78.7	
59	29	5. <b>52</b>	3210	157	100	99.7	

Table 4. (Continued)

Sample	Fig.	<del></del>	Pe	JA(v)dv	A(2168)	A(2116)	
No.	No.	(atmos cm)	(mm Ha)	(cm-1)		*	Remurks
			Path L	ength = $6$ .	35 cm		
60	24	0.073	10.2	1.8	1.4	1.4	Sample 60: 10
61		0.073	16.0	2.5	1.8	1.7	mm Hg of CO
62	24	0.073	27.0	2.9	2.2	2,1	alone, Samples
63		0.073	45.0	3.7	3.1	2.7	61-70: No added
64	24	0.073	86.0	4.8	4.1	3.5	w sample 60.
65		0.073	170	6.2	5.7	ű,á	
66	24	C.073	292	8.0	7.5	6.3	
6?		0.073	492	10.0	9.5	8.1	
<i>6</i> 8	24	0.073	735	11.7	11.5	9.2	
69	21:	0.073	1515	11.5	15.0	12.2	
70	24	0.073	2970	16.3	18.1	14.1	
10	<b>47</b>	0.015	2910	10.7	10.1	14.1.	
71	24	0.145	20.4	3.8	3.1	8.8	Sample 71: 20
72		0.145	32.4	4.8	4.O	3.7	man Hg of CO
73	24	0.145	49.8	5.6	4.7	4.2	alone. Samples
74	24	0.145	88.5	7.2	6.1	5.4	72-80: No added
75	24	0.145	152	9.1	8.0	6.7	to sample 71.
76		0.145	202	12.0	10.8	9.1	
77	5#	0.145	545	15.9	14.6	12.3	
78		0.145	797	18.7	17.5	14.9	
79		0.145	1545	23.6	23.2	19.3	
8ó	25	0.145	3070	28.7	29.5	24.0	
81	24	0.298	42.0	7.6	6.3	5.8	Sample 81: hl
82	24	0.298	64.0	9.2	7.7	7.0	mm Hg of CO
83	24	0.298	95.0	11.0	9.4	8.3	alone. Samples
84	24	0.298	1.8	13.4	11.4	10.4	82-91: No added t
85	57	0.298	231	16.3	14.3	12.3	cample 81.
86	24	0.298	333	19.5	17.0	14.9	compre or.
87	24	0.298	450	22.2	16.0	14.9	
88	25	0.298	747	27.5	19.8 24.2	17.1 20.5	
89	25	0.298	1470	35.8	32.8		
90	25	0.298	2260	42.0		27.7	
90 91	25	0.298	2960	46.3	40.0	33.0	
91	<b>4</b> 7	0.290	2900	40.7	44.2	36.2	
92	24	0.440	59.0	10.9	8.6	8.2	Sample 92: 58
93	24	0.440	84.5	12.5	10.7	9.6	mm Hg of CO
94	24	0.440	153	16.5	14.0	12.7	alone. Sumples
95	24	ō, 44ō	الماج	دو ، وابغ	17.9	15.7	95-102: 15
95 96	86	0.440	341	24.5	20.8	18.3	added to sample
97 <b>98</b>	26	0.440	4 <b>6</b> 6	28,9	24.0	20.8	92.
98	26	0.440	727	34.0	29.0	24.7	-
99	-	0.440	1240	42.4	37.3	31.7	
100	26	0.440	1660	48.7	42.8	36.2	
101		0.440	2160	52.9	47.6	40.3	
102	25	0.440	2955	58.8	53.3	45.3	
F/E	40	01 110	-277	,0.0	22.3	マフ・ブ	

Table 4. (Continued)

Sample	Fig.	(	P.,	JA(v)dv	V(5198)		Remark
No.	No.	(atmos cm)	(mar Hg)	(ca-1)	<u>*</u>	- \$	
103	26	0.800	112	19.9	16.2	14.3	Sample 103: 110
104	26	0.800	169	24.0	19.8	18.0	mm Hg of CO
105	26	0.800	247	27.9	23.2	20.8	alone. Samples
106	26	0.800	<b>38</b> 8	35.1	29.5	26.2	104-112: No
107	25	0.800	513	39.5	33.0	29.3	added to sample
1.08	26	0.800	74É	47.1	<b>i</b> .5.0	34.8	103.
109	25	0.800	1260	56.1	49.6	43.0	2071
110	ēč	0.800	1615	6.6	55.6	48.2	
111	26	0.500	2215	72.9	62.0	54.4	
112	26	0.800	2940	81.3	63.5	60.5	
113	26	1.48	207	37.5	30.0	27.5	Sample 113: 203
114	26	1.48	30¥	45.2	36.0	32.2	men He of CO
115	26	1.48	461	54.5	43.2	39.0	alone. Samples
116	26	1.48	752	36.5	52.8	47.2	114-120: No
117		1.48	1240	81.5	64.0	57.0	added to sample
118	26	1.48	1960	95.0	74.3	66.2	113.
119		1.48	2310	100	78.5	70.5	
120	26	1.48	3060	110	82.0	76.8	
121	25	2.74	412	70.2	56.0	51.0	Sample 121: 404
122	25	2.0	591	80.2	53.5	58.5	mana E <sub>c</sub> of CO
123	25	ક.94	853	93.1	76.0	66.0	alone. Samples
124	25	2.94	1215	106	80.3	74.5	122-127: No
125	25	2.94	1770	119	8, 8	81.5	added to sample
126		2.94	2295	127	92.8	88.8	121.
737	25	2.94	2 <u>350</u>	135	26.8	91.2	
128	25	5.55	711	112	84.1	79.2	Sample 128: 762
129	25	5.55	1180	129	92.8	89.2	Be of CO
130		5.55	1465	138	96.0	21.5	Alone, Sumples
131	25	5.55	1815	145	95.2	95.0	129-155: 12
132		5.55	2295	152	98,8	97.0	added to sample
133	25	3.55	2955	158	99.2	98.4	126.
134	25	10.9	1530	153	99.7	98.8	Sample 154:
175		10.9	1840	169	100	99.3	1500 mm Hg of
136		10.9	2295	172	99.7	99.5	CO elone. Sen-
157	25	10.9	<b>30</b> 85	1.19	100	99.5	ples 135-137: My added to emmple 134.

Table 4. (Continued)

Mo.	Fig.	(atmos cm)	Pe (ma Hg)	JA(v)dv (c≡ <sup>-1</sup> )	A(2),68)	A(2116)	Reserve
138 139	25 25	16.6 16.6	2325 3140	182 187	99.€ 99.7	99.6 39.7	Semple 138: 2280 mm Hg of
							ple 159: No added to sample 135:
140	કંડ	22.2	3115	194	99.6	99.6	Sample 140: 30% mm Hg of CO alone.
			Path I	ength = 40	<u>0 ca</u>		
141	27	1.32	3.0	6.0	4.4	4.5	Semple 141: 2.
142	27	1.72	5.3	6.9	5.2	4.9	<b>■ 3</b> 2 07 00
145 144	27	1.32	10.0 26.2	9.4 14.2	5. <b>8</b>	6.6	alone. Sample
145	27 27	1.)2 1.32	61.1	19.7	10.9 15.6	10.0 14.7	142-149: No
146	\$1 2	1.32	170	33.6	26.1	24.2	141.
147	27	1.32	503	51.8	42.2	36.3	~~~
146	27	1.92	1520	<b>8</b> 1.9	66.8	60.0	
149	27	1.32	2770	102.6	81.8	74.0	
150	98	2.62	5.8	16.0	6.8	6.6	Seastle 150: 5.
151	56	2.62	10.	13.6	9.0	8.5	make of co
152	26	£. <b>52</b>	29.2	20.5	15.0	14.9	alone. Sample
153	28	8.62	59.1	æy.0	21.1	19.4	151-157: 🌄
154 155	26 26	2.62 2.62	186 503	46.7	36.1 56.7	55.5 52.2	added to seep!
156	26	2.62	1520	73.2 109	32.0	75.6	130.
157	26	8.62	89/0	151.5	95.0	90,0	
358	26	4.82	10.9	17.8	13.6	12.9	Sample 158: 10
150	26	<b>4.6</b> 2	25.6	26.0	20.2	19.0	Re of CO
160	28	<b>⇒.</b> 92	65.5	78.6	30.8	28.6	wione. Sample
161	26	4.82	174	60.2	17.5	14.2	159-164: Ro
168	26	4.62	510	?3.ó	71.7	66.8	added to rempl
163	26 26	4.62 4.52	1520	191.5	94 4	ຊກ.ກ ສະ	140.
704	20	4.02	<b>297</b> 0	151.0	5.6	27.3	

Table .. (Continued)

Rocal ke	¥(5779)	म्(श्रेळ) •	$\int_{\infty}^{\Lambda(\nu)d\nu}$	a) (m lg)	(atmos. c	Fig.	No.
Sample 165: 32	35.4	39.4	50.7	<b>52.</b> 5	14.4	26	165
max Reg of CO	53.3	56.9	73.8	93.6	14.4	28	166
alone. Samlo:	75.7	Al 2	110.6	237	14.4	28	167
166-170: 10	96.6	98.1	150.9	756	14.4	26	168
added to sea-	99.9	100	170.1	1520	14.k	26	169
ple 165.	100	100	183.4	2970	14.4	26	170
Sample 171 · 100	82,3	89.0	123.3	101	45.6	28	171
En of CO	97.4	98.8	156.8	257	45.6	26	172
alone. Samples	100	100	184.6	771	45.6	26	173
172-175: %				•••			
added to sample							
171.							
Samples 174-179	1.8	1.8	2.47	1.0	0.46	27	174
CO alon:	2.9	5.1	4.35	2.0	0.91	27	175
	<b>6.7</b>	<b>⊹.8</b>	5.76	3.1	1.37	27	176
	6.2	6.7	8.70	5.1	2.28	27	177
	23.5	25.6	31.90	20.4	9.10	81	178
	53.2	57.6	73.90	51.0	22.8	27	179

Table 5. Data for the "760 cm-1 CO Bend

The Fig. No. corresponds to the figure in which the spectrum of each sample is shown. Absence or Fig. No. indicates spectrum is not shown. Uncertainties of the values of  $JA(\nu)d\nu$  are discussed in the text.

mple No.	Fig.	(atmos. cm)	Pe (■ Bar)	Path (cm)	JA(v)dv (cm-1)	Bene rhe
<i>7</i> 0.		(2000) (8)	\	1687	· · · · · · · · · · · · · · · · · · ·	
1	40	36.6	54	625	8.2	49.9 m lg of
î	39	72.0	5 <b>h</b>	1232	10.3	coin cell alone.
3	\$ 50	143	54	2445	15.5	
4	39	265	54	4875	22.1	
5	40	36.6	217	625	16.2	H2 added to above
6	22	72.0	217	1232	22.2	sample to total
7	#C	143	217	2445	30.1	pressure of 213
8	959	285	217	1075	43.0	mm Rg.
9	40	36.6	<b>hoi</b>	625	21.0	No added to above
10	39	73.0	404	1232	zô.2	sample to total
11	FU	143	# C#	2445	39.7	presente of 400
12	39	265	POP	1075	56.7	m I4.
13	40	36.6	744	625	25.8	Il, added to above
14	30	72.0	744	1272	<b>30.</b> 0	Labor of edgeste
15	40	343	744	2445	52.6	pressure of 143
36	39	285	764	1875	72.2	m 4.
17		<b>20</b> 5	766	¥375	71.0	Sem. 17 is repus of sem. 16.
19		74.4	106	Carr	٠.0	100 m & of CO
19		103	106	12.12	29.2	'a cell alone.
ŁŪ		261	.08	2445	30.6	
Pì.	<b>PC</b>	wi	108	MIT	12.9	
22		72.4	poß.	625	19.9	Es added to above
3		143	200	1232	29.5	sample to total
50		26)	206	2445	¥Q. 8	pressure of 200
53	60	564	\$03	<b>M7</b> 5	<b>%.6</b>	m Se.
26		72.4	100	625	26.0	By added to show
<b>2</b> 7		145	•08	ענע	٠٥, <	ecopie to ecit.
20		<b>৮</b> শী ৭	<b>206</b>	3664	خلا ۽	pressure of 100
<del>&gt;</del> 9	40	364	£00	1375	4 4 4 4 4	- *:
30		72.4	749	623	36.9	Eg added to above
31		263	765	27.2	52.5	ecopie te total
<b>3.</b>		கூர் 3	768	2645	79.2	pressure of 740
11	60	54m	760	M075	94.U	m Ke.

140

Table 5. (Continued)

Sample No.	Fig.	(atmos. cm)	(mm Hg)	Patin (cm)	JA(v)dv (cm-1)	Romarks
3h 35 36 37	39	լև6 9 <b>88</b> 5 <b>7</b> 2	216 216 216	625 1232 2445	30.3 h3.h 58.5	200 mm Hg of CO in cell alone.
37	39	1140	216	4875	81.0	
<b>3</b> 8	39	146	420	625	41.7	No added to above
39 40		288 572	1150 1150	1232 2445	58.0 77.5	pressure of high
41	39	1140	420	4015	105	m Hg.
42	39	146	756	625	53.5	No added to
43 44		288 572	756 766	1232 2445	70.6 <b>94.</b> 6	above sample to
45	39	1140	756 756	4875	125	total pressure of 740 mm Eg.

Tab.s 6. Date for the Off, Bods

The Fig. 60. corresponds to the figure in which the spectrum of each sample is shown. Absence of Mg. No. indicates spectrum is not shown. Underwistes of the walks of  $\int A(y) dy$  are disquested in the tast. Attended of a talking walks of  $\int A(y) dx$  for the 1550 on  $^2$  band indicates it was too what, be seen to

	ě	7	7	1550 1506	150b am-16ane	J.(v).	
*(**. **)	. <u>.</u>		(	(mm Re)	<b>71⊈.</b> ₩0.	[cm-1 1405 cm-1 bard	Remaria
			Ž.	Ne Lagta - (.3) 94	.15 ex		
51771	3.6			8.8	1.1	1.15	HAD LESS OF STREET
ş	~		<b>5</b>	5.8		91.1	alous sam. 2-f. H2
613	10.7		3.10	10.8		66.1	
6.63	Š		512	30.6		. f.	
0.435	9.		2.tb	93.8		Q.	
(1)	<u> </u>	<u></u>	3.80	Ŕ		2,30	
0.03,	2. ž	\$	3.70 4.10	103 203 203 203		2,50 2,78	
a. 636	*	\$	\$	<b>8</b> .	1,1	1.70	Maa. 9, 3.5 mm Hg Off., 6. com j
3	8,1		3.00	9.1		9	
3	4.3		ž	19.6		8	to San. 9.
00 (B) (C) (C) (C) (C) (C) (C) (C) (C) (C) (C	£3	<b>\$</b>	5.3	\$. 78 \$. 78 \$. 20		3,00	
2 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5	7. A.		4.5	137		iv.a.	

Sable 6. (Continued)

		28%	SCO CET bead	3	1550 £ 1306 on "band	Can Joans	fa(v)av	
r i	No. (atmes. cm) (um f.g.)	(m f.c.)	i L	J.(-) (-=-)	P. (m. Rg)	F1g. Bo.	(cm <sup>-1</sup> ) 1306 cm <sup>-1</sup> band	Remarks
32	0.086 9.086	1005 3073	\$3	6.9 7.1	1005 3075		4.6 6.4	
2	<b>8</b> (0.0	9.1	\$	1.4	7.6	L.	2.7	Вел. 18, 7.0 ши Ис СИц. «Іспе)
28	त्र त्र १८००	4.5.8 9.7.8	\$4	<b>9</b> 10	24.0 25.5	14	% % &&	Ban. 19-26, We added to Ban. 18.
ឌដ	85.9 88.9 88.0	3.8 0.	\$	6.8 7.8	% % 0.0%		4.7 9.4	
2/8	9.; 88	88	\$	2.6	<u>%</u> 8	4	1.0 7.9	
22.8	85.0 0.0 0.0	1515 3075	\$	12.4	1,500 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000 3,000		7.8	
72	0.073	7.52	\$	5.9	13.5	L <sub>A</sub>	3.8	Sam. 27, 9.8 mm Hg CEL, alone;
82	0.073	23.2	Ş	6.8	24.0	L a	5.9	Sam. 28-35, H <sub>2</sub> added to Sam. 27.
22	0.073 0.073	શ્લુ	:?	9.5	85.7 181	Ŀ,	7.1 8.5	
××	0.073	3178	\$	13.0	<b>3</b> 5	Lte	9.6 10.5	

Table 6. (Continued)

		ě	1.83	And	1550 £ 1306	ca_band	CA(v)dv	
F. 16	'ataca. ca) (su Ng; No. (ca	Pe Kg;	2 2	(cg-1)	Fe Fig. (== Hg.) Ho.	F16.	1306 cm <sup>-1</sup> band	Remerks
メン	0.075	15 yo	*	15.6 16.8	1530 3CT5	L. 9.	11.5	
9,	O.1%	**	#	9.3	24.9	F.4	6.9	Sam. 36, 18 um Eg CEh, alone;
<b>≈%</b>	0.1% 0.1%	27.08 5.0	3	20.5	39.2 57.0	14	9.5	Sam. 57-43, Badded to Sam. 36.
22	0.13 0.13 4.1.0	239	#	33.6	211	14	10.5	
7.3	3 X C C C C C C C C C C C C C C C C C C	50 K	4	1.61 9.15	4 6 7 7 5 8 5 1 5 1 5 1 5 1 5 1 5 1 5 1 5 1 5 1	Lq	15.6 16.6	
\$	0.1%	15:5	3	24.8	1530		0.81	
1	0.223	2	3	9.41	<b>41.</b> 4		10.4	Sam. 44, 30 mm Hg
₹.	0.233	Ē		16.0	<b>₹</b> .83	Į.	9.11	CM, Bloms; Sem. 45-51, M2 added to Fem. 44.
35	9.5 2.23		33	18.5 22.	0110 225	Lq	4.81 18.4	
22	0.23 0.23	7 <u>7</u>	3	%% 5.50	89 <u>7</u>	L#	25.1	
<b>9</b> ;4	7.225 U. 223	15.55 25.55 25.55		, , ,	1530 3630	14	3 <b>6.</b> 0 27.1	

Stalle 6. (Contaried)

		200	220 cm 1	A)	1550 £ 1306 cm-1band	ca-1band	JA(v)&v	
ė	(Atmos. :a) (ms Eg)	(File 18)	Žú	(12-1)	(m Eg)	716. 80.	(cm <sup>-1</sup> ) 1306 cm <sup>-1</sup> band	Remarks
×	0.431	75.5		22.9	<b>8</b>	1.4	16.0	Sam. 52, 58 mm Hg CH <sub>h</sub> alone;
24	0.431	.18 21 22	11	25.5	128 24.3		19.6	Sam. 53-58, No added to Sam. 52.
28	0.431	2.85 2.85	3	13.4	şÊ	LA	28.0 32.8	
23	25.451	2.5 2.8 2.8	3	\$6.0 \$4.2	11 8 98 98	14	36.54 80.64	
\$	9.728	ä	3	**				Ban. 59, 98 mm Hg CHia.
\$6	8.7.3 88	38	33	36.5				Sem. 60-52, M2 Edded to Sem. 59.
3	6.73	8 8 8	3	73.0				
				ā	Att Length - 400 on	8		
ij.	C.7	0	<b>3</b>	18.9	2.1		1.7	Seut. 63, 1.52 mm fig CHi, elone;
3	C. 7	3.3	\$	29.8	3.4		8.0	Sea. 64-73, W2 added to Sea. 63.
23	00.1	5.7	33	23.2	5.8	<b>3</b> 3 33	9.0 11.5	

Table 6. (Continued)

1		, x	1.50 ca.	10(1)	1550 4 1506 cm band	A. Land	J.k.(v)dv (cm-1)	Remarks
ف	(stros. es,	(¥	á	(0.1)	( See)	No.	1506 cm-1 band	!
53		21.7	3	% 8.3 8.5	21.5	3.3	16.4	
38	LL Uii	21		<b>%</b>	27.5	33	¥.33	
22	;; ;;	120		 	745	33	4.6.3 4.9.4	
22	п.э	ĝ		76.0	0204	<b>9</b>	4.15	
2	1:35	5.7	ï	28.4	3.8	3	10.9	Sem. 74, 2.75 = Eg
t	1: 35	8.8	3	\$	6.1	63	22.8	CH, alone; Sam. 75-83, H2 added Sam. 75.
<b>%</b> E	ጓጘ ፚ፞ፚ፞	23.5	<b>₹</b>	XX XX	10.4	33	14.5 20.2	<u>:</u>
23	:: 20	25. 25. 25.	2	\$.5 \$.5	53.6	33	:: **	
84	44	<u>8</u> 2	33°	36.0	***	33	45.5 80.5 80.5	
85	27 11.	3, §	22	8,8j	085 040 040	22	4.88.17	

Table 6. (Continued)

		ži	5	Pard	1550 2 1506 as 1 band	Pand 1 ec	JA(v)dv	
o c	(atan. ca)	() ()	73 g.	(1. ()	( ) ( ) ( ) ( ) ( ) ( ) ( ) ( ) ( ) ( )	71g.	(cm <sup>2</sup> .)	Remarks
£		6.9	*	4.54	7.5	5		Sean. (14, 5.7 mm Hg.
£	**	10.A	<b>;</b>	45.2	11.2	\$	21.5	CH <sub>L</sub> alone; Sam. 85-94, M <sub>2</sub> added to Sam. 84.
25	34 G 6 M	16.5	3,	47.6	16.9	33 <b>3</b> 3	23.5 28.5	
88	5 Q 4 3	\$5.0 \$3	*	65.5 75.2	55.5 130	33	35.0 37.9	
88	OR SER	8	Š	8.0	<u>8</u> 8	22	39.9 61.7	
82	3 X 3 3	<b>8</b> %	ž,	8 8 A	95 28.	33	76.0 90.0	
3	Ø	<b>9</b> 6	¥	153	30%	2	102.0	
				ě	Ath Length - 1600 or	2009		
2	5.5	9.9	\$	39.0	4.1	64	28.0	Sam. 95, 3 mm Hg CH,
×	<u>.</u>	9.6		65.5	æ.	6	59.0	alore; Ban, 96-134, N; alded to Hen. 95.
<b>58</b>	10 to	26.9	<b>TT</b>	71.0 71.0	15.1	66	35.0	
88	20 <b>20</b> 20 20	21.5 286	2	67.8 106	51.¢	664	51.7 67.3	

Table 6. (Continued)

		8	020 cm 1 1434	1624	1550 ½ 1306 cm <sup>-1</sup> bend	ca band	JA(v)dv	\$4(v)A	
	(atmos. cm) (mm Bz)	E ( )	7. 5.	(A(v)4y (Cal-1)	P. (m. 18g)	Mg.	(cm <sup>-1</sup> ) 1306 cm <sup>-1</sup> band	(cm <sup>-1</sup> ) 1550 cm <sup>-1</sup> bend	Remarks
	301 141		54	प्रदे इ	74:	<b>\$</b> \$	85.9 107		
	1580 2865	0.15	*	193 420	1520	<b>₹</b>	126 141		
	7	7.8	<b>*</b> 2	82	8.3	6 <del>1</del>	38.4	1.9	Sam.105, Sam Hg CH <sub>t</sub> alone;
	ង	เล	\$	103	9.11	<b>6</b> 4	45.0		Bam. 105-215, Mo madded to
	%	26.2	. · ·	111	26.7	Ç.	55.9		<b>Sen</b> . 105.
	51.8	<b>6</b>	45 45	%19 %19	52.3 121	\$ <b>\$</b>	4- <b>69</b> 4-606	3.3 4.1	
1.1. 1.1.	32		24	197 230	82	<b>Q</b> , <u>0</u> ,	115	7.0	
	15% 28%		¥5.	<b>%</b> &	1575 2965	<b>\$</b> \$	160 171	8.0 10.0	
	16.3		<b>%</b>	155	17.3	50	73		Sax. 116, 12.5
	25.9	o.		191	86.9		83	5.7	alone; Sem. Lif-121,
23.8	53.8	æ	<b>%</b>	<b>76</b> 1	55.0	50	<b>10</b>		Sar, 114.

Table 6. (Continued)

		j			<b>.</b> 25	± 123, € 123,				وجي کا	13% 50%
	Remarke				Sem. 122, 25	alone; Sem.127-129	36m. 122.			Sem.130,50 rom Hg CH <sub>h</sub> alone;	Sem.151-136, %2 added to Sem.130.
f & ( ) d	(cm-1) (1550 cm-1 bend	9.6	14.7	21.0	12.0	4.41	16.7	21.3	30.2 34.9	19.5	24.7 25.3
(4/4)	1306 cm <sup>-1</sup> band 1550 cm <sup>-1</sup> ba	130 254	180 195	208	115	135	154 179	193 209	222 234	175	199 214
cm-lband	Fig. No.	50	20	50	50		50 50	50	50	น	51
1550 \$ 1306 cm-1band	(m) Hg)	151	748 1525	3040	34.5	59.5	110 242	410 750	1530 3025	8	133
	1.00 1.00 1.00 1.00 1.00 1.00 1.00 1.	233	33.0 574	<b>†17</b>	237	593	1.7.6 1.7.6	361 413	450 479	360	07 <del>1</del> 1
-	F16. JA(v. Rg. No. (cm	75.73	5 9 4 9	<b>3</b>	,ç,	\ <b>८</b> स्र	94	91	94	97	\$ <sup>2</sup>
300	P Hg	55 56 56 56	747 1525	3040	32.5	37.5	240	406 7448	1530 3030	65	612 ६२
	(a.znos. cm) (zm Hg.)	23.8 23.8	23.3 23.3	23.8	47.3	6.74	47-3 47-3	47-3 47-3	47.3 47.3	95.0	95.0
	Sample No.	11. 11. 11.	677	127	83	ध्य	<b>3</b> 53	<b>8</b> 8	8 B	8,	£ 55.

Table C. (Continued)

		202	7	1	70.0	1			
Sample Ko.	(atmos. cm) (am Hg)	P Rg)	Fig. (A)	(cm-1)	P. F18.	Fig.	JA(v)dv JA(v)dv (cm-1) (cm-1) .505 cm-1 band 1550 cm <sup>-1</sup> band	JA(v)dv (cm <sup>-1</sup> ) 550 cm <sup>-1</sup> band	Remarka
	95.0 95.0	413	,54	1.80 5.6	41° 759	15	227 242	31.4	
13.35	92.66 95.6	1540 2830	<b>%</b>	% % %	1545 2835	51	257 267	0.11.0	
137	18t	1,50	Ą	ट <sup>म्म</sup>	1,78	댅	220	32.8	Sam.137,130 nm Hg CH <sub>ll</sub> slone;
<b>8 3 3</b>	186 186	232 235	ጀ	LB5 522	643 643	51 51	238 253	39.0 44.2	Sam. 136-142 No added to
1:40 1:41	1.86 1.88	776 1540	¥	550 597	784 1550		265 380		Sem137.
2	188	2950	3	628	2950	51	290	61.5	

Table /. Summary of Empirical Constants

1.  $\int A(v)dv = c \left[ vP_e^a \right] c$ 

II.  $\int A(v)dv = C + D \log_{10} \left[ \Psi P_{e}^{B} \right]$ 

A(v)dv can be calculated by use of tabulated empirical constants in I and II above for values of  $P_e$  and A(v)dv within specified limits

Dand (cm <sup>-1</sup> )	Band limits (cm <sup>-1</sup> )	c	ь	<b>a</b>	С	D	P <sub>e</sub> limits (mm Eg)	JA(v)dv limits (cm <sup>-1</sup> )
يانونود د - الا	3100 - 2300	18.0	0.53	0.7			10 - 25	10 - 45
N2Ω				0.7	15	40	10 - 760	45 - 120
2143	1975 - 2275	2.75	0.55	0.8			20 - 250	4 - 40
		2.75 3.20	0.55					10 - 40 40 - 120
3020 CB4	2ª00 <b>~</b> 3400	15.5	0.55	1 0.4	-106	61	20 - 3000 10 - 760	120 - 200 15 - 250
130%		7.3	0.45		-375	545	10 - 3000 10 - 760	250 - 500 20 - 150
CH <sub>H</sub>	1050 - 13551			4.6	-190	115	10 - 760	150 - 500
1550 cm Odi <sub>k</sub>	- 1535 - 1800	ಗ≎ ≎ಗ	pirica	l agu	# t] uun	أدهان	ved.	

<sup>\*1535</sup> cm-1 is houndary between 1806 cm-1 and 1550 cm-1 CH4 behale; other band limits represent frequencies beyont which there is negligible absorption for samples used in the present 250My.

Table 9. Summary of Band Intensities

Band	fand Assignment	∫k(+)d+ atmos=1 cm=2
222+ m-1 #50	٧3	1350 ± 185
2145 cm - 20	v <sub>1</sub>	260 ± 26
4260 : <b>x</b> *1 30	2 <b>4</b> 1	Not wearing
New many mag	٧٠	520 ± <b>48</b>
1306 mm 1 18 <sub>6</sub>	₽ <sub>14</sub>	185 ± 28
1550 m <sup>-1</sup> 18 <sub>4</sub>	₹2	2.5 ± 0.5

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#### APPENDIX T

#### ADSORPTION EFFECTS

It is a well-known phenomenon that gases are adsorbed on the surfaces of containers. Because of the sampling procedure used in the present study it was necessary to investigate the possibility of sampling errors which might arise as a result of such adsorption.

The absorber concentration w for a given sample was determined by measuring the partial pressure of the absorbing gas. If one could assume that a negligible amount of gas was adsorbed by the cell vindows, the amount of absorbing gas traversed by infrared radiation would be given by the product of the cell length and the partial pressure of the gas, regardless of any adsorption on the cell walls. However, when nitrogen is added to a sample of absorbing gas as was done in the present investigation, the question arises as to how much absorbing gas remains in the optical path. It seems possible that the nitrogen might displace some of the absorbing gas from the cell walls; this effect would result in a greater amount of absorbing gas in the optical path.

Penner and Weber<sup>21</sup>, in a study of infrared absorption by GO at high pressures, found that the CO concentration did increase as nitrogen was added to produce high total pressures. They attributed the increase to an "adsorption-desorption" process in which CO was adsorbed on the cell walls and then was displaced by  $N_2$  which was added later. The CO concentration in portions of the sample taken from the cell was determined by Penner and weber with a mass spectrum ter.

In the present study to mass spectrometer was available to use to monitor the absorber concentration. However, reverst experiments, which involved the three gases CO, H2O, and CH<sub>0</sub> in the three different absorption cells, were performed in order to gain a better understanding of such an adsorption - description process and to determine which sampling procedures could be expected to yield sintant sampling error. As a result of these experiments, along with what is generally known about adsorption, the following conclusions were made and the corresponding sampling tiredures were decided upon:

(1) It was found that the adsorption was essentially complete within a period of a few seconds after the absorbing gas had been added to an evacuated cell. This phenomenon was checked by adding a sample of absorbing gas (RgO, CO and (Pg)) to a cell while monitoring the infrared absorption by the sample. The infrared absorption was inver found to decrease significently after the sample had been added, thus indicating that no absorber was disappearing from the optical path as a result of slow addription by the cell valls. This experiment, of course, did not prove that there had been any adsorption at all; but initial adsorption was confirmed by the next experiment.

(ii) In the second experiment, come absorbing gas was added to an initially evacuated cell; No was then added and an absorption spectrum was obtained immediately. The sample was allowed to remain in the cell and spectra were recorded periodically for several hours or even days. in most cases the infrared absorption was found to increase, indicating that the amount of absorber in the optical path was increasing. This increase was found to take place rather slowly; within a few minutes only a elight increase in absorption was observed, but after a few hours the increase in absorption might correspond to as much as a 10 to 20% increase in absorber concentration in extreme cases. Because of the slowness with which the "desorption" took place, the sampling error was minimized by making a set or runs within as short a time interval as was practical. In no case were data used which were obtained from samples left in the absorption cell over night. In most cases samples were not in the absorption cell more than two - three hours.

(iii) It is to be expected that the relative amount of gas adsorbed and displaced by No decreases with increasing partiel pressure of the absorbing gas. It is known that the amount of was adsorbed is approximately proportional to ya where p is the pressure of the gas and a has a value that may not be constant but is always less than unity; Thus, it is to be expected that the fraction of the gas that is adsorbed decreases with increasing pressure. This result was also confirmed experimentally in the present study: a negligible increase in infrared absorption, if any at all, was observed after adding No to samples having partial pressures of the absorbing gas in excess of 200-300 mm Ng. The greatest relative increase was observed in the case of samples in which the partial pressure of the absorber was between I and 10 mm Mg. For this . cason very few results are included which involve samples formed by adding No to samples of abscribing me baving a partiel pressure less than " or 5 me Mg. Short absorption cells (1.55 cm and 6.55 cm) were used in order to nheat remains of low abouties concentration without using was easiwaly low partial pressures of the abcording cas.

(iv) Considerable time is required to remove the adsorbed gas from the cell valls by pumping, a phenomenon vell-known by anyone who has devit with high-vacuum systems. However, some cell for results of the experimental checks of the phenomenon seek forther of discussion. A hypical experiment might so as foliate: A sample with a pressure of approximately 100 mm Mg of MgO was put in the cell and allowed to remain for a few simules. The null was then opened to the vacuum pump and pumper for approximately one minute after the pressure had been reduced to 0.1 mm Mg. The valve to the cell was then closed and a spectrum of the 22th cm<sup>-1</sup> region was immediately accounted with the cell in the optical path. Mo shourption of

the infrared radiation could be detected; but if the cellwere allowed to remain for several minutes, or even hours, it would be found that same radiation was being electred, and the agoust of absorption increased with time. This result indicates that some of the annurum SpI was leaving the walls of the cell and empering the optical path. 3

A second now experience experiment, algir typically go as fallows. The call was filled with EpO and pusped as lattice, but instead of allowing the call to remain emounted, Epusa added to a pressure of Toll to Epuse a spectrum was nativable in a pressure of Toll to Epuse a spectrum was nativable if any, changed on the infrared remainship, indicating it by the far the optical parts. Someone, spectra changed periodically afternated would soon increasing electromation and periodically afternated would soon increasing electromation on the letter once the increase or analytical soon and regions in the letter once the increase or analytic experiment. In the letter once the increase or analytic experiment, in the letter once the increase or analytic resourced acre regions to being displaced by Ep.

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which were formed by first introducing some  $N_2$  (say 20-30 mm Hg), then adding the adsorbing gas, end finally more  $N_2$  to the desired pressure. It is presumed that the first  $N_2$  introduced was sufficient to form a layer of  $N_2$  on the cell valls before the absorbing gas was introduced. The amount of absorbing gas which was adsorbed on the valls was apparently reduced because of the  $N_2$  layer on the cell valls.

This result suggests that the samples should be formed by adding M2 before the absorbing gas, then by adding more M2 to the desired pressure. Although it would still be difficult to be sure, one might expect that some absorbing gas would displace some M2 from the walls, and the amount of absorbing gas in the optical path would decrease. This method of forming the samples has the disadvantage that the partial pressure of the absorbing gas cannot be measured to accurately when it is being added to M2 as when the cell in evacuated. For example, if M2 were added to a cell to a pressure of 100 mm Mg and M20 added to this until the pressure was 105 mm Hg, the partial pressure of M20 (5 mm Hg) could not be determined as accurately as if the M20 were added to a evacuated cell and the pressure measured with the absolute pressure gauge.

The method of adding H<sub>2</sub>, then the absorbing gas, and finally move H<sub>2</sub>, has the further disadvantage that no spectra of the sample can be obtained for values of pressure less than the pressure of the H<sub>2</sub> initially added plus the partial pressure of the absorbing gas. The total absorptions of several samples formed by this method were compared with those of samples believed to have the same values of shearies connextwation and pressure but which were furnit up introducing the absorbing gas 'n the cell before the H<sub>2</sub> was added. If the letter sample was formed by adding an absorbing gas to the cell after it had seen rimbed with H<sub>2</sub> and pumped only sufficiently long to reduce the greature to approximately U.SZ-0.1 mm H<sub>3</sub>, it was found that the agreement was sood. This result suggests that a layer of K<sub>1</sub> still remained on the cell value efter the short pumping period and served to reduce the adsorption of the absorbing gas.

Another method used to estimate the amount of sampling error arising from autorption-description effects whe to nompere spectra of samples in two different cells but with values of absorber concentration and equivalent pressure believed to be equal. (Equivalent pressure to defined above in the notices on self-broadening.) For example, 8,0 might be added to the 1.55-on cell to a pressure of 20 at 1.55/0.35 = 5.0 mm Mg. Mg was then added to the first-one cell until the equivalent pressure was the same in both cells. Buth cells then had the ease absorber concentration except for changes in the 6.55-on cell due to adverption or description. Since the aborter cell contained only Mg.0. To error in the determination of the absorber concentration should arise from adverptions or description. Comparison of the absorption spectra of the two camples served to give an indication of the amount of adverption-description which tank plane in the cell containing both Mg.0 and Mg.

Samples of very small absorber concentration for CO and N2O were formed by adding to the cell mixtures containing a known fraction of m-sorbing gas (1/10 or 1/50) in N2. The mixtures were unde in a separate tank by adding N2 first, and then the proper amount of absorber (CO or N2O) in order to reduce the adsorption of the absorber or the walls of the tank. The mixture was formed at pressures sufficiently high so that the relative error in measuring the partial pressure of thereboorbing gas was small. An electrically driven fan in the tank was used to incure uniform mixing.

A portion of the mixture was then admitted to the absorption cell after it had been flushed with H2 and evacuated. By this method it was possible to use samples having a partial pressure of the absorbing gas equal to a fraction of a mm of Hg, without having to measure pressures less than 5-10 mm Hg. This procedure should increase the accuracy of sampling. Higher pressures were then obtained by adding more H2 to the mixture.

Since the N<sub>2</sub> was added to the mixing tank first, and the cell was flushed with N<sub>2</sub> before being evacuated, one would expect on the basis of the previous discussion that these containers would have a layer of N<sub>2</sub> on the walls which would tend to reduce the adsorption of the absorbing was. One would further expect that what adsorption did tabs place would occur slowly after the sample was admitted to the cell, and that any change in the absorber concentration would be a decrease. This is in contrast to the expected increase in absorber concentration caused by the displacement of the absorbing was by N<sub>2</sub> when the absorbing was admitted first in the usual way. Several samples formed by mixtures were majored with samples which were believed to have the same values of absorber concentration and pressure but were formed by suding N<sub>2</sub> to the cell after the absorbing gas was admitted. This comparison served as another indication of the annual of error caused by alsorption or description.

It should be noted that the experiments performed were issufficient to prove that the processes described above, such as the adsorption, the displacement of one gas on the cell walls by another, the forming of a layer of one gas, etc., actually occur in the manner suggested. It can make be sent that these proposed processes need resonable in view of want is to be found in the internative concerning adsorption, and that they would account for the results of the many sampling experiments which were performed.

As a result of these experiments it can be concluded that, in studies such as the present one, care must be taken to prevent emeasive compling error from adsorption effects. It is believed that, by following the campling procedures just described, errors in the determination of total absorption arising from adsorption effects were held to less than two or three percent for supples having the partial pressure of the absorpting man greater than 50 mm 3g. And to less than five servent for even the smallest sumples used.

## PART B

ABSORPTION BY CARBON DIOXIDE

by

Darrell E. Burch, David Gryvnak and Dudley Williams

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### ABSORPTION BY CARBON DIOXIDE

#### I. THIRODUCTION

This report is the second of three to be published in the present research program, which involves the investigation of the infrared absorption by gases occurring in the atmosphere. Included in this report are the results of the investigation of the absorption by CO<sub>2</sub> bands at 3761 and 3609 cm<sup>-1</sup>, at 2550 cm<sup>-1</sup>, at 1064 and 961 cm<sup>-1</sup>, and in the spectral range from 875 - 495 cm<sup>-1</sup>. The first report<sup>1</sup> (hereafter referred to as Report 1) in the present program contains results of absorption measurements of the following bands: 2224 cm<sup>-1</sup> H<sub>2</sub>O<sub>2</sub> 2145 cm<sup>-1</sup> and 4260 cm<sup>-1</sup> CO; and 3020 cm<sup>-1</sup>, 1506 cm<sup>-1</sup> and 1550 cm<sup>-1</sup> CH<sub>0</sub>. The third report<sup>2</sup>, which will be called the Final Report, will include results of the measurements of other H<sub>2</sub>O bands and some strong H<sub>2</sub>O bands as well as a section dealing with the "relative broadening abilities" of different gases. The contents of these three reports are intended to supplement earlier work reported by Foward, Burch, and Williams (hereafter referred to as HM) which included measurements of CO<sub>2</sub> and h<sub>2</sub>O absorption.

The work by HBM was limited to frequencies greater then approximately 670 cm<sup>-1</sup> because of the limitations of the spectrometer. Since a large portion of the absorption by CO<sub>2</sub> in the 675 - 195 cm<sup>-1</sup> region occurs at frequencies below 670 cm<sup>-1</sup>, it seems desirable to study the absorption in the entire spectral region, 675 - 195 cm<sup>-1</sup>, by the use of apparatus developed since the HBM work was completed. Besides extending the spectral region covered, it has been possible with the new apparatus to obtain absorption measurements which are comewhat more accurate and include ramples corresponding to aborter paths.

Same as commutating smaller amounts of CO<sub>2</sub>, and, therefore, corresponding to sourcer paths than those employed by RMs were used in the investigation of the absolution by the banks at 5'to. When, and 2550 cm<sup>-1</sup>. The two banks at 1054 and 901 cm<sup>-1</sup> are so week as to be of only minor importance in the study of the certh's atmosphere; however, a knowledge of the atmosphere makes a containing large quantities of CO<sub>2</sub>. For this reason as in the study of the atmospheres of other planets containing large quantities of CO<sub>2</sub>. For this reason as investigation of these banks was included in the present research program.

Several experiments which have been proposed recently involve the measurement of the absorption and the melanton of atmospheric CO<sub>2</sub>. These measurements will be used by the use of instruments in airylanae, ballooms, earth natellites, and even planetary probes. In order to interpret data obtained in these experiments it is necessary to understand CO<sub>2</sub> absorption and causaious as. Its dependence on some parameters as path length, temperature, absorber partial pressure, and total pressure. It is based that the present data will further the knowledge of CO<sub>2</sub> absorption and its relationship to these parameters. The data will, as in earlier reports, by presented by the use of absorption curves, tables, graphs, and empirical enteriors.

## A. UNITS, SYMBOLS, AND DEFINITION OF QUANTITIES

Spectral regions and frequencies are expressed in terms of frequency with units of cm<sup>-1</sup> throughout the present study, except in reference to other work. All tabulated values of partial pressure, p, total pressure, P, and equivalent pressure, Pg, are expressed in un Mg; and absorber path lengths are in cm. Values of absorber concentration were in units of atmos cm, and are determined by multiplying the absorber path length by the partial pressure of the theoreting gas in atmospheres, corrected to atmosphere corrected to in degrees centigrade, except in instances where it is advantageous to use degrees Felvin, for which the symbol # is used.

A'(v) is the true fractional absorption at frequency v, which would be observed with an instrument of infinite resolving power. A(v), the observed fractional absorption at v, depends on the spectral slit width and is expressed as a fraction or as a powentage. The total absorption of an absorption band JA(v)dv is equal to JA'(v)dv and is independent of the spectral slit width provided the limits of integration include the entire band.

In section VI a quantity solled the seas fractional absorption is used; this quantity is given by

$$\bar{A}(v_1 - v_2) = \left(\frac{1}{v_2 - v_1}\right) \int_{v_1}^{v_2} A(v) dv$$
, (1)

and represents the mean resctional absorption in the spectral region  $\nu_1 = \nu_2$  .

## 11. EXPENSIONERAL PRESIDENCE AND AXALTHIS OF DATA

, we different epectrometers were need to obtain the absorption curve: for the present investigation. A forbin-Einer Model 21 equipped with dust multiple-pass absorption critic was used for absorption paths between w00 and \$207 cm. Thereis absorption cells of lengths 1.55 and 17.8 cm were exployed with a Forbin-Einer Model 99 spectrometer which was exclosed in a vacuum tank. All samples were compared of either  $CO_2$  alone or of mixtures of  $CO_2$  and  $R_2$ . The exchangular used in sampling, as well as the investmention of the accomplion cells, and the method of determining the total absorption from the absorption curves were discussed in Report 1.

The The date reported in the present study can conveniently be duried in the four spectral regions which are referred to as follows:

(1) the 5716 and 5600 cm<sup>-1</sup> tanis, (1); the 250 cm<sup>-1</sup> tend, (111) the 1664 and 961 cm<sup>-1</sup> tanis, and (iv) the 875 - 475 cm<sup>-2</sup> region. The data for each of these spectral regions were obtained separately and the results are presented in sections III. IV, V, and VI, respectively

Each of these four spectral regions covers a rather wide spectral region. It therefore mened advantageous to divide each region into smaller sub-regions and to determine either the total absorption or the muan fractional absorption of each sub-region. It is desirable to divide the larger spectral regions into small sub-regions in such a way that the total absorption determined for each sub-region is virtually independent of the spectral alitavidah and upuas to the value videb would be determined with infinite resolving power. As mentioned earlier, the total absorption JA, v)dv of an entire band has been shown? to be independand of the openional alit-width, provided there is no absorption beyond the limits of integration. It follows that the total absorption of the sub-regions wild be independent of spectral slit-width if they were divided at frequencies where there was zero absorption. However, it will be seen in sections III - VI that there are no frequencies within the larger spectral regions at which the shaorption is zero for samples having large values of v. To was therefore necessary to choose optimus frequencies, at which there was absorption, as boundaries between the sub-regions such that the dependence of total absorption of each sub-region was least dependent on the spectral slit-width.

If  $v_i$  is the frequency at which two sub-regions are divided, the observed  $\lambda(v)$  just to one side of  $v_i$  is partially due to absorption at frequencies slightly to the other side of  $v_i$ , as a result of the finite spectral slit width  $\Delta v$ . The boundary frequencies were cases so that the fractional absorption was virtually symmetrical about  $v_i$  for approximately  $\Delta v$  wave numbers on each side of  $v_i$ ; i.e., the fractional absorption was nearly equal at a given distance on either side of  $v_i$ . If the fractional absorption is symmetrical about  $v_i$  and the spectral slit function is also symmetrical, one would expect the total absorption of sub-regions divided at  $v_i$  is have very little dependence on spectral slit-slith width.

The boorption curves for the 3716 and 3609 cm<sup>-1</sup> bands were sivided into two megious with the division at 5600 cm<sup>-1</sup>, which corresponds to minimum in absorption. At will be seen in section III that the absorption corresponds to make the absorption curves are nearly symmetric correct 5000 cm<sup>-1</sup> for a few cm<sup>-1</sup> car cash side, thus making it a nottable frequency at which to divide the bands. The total absorption measured from the portions of the absorption curves at frequencies greater than 5600 cm<sup>-1</sup> is referred to as the 5716 cm<sup>-1</sup> band; and that on the low-frequency side of the interval is referred to as the 3600 cm<sup>-1</sup> band;

The 100 and 961 cm<sup>-1</sup> bends are visitarly divided at 1000 cm<sup>-1</sup>, the frequency corresponding to an absorption minimum between the backs. For the targer camples studied, the fractional absorption did not decrease to zero on the low-frequency side of the fill the land, but

reached a minimum at approximately 875 cm<sup>-1</sup>, below which the fractional absorption increases. The limits of the 961 cm<sup>-1</sup> band were therefore chosen to be 1000 cm<sup>-1</sup> and 875 cm<sup>-1</sup>.

The 2350 cm<sup>-1</sup> hand has divided at the band center and the total absorption on each side of this boundary was determined. The total absorption of the entire band was found by adding the values of total absorption of the two portions.

There are several etrong and medium bands in the spectral region between 875 and 595 cm<sup>-1</sup>, the etrongest band having its center at 657 cm<sup>-1</sup>. The frequencies curresponding to the absorption minima are not the same for all samples; and those frequencies are therefore not suitable for bundaries between the sub-regions. The major portion of the absorption arises from bands having prominest § branches. The frequency of maximum absorption to each of these § branches is essentially the runs for all samples; and the absorption curres are courl to each side. On the basis of the above discussion, it was decided that the conters of the § branches at 730, 667, 617, and 545 cm<sup>-1</sup> would be used as boundaries between the aportical sub-regions. For even the largest samples used, the absorption was essentially more at frequencies below 95 cm<sup>-1</sup>, and absorption at frequencies greates these 875 cm<sup>-1</sup> was treated as a part of the 961 cm<sup>-1</sup> band. The 875-895 cm<sup>-1</sup> region was increated as a part of the 961 cm<sup>-1</sup> band. The 875-895 cm<sup>-1</sup> region was increated as a part of the 961 cm<sup>-1</sup> band. The 875-895 cm<sup>-1</sup> region was increated as a part of the 961 cm<sup>-1</sup> band. The 875-895 cm<sup>-1</sup> region was increated as a part of the 961 cm<sup>-1</sup> band. The 875-895 cm<sup>-1</sup> region was increated as a part of the 961 cm<sup>-1</sup> band. The 875-895 cm<sup>-1</sup>, pagion was increated as a part of the 961 cm<sup>-1</sup>.

"so men fractional absorption, which is related to the stription by Eq. (1) for each of those sub-regions, is related to the parameters y and Po by curves given in Section VI. The quantity "mean fractional absorption" return than "buttl absorption" was found for each of the sub-regions in the 575-595 or 5 region. This was demo class that term "betm" observation structured absorption for the interval v<sub>1</sub> - v<sub>2</sub> is equal to the fraction of the relations power in that interval which is absorbed, provided the incident relation to the seaso for all frequencies.

# A. AND-MOADMING CONFACINITY AND DESITALIST PROBERTS.

The absorption of a binary mixture of \$\ \text{and an absorbing gas baving a given absorber consentration y is known to depend on the partial procure p of the absorber as will so upon the total pressure P. This phononome is a result of the different breakening attitities of the two levelet. Soveror, it was found that the breakening affects of you gener can be expressed in terms of a single variable called the equivalent pressure \$\frac{1}{2}\$ which is defined by

$$r_n = p + (p-1)p = p_{p_n} + p_p \tag{2}$$

where B, the self-broadening coefficient, can be determined engarimentally.

The value of B was determined by the method described in Report 1 to be approximately 1.30 for the CO<sub>2</sub> bands included in the present study. It was shown in Report 1 that the value of B may be somewhat different for different bands of the same gas; it may even be alightly different for different frequencies within a given band. Although there may be small differences in the value of B for different frequencies, the value of 1.30 is believed to be correct to within ±6 to ±8% for wirtually all portions of the bands studied.

By substituting the value R=1.30 into Eq. (2), one obtains the following expression for the equivalent pressure of the  $\rm CO_2-R_2$  samples:

$$F_e = P + 0.50p = p_{\overline{N}_2} + 1.50p$$
 (5)

It is noted that no quantity defined to be equivalent pressure was used by HBW. However, the quantity (P+o), which was called a "weighted pressure", was used as a single parameter in the same manner that  $P_{\rm g}$  is used in the present study. Using (P+p) as an equivalent pressure is, of course, the same as using a value of B = 2, which is to be compared with the value of 1.30. The latter value is based on rather careful with the value of excently and is in good agreement with theoretical predictions by two well-known theorists, Lewis Kaplan" and William Benedict. Edwards in a recent study of the total absorption of CO<sub>2</sub> bands used an expression for  $P_{\rm g}$  which is equivalent to the following:

$$P_e = (P + p + 0.5 p^2)$$
. (pressure in atmospheres)

This expression for  $P_{\rm e}$  is equivalent to using B = (2 + 0.5p), a value which is greater than that used in the present study or by HHW.

B. DEFICIENCIES OF DOUBLE-BEAMING TECHNIQUES IN THE STUDY OF STRONG ATMOSPHERIC BANDS

It is well known that, when a conventional ringle-beam spectrometer is employed with the optical path open to the atmosphere, there are spectral regions over which the radiant power incident on the detector is required as a result of absorption by atmospheric games. It will be shown

<sup>&</sup>quot;The Final Report will contain more detailed information concerning the determination of B in a section devoted to the relative broadening abilities of various gases.

that care must be taken in order to determine the correct value of the fractional absorption  $A(\nu)$  of a sample at frequencies in a region of atmospheric absorption. It will also be demonstrated that atmospheric absorption within a double-beam spectrometer can give rise to error an the measurement of the fractional absorption of some samples, notably atmospheric games. In order to demonstrate these phenomena the following qualities are defined:  $R_0(\nu)$  is the radiant power incident on the detector with the spectrometer set at frequency  $\nu$  if there is no absorption by the atmospheric gases in the optical path. Ine power  $R_0(\nu)$  includes radiation within the frequency interval As passed by the spectrouster slit. The spectral intensity distribution of the power R. over the interval as depends, of course, on the spectral intensity distribution of the source as well as upon the "slit function" of the spectromster. In spectrometers such as the ones used in the present study the intensity of the source is nearly constant over my and the spectrometer wift function is such that the intensity of  $R_0(\nu)$  changes gradually over Aw with its maximum at frequency v.

If a sample is placed in the bear and the radiant power incident on the detector is reduced to  $R(\nu)$ , the fractional absorption  $A(\nu)$  of the sample is defined to be

$$A(\nu) = \frac{R_O(\nu) - R(\nu)}{R_O(\nu)} . \tag{4}$$

In the case of a sample for which  $A^*(\nu)$ , the true fractional absorption as viewed with infinite resolving power, changes rapidly, it is apparent that  $A(\nu)$  will depend on the value of  $\Delta \nu$  and on the slit function.

If frequency  $\nu$  occurs within an absorption band of one of the atmospheric gases present in the optical path, the radiant power incident upon the detector is reduced to  $R_0^2(\nu)$ . If the sample having fractional absorption  $A(\nu)$  is placed in the beam, the power meaching the detector is reduced to  $R^2(\nu)$ . The apparent fractional absorption of the sample  $A^2(\nu)$  is then given by

$$A^{\phi}(v) = \frac{R_{\sigma}^{\phi}(v) - R^{\phi}(v)}{R_{\sigma}^{\phi}(v)}. \qquad (5)$$

\*

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Care must be taken not to assume that  $A^*(\ ) = A(\nu)$ . This equality is approximately valid in many cases; however, it has been shown by #FW that  $A^{\pm}(\nu) < A(\nu)$  if the sample being studied is the same gas as the atmospheric panerousing the absorption. The inequality immediately above can be explained in an elementary way if one thinks of the radiation as having ressed through the atmospheric path before being incident on the sample was. (Actually it does not matter whether the radiation passes through the sample first or last, but for the purpose of discussion it will be considered as having passed through the sample last.) The intensity of the radiation incident on the sample will not be constant as in the case of no atmospheric absorption, but it will have minima at frequencies corresponding to absorption lines and maxima at frequencies between the lines. Since the sample being studied is the same gas as the atmospheric component giving rise to the absorption, the radiation incident on the sample is greatest at frequencies for which it absorbs least and is least at frequencies for which it absorbs most. The net result is that the fraction  $A^{*}(v)$  of the radiation absorbed by the sample is less than  $A(\gamma)$ , the fractional absorption observed with no atmospheric interference.

It is apparent that if the sample being measured is different from the atmospheric cas giving rise to the absorption,  $A^{\mu}(\nu)$  might be greater or less than  $A(\nu)$ , depending on the relative positions of the absorption lines of the sample and the atmospheric gas. If the absorption lines of the sample appen to occur at frequencies between the lines of the atmospheric gas,  $A^{\mu}(\nu)$  would be greater than  $A(\nu)$ .

In a double-beam type of instrument such as the Perkin-Elmer Model 21 used in the present installigation, error is also introduced by the atmospheric gases in the optical path when the absorption of  $CO_2$  or  $E_0O$  samples is being studied. Such an instrument records directly the ratio of the transmission of the sample beam to the transmission of the reference beam. If  $R_0^n(\nu)$  is the radiation power incident on the detector from the reference beam,  $[1-A^n(\nu)]$   $R_0^n(\nu)$  is the power incident on the detector from the sample beam which has passed through a sample having fractional transmission  $A(\nu)$ . The apparent fractional absorption recorded by the double-beam instrument is

$$1 - \left\{ \frac{[1-A^{*}(v)] \operatorname{Rg}(v)}{\operatorname{Rg}^{*}(v)} \right\} = A^{*}(v) .$$

Thus, the finite-date instrument grains the same Tiles residing \$4(a)) as a single-last instrument during the same specime, 2.27 finition and the same specimen, 2.27 finition

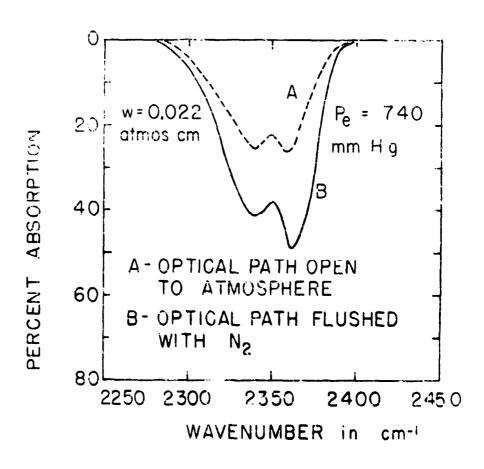
The fact that a limited line trained which is spin to the attraction on small a value of fractional absorption for a We couple is illustrated in Fig. 1. There is no strained with the couple indicated one state the tractionary open is the statement. The apparent fractional absorption is seen to be mostliarably less immegated the band than in more 3, which was interested with the same fag couple with the apparent matter and the spin trained of the present.

he would of reducing the course satisfing from examplement absorptions as to filter the approximation with \$\frac{1}{2}\$, as was done to absorb more \$\frac{1}{2}\$. In success, the argument is an analysis of the filter approximation of the course of the argument is more supported by the absorber monomicalism of the course done are attenuable to more greater than at the attenuable file. The rest are attenuable to success absorption is negligible. In the Trail have the activities of the filter of the approximation approximation approximation are filled and the area attenuable with a service of the file of the file

because the 's consequent's eleminative model. If notice, he eliminately incurred the expectations. Since access improvement, if not improved with the note of the field. It, however, a recouse that was helds to exact me field. By which was used with the 1.5% and 12.5 as while to struck meaning that the field. Since we have the improved with the field of the field of

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The 1. Absorption curves for the 2550 cm<sup>-1</sup> CO<sub>2</sub> band illustrating the error maising from air in the optical path of the apectrometer. The absorber concentration of the sample is approximately one-third as great as the absorber concentration of the CO<sub>2</sub> in the autoupheric path of the sample beam.

## III. EXPERIMENTAL RESULTS FOR THE 3716 AND 3609 cm<sup>-1</sup> CO<sub>2</sub> BANDS

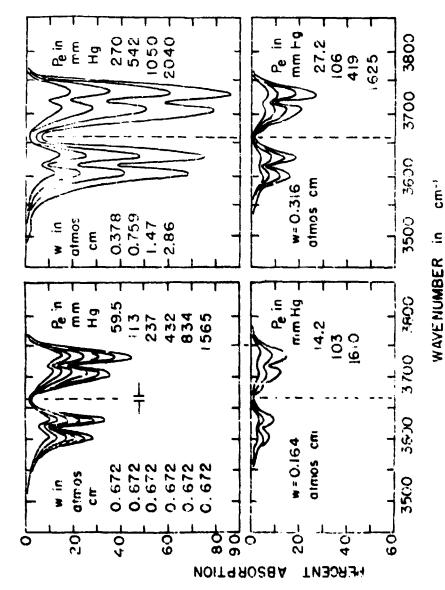
Tracings of many of the absorption curves for the 3716 and 360y cm<sup>-1</sup> bands of 60<sub>2</sub> are shown in Figs. 2 and 3 with the spectral slit-width indicated in the lower left-hand portion of Fig. 3. The wertical dashed line at 5660 cm<sup>-1</sup> represents the boundary between the "bands".

The data are tabulated in Table 1, which includes the values of absorber concentration  $\gamma$  and equivalent pressure  $P_0$  for each sample as well as numbers of the figures in which many of the absorption curves appear. Also included are remarks concerning composition of the sample and the values of total absorption of each of the two bands along with their sum. The value of the sum of the total absorptions of two bands is included for comparison with data reported by HBM and for the convenience of other workers who treat these two separate bands as one. The temperatures of the samples, which were near room temperature and varied by approximately  $z^{ij}$  degrees C, were not included since the total absorption of these two bands varies only slightly with temperatures.

Representative data for the 3716 cm<sup>-1</sup> band are plotted in Fig. 4, where the different curves corresponding to various values of v relate the total absorption to  $P_a$ . A qualitative explanation of the Testures of curves such as those in Fig. 4, as well as of other figures in this report, was given in Report 1 and will not be repeated.

In order to show the dependence of total absorption on absorber concentration, values of absorber concentration were taken from the curves of Fig. 4 and used to obtain the "smoothed" curves of Fig. 5, where each curve corresponds to a given value of  $P_{\rm e}$ . The results of measurements of the 3609 cm<sup>-1</sup> band are presented in a similar manner in Figs. 6 and

In Fig. 6 is shown the sum of the total absorptions of the two bands plotted against  $P_{\rm e}$  for various values of w. The solid curves are based on data obtained in the present investigation, while the dashed curves are based on HDM data which are shown for comparison. This figure permits a direct comparison of the present results with the earlier work for values of y=k.7k and 9.48 atmos m. If one were to interpolate between the two curves corresponding to 3.02 and 6.08 atmos m for values of total absorption corresponding to y=k.7k atmos m, one would find that these values differed by approximately 5 to 105 from values taken from the curve corresponding to y=k.7k atmos m calles taken from the 9.48 atmos m curve is seen to occur close to where one would expect it to be relative to those for 6.08 and 12.7. It should be noted that the HBM data used in the comparison represent the smallest samples investigated. For which one would expect the experimental error to be the greature. The satimated error in the RBM results, as well as the present results, is approximately  $xy \le t$  thus, one concludes that there is fairly



34. 2. Absorption curves of the 3716 and 3:09 cm<sup>-1</sup> Ct<sub>2</sub> bands. The spectral sitt-width indicated in the upper-left jortion of the figure applies to the curves in Figs. 2 and 3.

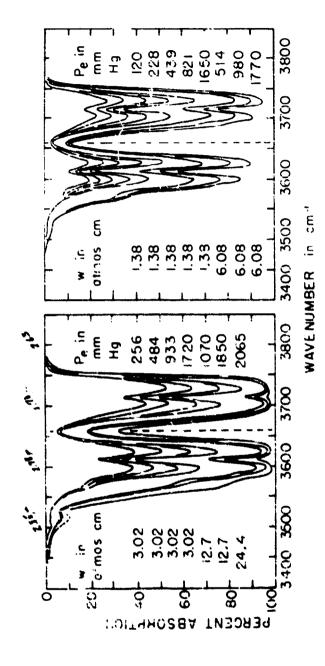


Fig. 3. Assemption curves of the 3716 and 3509 on 1 322 bands.

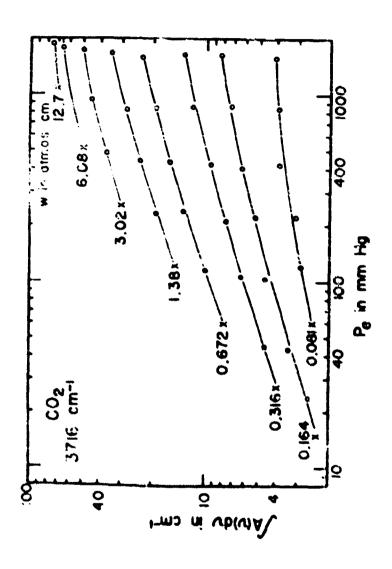


Fig. h. The wiml absorption of the Fils curl (O. tand versus equivalent presents. N's correspond to sample of City alone and O's to mintures of City and H.



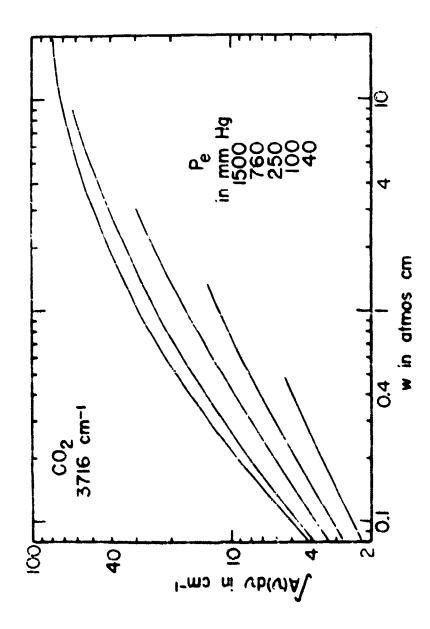


Fig. 5. The termi sheorytion of the 5736 cm<sup>-1</sup> CO<sub>2</sub> band wrate absorber concentration.

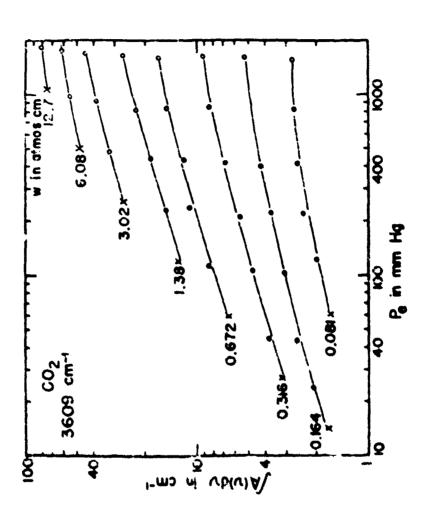
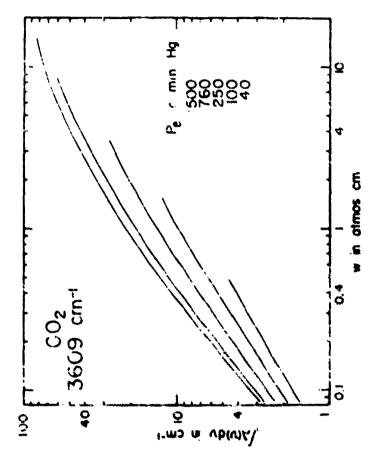


Fig. 6. The total absorption of the 1609 cm<sup>-1</sup> CO<sub>2</sub> tand wereld equivalent pressure. N's correspond to bamples of CO<sub>2</sub> alone, and O's to minture of CO<sub>2</sub> and H<sub>2</sub>.

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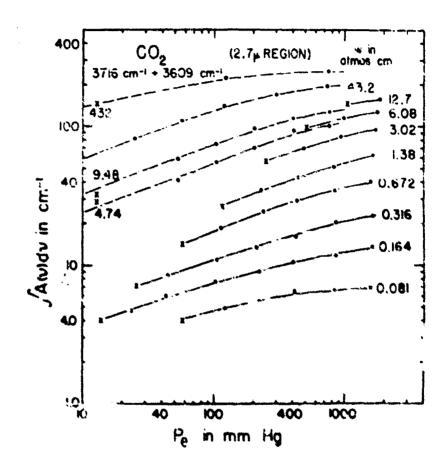


Fig. 6. The title appropriation of the Jill and Jill to 100 tands regress equivalent research. The sould curres are based on this citations in the process investigation; the cartes current are based on the from the MM report and are shown for comparison. I'm narrespend to samples of the comparison. I'm narrespend to samples of the comparison.

satisfactory expresent between the present and the carries were form then of whend  $P_{\rm c}$  repercedby seth

Values of total absorption ... Fig. ", we well as from HEF lath which are not shown in Fig. ", were used to obtain the correct f Fig. , where total absorption is justed assigned for the various values of equivalent organize fadicated. By interpolating netwern one curves of Fig. (1) is possible to determine the total absorption of the 3715 and  $609~\rm cm^{-1}$  bands for samples for which the values of y and  $P_{\rm e}$  are included by the curves.

It is commetimes desirable to have empirical equations which release the total absorption to the parameters  $\underline{\mathbf{v}}$  and  $P_{\underline{\mathbf{e}}}$  although the values was monutedly be determined more accurately from curves such as those in Fig. 9. Two different empirical equations of this sort were reported by HBW for this band: one for values of total absorption greater than 50 cmf. and another for smaller mines. The equation intended for small values of toral alabeation was based on the results of unly approximately 10-12 camples, the maximum pressure of which was less than 50 mm Hg. Many of the samples used in the present investigation yielded values or total absorption less than 50 cm., and the equivalent pressure of these samples varied from a few mm: Hg, to approximately 1300 mm: Hg. Since the present data represent greater ranges of weard Pa, and since the equivalent: pressure used in the present study is believed to be more reliable than the 'weighted pressure" used by HEW, it seemed desirable to revise the empirical equation to be used for values of wital absorption less than 50 cm<sup>-1</sup>.

It was demonstrated in Report 1 that was equation of the form

$$\int_{0}^{\infty} \frac{dx}{x^{2}} = c \left[ \Psi F_{\alpha}^{(1)} \right]^{\frac{1}{2}} \tag{3}$$

was useful for reall values of Alv dv, and that the values of the constants a and b could be easily determined graphically. In order to determine the value of the empirical constant a, values of total absorption were taken from the curves of Fig. 8 for different values of 'P' and plotted against  $vP_a$ ': Different values of a were tried, and the set of curves produced by using a = 0.65 are shown in Fig. 13. The value of 0.65 was selected since in Fig. 16 the five curves corresponding to  $P_a < 760$  mm. Hg are nearly coincident for values of botal absorption between 30 and 30 cm<sup>1</sup>. Since these portions of the curves are marriy coincident it is apparent that the total absorption in he expressed as a function of the plague variable  $vP_a^{(0.5)}$ . For the values of  $P_a$  and total absorption represented. Portions of some of the curves of Fig. 13 were married to prevent of servery, points are included to indicate positions of the curves.

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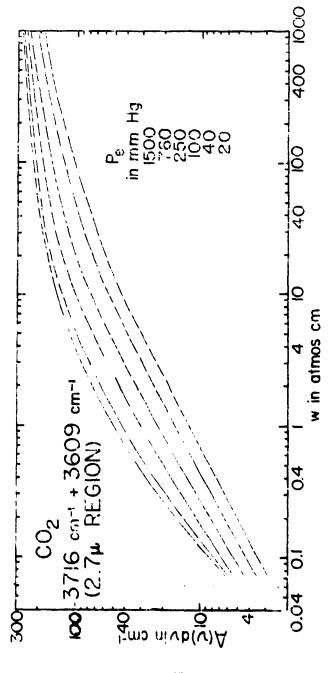
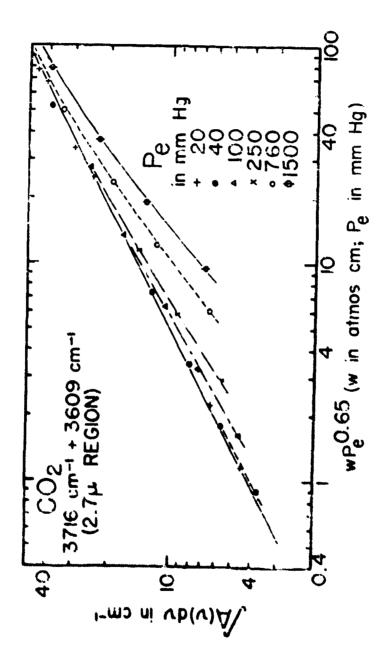


Fig. 9. The total absorption of the 3716 and 5609 cm<sup>---</sup> CO<sub>2</sub> bands were a shorter concentration. The scill portions of the curves are based on days obtained in the present investigation, while the dashed portions are based on data from the IBA meport.





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7:8. 10 The total absorption of the 3716 and 3600 cm<sup>-1</sup> CC2 bands wereas who 0.55.

Thus, if  $F_a$  is limited to less than 250 mm Hg, values of total absorption between 10 and 50 cm<sup>-1</sup> can be expressed as a function of  $\Psi_0^{-0.55}$ . Since both the ordinate and the abscissa of Fig. 10 are logarithmic, the value of b for Eq. (6) was determined from the slope of the single straight line which best fitted the points corresponding to the values of total absorption and equivalent pressure specified. The value of b way, found to be 0.58.

On the basis of the above discussion the following empirical equation and its region of validity were derived for the 3716 and 3609  $\rm cm^{-1}$   $\rm CO_p$  bends.

$$\int A(v)dv = 3.5 \left[ v P_e^{C.65} \right]^{0.58} \tag{7}$$

for

and

$$10 < \int A(v)dv < 50 \text{ cm}^{-1}$$
.

It can be used for  $20 < F_0 < 760$  um Hg if  $30 < \int A(v)dv < 50$  cm<sup>-1</sup>.

This aspirital equation one be compared with the corresponding equation from FDM, which it is intended to seplace:

$$\int A(v)dv = 325 v^{0.5} (Pep)^{0.43}$$

The curves of Fig. 10 which correspond to higher equivalent pressures could be chifted to the left, and made to occur element to the nurve corresponding to lower pressures, by using an exponent a for P. emiliar than the value of 0.64 which was used. However, if the santific exponent were used, the curves corresponding to the high pressures would occur to the left of and above the curves of lower pressure if for larger values of total absorption. Thus our concludes that the t cal absorption and small values of v. This result was discussed in some detail in report 1 and is confirmed in the present investigation. It is probably true that by limiting the ranges of values of |A(v)| and  $P_0$  for each equation and by using different values of  $a_1$ , one could find a series of equation which would eatisfactorily relate |A(v)| to v and  $P_0$  for all the data reported. It is apparent that several such equations would be necessary, and the sech would be valid only over very limited ranges of values of |A(v)| and  $P_0$ . It did not seem surthwails to derive as

many equations as would be necessary; and it is recommended that the curves of Fig. 9 be used for values of  $A(\nu)d\nu < 50$  cm<sup>-1</sup> for which Eq. (?) is not suited.

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## IV. EXPERIMENTAL RESULTS FOR THE 2550 cm<sup>-1</sup> CO<sub>2</sub> BAND

Tracings of many of the absorption curves of the 2550 cm<sup>-1</sup> CO<sub>2</sub> bend are shown in Figs 11 and 12. The data are listed in Table 2 in a manner similar to that used in Table 1. Before measuring the total absorption of such spectrum, it was divided into two portions at 2550 cm<sup>-1</sup>, and the total absorption of such portion was then determined. Values of total absorption of each partion are given in Table 2 along with the sum of these values, which represents the total absorption of the entire band. It hough the results for the two portions are not analyzed separately in the present study, they are tabulated for possible future use by others.

Representative data are presented in Fig. 13 where the total encorption of the entire bend for various values of absorber concentration is plotted equivalent pressure. The solid curves are based on eater obtained in the present investigation. The dashed curve representing EN data corresponding to 4.7 stace on is seen to occur slightly showe the solid curve corresponding to 4 99 atmos on for values of P. between approximately 400 and d00 mm Mg. The values of total absorption indicated by the 4.7 atmos ca curie are approximately i to 05 greater than one would predict on the basis of the rolld curves. Other EW results, which are not shown, were compared and found to be dear than 86 different from corresponding results in the present about. The estimated error in the Ell results as well as the present results is approximately 196. Thus, one concludes that the marketing terms the earlier work and the present work to within the indicated experimental untertalities for values of w and P opvered in both investigations. As is the uses of the 3716 and 3609 cm bands, the EMM results which were compared with the present results represented the samples of least absorber comcentration, for which the uncertainty would be the greatest.

Values of total shearptics were taken from Fig. 14 and from HM data to produce the curves of Fig. 16, share the total shearption is ulothed egainst absorber concentration. By interpolating between the tree it is possible to determine the total absorption of a known sample for which the values of y and P<sub>0</sub> occur within the range coveres by Fig. 16.

In the EDS report an empirical equation was derived to relate values of total absorption less than  $50~{\rm cm}^{-1}$  to y and  $P_{\phi}$ . In order to derive an expirital relationship for the smaller ramples for which the values

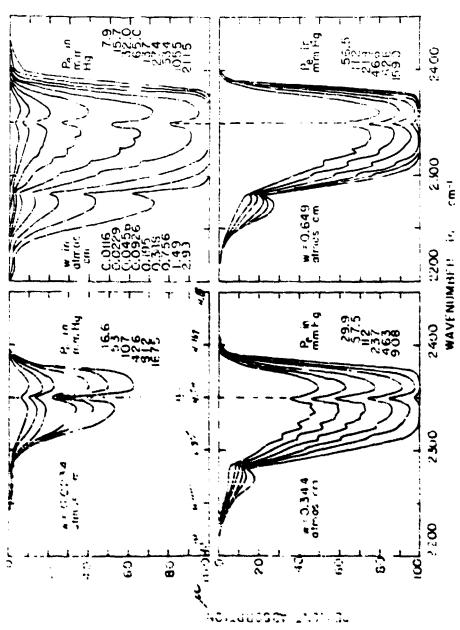
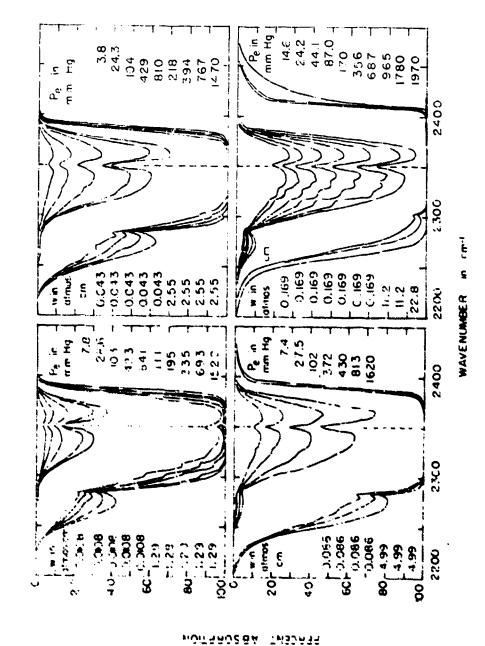


Fig. 11. Absorption sirves of the 2350 cm<sup>-1</sup> CO, land. The spectra eliteviéth indicated in the upper-left kertion of the figure applies to the curves in Figs. 1. and 12.

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Tig. 12. Absorption curves of the 2550 .m. Land.

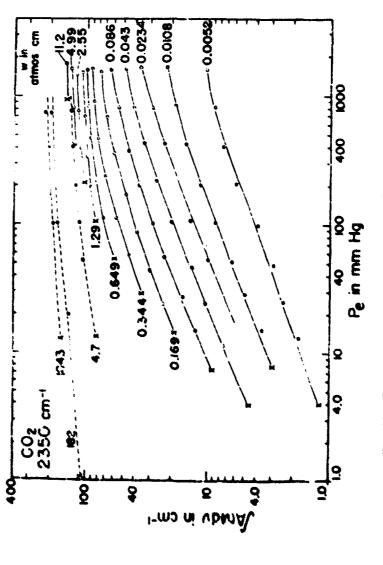
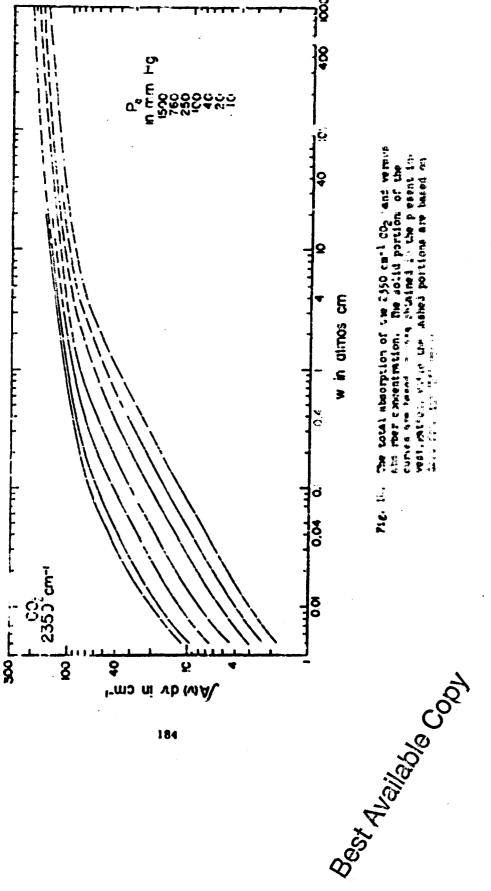


Fig. 15. The total absorption of the 2550 cm<sup>2</sup> CO<sub>2</sub> band versus equivaling grassume. The solid turnes are based on data obtained a the present investigation; the dashed curves are base tool data from the NEW region; and are shown for comparition; "Y's correspond to sampling off CO<sub>2</sub> alone; "O's" and 'C's" to mixture off C)<sub>2</sub> and N<sub>2</sub>.

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The total absorption of the 2550 cm<sup>-1</sup> CO<sub>2</sub> and versus this free concentration. The solid portion of the curves are based a long thinging 10 the peant in vestigation, with the Landau portions are based on this first first the life was a 716. II.

of total absorption are less than 50 cm<sup>-1</sup>, value curves of Pig. 13 and replatted in Pig. 15 again of the expirital constant 0.75 was determined the case of Fig. 10. From this figure the feir and to wall? (In taking of F and Are)4v

The Laken From the up Coll. The two the te same manner as in ag equation was derived ified.

JA(v) 10 - 15.0 (VE, 0.35)0.54

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 $\dot{z} < \int A(v) dv < 50 \text{ cm}^{-1}$ 

10 < P, < 250 mm Hg .

Equation (?) to also walld for

10 < P. < 760 xm Hg

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 $15 < \int A(v)dv < 50 \text{ cm}^{-1}$ .

V. PAPERLINGUEDAL RESULTS FOR THE 1064 AND 961 on 1 BANDS

Tracings of reveral absorption curves of the 1054 and 961 cm<sup>-1</sup> bands are shown to Figs. 16, 17, 10; those in Figs. 16 and 18 correspond to rapples consisting of CO<sub>2</sub> alone, while the curves in Fig. 17 correspond to samples consisting of CO<sub>2</sub> and R<sub>2</sub>. The absorption curves of Fig. 17 were extended to approximately 700 cm<sup>-1</sup>; and the results of the reality of the sections of these curves between 875 and 720 cm<sup>-1</sup> are discussed in the reality rection. The absorption curves in Fig. 16 represent a given sample of the absorption curves above sablent; the results of those curves, as well as others obtained from samples at elevated temperature, 41) to discussed near the end of this section.

It is seen from figs. 16 and 17 that the absorption is essentially sero at approximately 1125 cm<sup>-1</sup> for even the largest samples, and that the bands overlap with an absorption minimum near 1000 cm<sup>-1</sup>. On the law frequency side of the bands another absorption minimum openess at 875 cm<sup>-1</sup>; the vertical dashed lines at 1000 cm<sup>-1</sup> and 875 cm<sup>-1</sup> to Figs. 16-18 represent the toundaries between the bands which were discussed in section 1;

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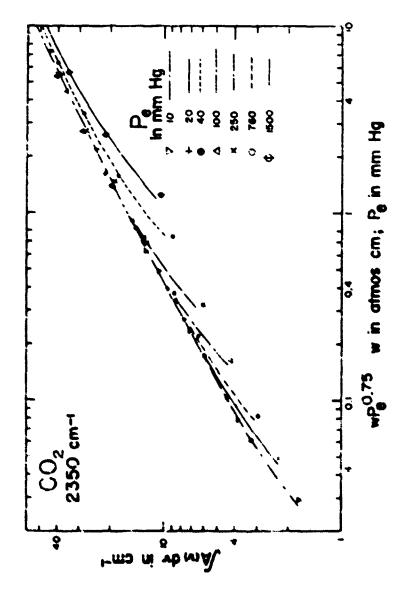


Fig. 1. The total absorption of the Pricial Copies.
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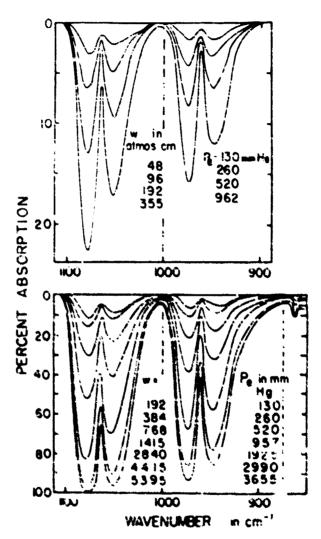


Fig. 16. Absorption curves of the (Non-and 96) curl CO2 bands. The ordinate scale of the approximater was expanded by a factor of 5 while obtaining spectra shown in the upper portion of the figure.

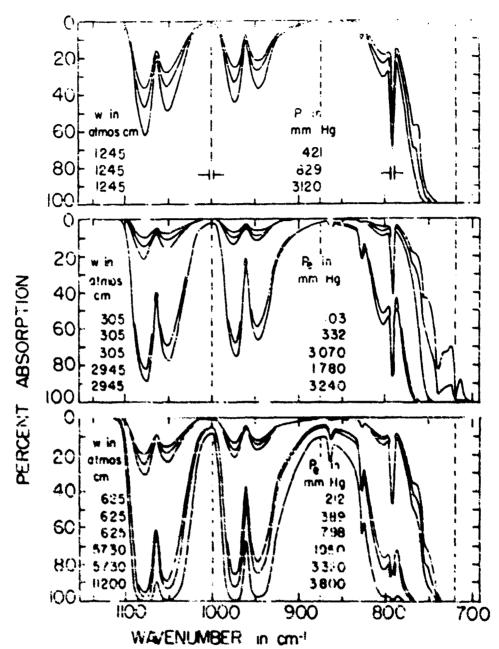


Fig. 1 Allowaystan curves of the 1984 and 981 cm<sup>-1</sup> CD, bands and the 1984 in 1984 in 1984 and 981 cm<sup>-1</sup> CD, bands and the 1984 in 1

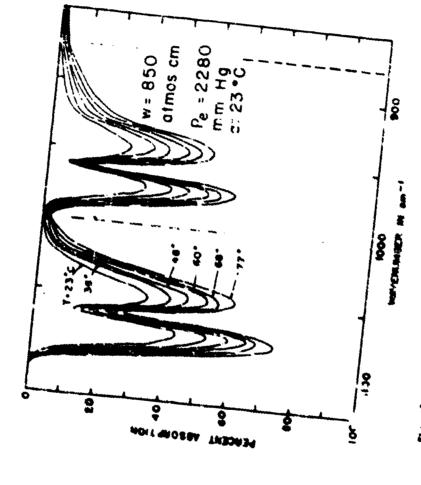


Fig. 18 Amagnites, curves of the 1986 and 16.1 cm<sup>-1</sup> (3) binds for the figure shows and the cm<sup>-1</sup> (3) binds for

The data for the 1064 and 961 cm $^{-1}$  bands are shown in Sable 5, which includes the sample temperature, to the measure 0.5 degrees C, as well as the parameters listed in Sables 1 and 2. Samples consisting of CO $_2$  -  $B_2$  mixtures are limited to those having an absorber path length of 3200 cm.

It is recalled that the total absorption of the 1056 and 961 cm<sup>-1</sup> bands is riveryly dependent on temperature; it was therefore necessary to correct the values of total absorption to the same imperature. The values of total absorption of all the samples studied mear room temperature was surrected to 26°C, which was approximately the average temperature. The necessary corrections were determined from the information given mear the end of this section concerning the growth of the bands with increasing temperature. All figures which relate the total absorption of these heads to v and Pa therefore correspond is possible at 15°C.

In part A of Fig. 19, we value of total absorption of the complex of CO<sub>2</sub> alone are platted against  $P_0$  for the 1054 on hand. Buch entre corresponds to a given path length and  $\gamma$  is therefore proportional to  $P_0$ . In part B of Fig. 19 total absorption is also plotted against  $P_0$ , but each curve corresponds to a given value of  $\gamma$  as indicated.

Values were taken from the express of Fig. 19 to obtain the serves of Fig. 20. In part A the total absorption was plotted against y for the different values of  $P_0$  indicated. Fart 3 was included to denominate the relationship between total absorption and the variable  $wP_0$ . Note: the value of the capational constant was determined in the sum manner as for Fig. 10. All of the points about it part 3 of Fig. 20 were obtained from the curves of Fig. 19; and it is seen that a single curve fits pretty well all of the points corresponding to the different values of  $P_0$ . From this error the following equation was derived and the limits of validity were determined for the 1600 cm<sup>-1</sup> band:

$$\int A(v) dv = 0.025 \left( v R_0 0.50 \right)^{0.75}$$

$$1 < \int A(v) dv < v 0.00^{-5}$$

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Curves similar to those in Figs. 19 and 20 were drawn for the 961  $\rm cm^{-1}$  CO<sub>2</sub> bank in Figs. 21 and 20. It is seen in part 3 of Fig. 32 to: the toing greecytics can be represented by a function of the variable  $\rm vf_{\rm g}^{-0.25}$ ; from this portion of the figure the following equation

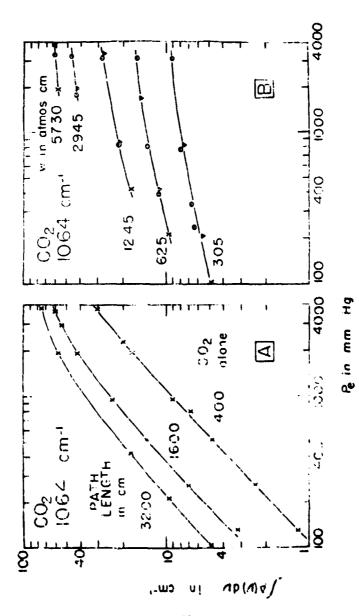
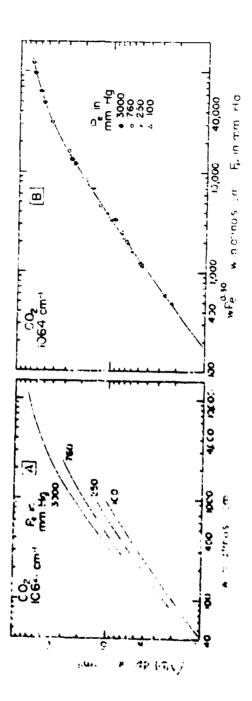


Fig. 19. The total absorption of the 106k cm<sup>-1</sup> CO<sub>2</sub> band versus equivalent pressure. In part / each curve corresponds to a given ecorier path ingth and wife therefore proportional to Pe. In part H each curve corresponds to a given value of w. N'S represent samples if CO<sub>2</sub> and Bo. T'S in part B represent samples of CO<sub>2</sub> and Bo. T'S in part B represent walues taken from curves of part A.



16. 20. The total everytion of the 1064 cm-1 CO2 land remote y in part A: and versus wig 0.30 in part B

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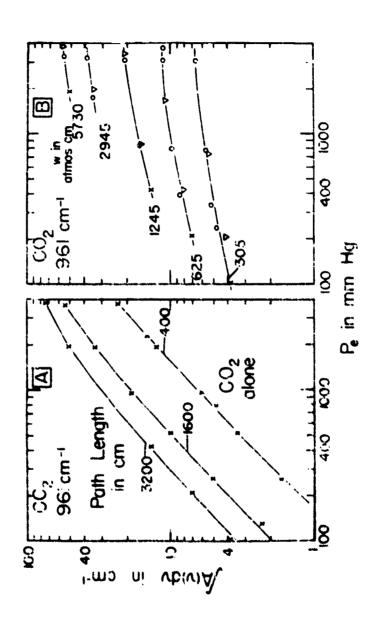


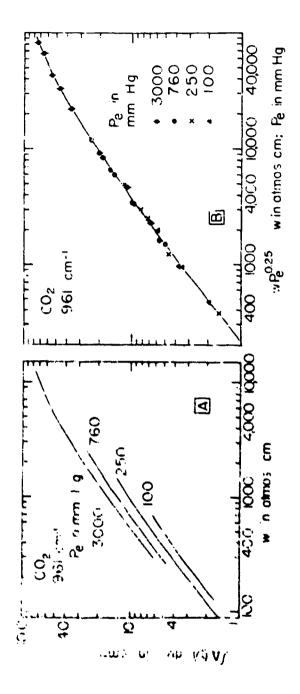
Fig. 2.. The total theorphics of the 961 cm<sup>-1</sup> 302 band when requisions pressure. In part A cach curve corresponds to a given absorption path length and will therefore proportional to re. In part B such curve corresponds to a city of the original of M. Wis represent samples of 302 along to a proportion that is an expense of samples of 302 along the B represent samples of 303 and \$2... Will in part B represent walker taker from curves in part A.

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Pig. 22. The total attorption of the  $961~{\rm cm}^{-1}$  CO<sub>2</sub> band versus  $\frac{1}{2}$  in part A, 414 versus  ${\rm vP}_0^{-0.25}$  in part B.

was derived and its limits of validity were determined for the  $\%1~\text{cm}^{-1}$  CO<sub>2</sub> band:

$$\int A(v)dv = 0.016 \left[ve_0^{0.25}\right]^{0.78} \tag{10}$$

. ...

for

$$1 < \int A(v) dv < 35 \text{ cm}^{-1}$$

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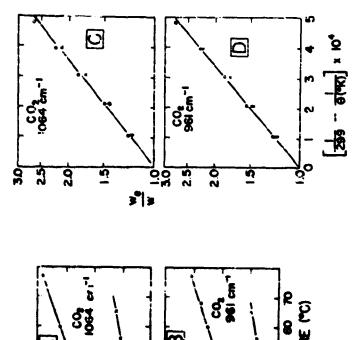
## A. EFFECT OF TENEFERATURE ON TOTAL ABSORPTION

In order to obtain information on the temperature dependence of the total absorption, absorption curves were recorded for three different sets of samples at temperatures above ambient. A thermometer was placed in each end of the cell and a temperature difference between the ends was found to exist. This difference was approximately 1°C for temperatures near ambient and increased to approximately 5°C for temperatures near 75°C. The temperature tabulated in Table 3 is an average of the two to the meanest 0.5°C. It was not possible to used the absorption cell.

Absorption curves corresponding to one of the sets of samples investigated at different temperatures are shown in Fig. 18; and the fractional absorption is seen to increase with temperature at all fraquencies. Values of total absorption for both backs are plotted against temperature in parts A and B of Fig. 25. The curves of parts A and B as well as parts C and D are discussed below in determining a quantitative relationship between total absorption and temperature.

There are several factors which might contribute to the change in total absorption with temperature. One of these factors involves the collision frequency of the absorbing unlecules, which is proportional to the line half width a. The change in a with temperature is believed to contribute only a minor part of the increase in total absorption of the temperature of the two bands under consideration. This boilef is based on the first that the total absorption of the 2550 cm<sup>-1</sup> CU<sub>2</sub> band, as well as other fundamental bands, increases only slightly with unique rure, and one would not expect the half-width of the lines of one CO<sub>2</sub> band to increase more rapidly with temperature than the lines of enother CO<sub>2</sub> band.

indoubtedly, the most important factor contributing to the increase is total absorption with temperature is increase in the population of



The permitter dependence of the total absorption of the 100 and 961 cm. Cop bands. Parts A and B relate the Lotal absorption to temperature for three different snapples. Incre 2 and B relate the different concentration to a quantity based on the calculated change in primarism of the lower vibrational energy latel. The + r. O's and C's in C and D each correspond to the same sample as in A and B. F16. 23.

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the lower vibrational energy state giving rise to the absorption. Another less important factor is the reductribution of the population of the rotational energy levels associated with the lower vibrational state. The would expect that the latter factor would result in a change in the shape of the band, but its effect on the total absorption would be small in comparison with the effect of the change in population of the lower vibrational energy state.

The lower vibrational states of the 1064 cm<sup>-1</sup> band and the 961 cm<sup>-1</sup> band are designated by (02°0) and (10°0), respectively. Their energy levels are 1266 cm<sup>-1</sup> and 1588 cm<sup>-1</sup> above the zero-point energy level (00°0).

The quantity  $r(\theta)$  is defined as the fraction of the molecules of a Boltzmann gas in the energy state  $E_1$  at temperature  $\theta$  (in  ${}^{0}$ K). If  $E_1$  is the energy above the zero-point energy level of one of the lower vibrational states involved in the present investigation, and  $\theta_1$  and  $\theta_2$  are two temperatures near room temperature, it can be shown that the following expression is approximately true:

$$\frac{\mathbf{r}(\theta_2)}{\mathbf{r}(\theta_1)} = \exp \left[ \frac{\mathbf{g}_1}{\mathbf{k}} \left( \frac{1}{\theta_1} - \frac{1}{\theta_2} \right) \right] \tag{11}$$

If one assumes that an increase in the population of the lower energy state by increased temperature produces the same effect on the total absorption as the same relative increase in absorber concentration y at occupant temperature, it follows from Eq. (11) that an effective concentration v<sub>p</sub> can be found by

$$v_{\theta} - v \exp \left[ \frac{E_1}{\kappa} \left( \frac{1}{\theta_1} - \frac{1}{\theta_2} \right) \right],$$
 (12)

where  $v_0 = v$  at temperature  $\theta_1$ . On the basis of Eq. (12) we would expect that experimentally determined values of log  $(v_0/v)$  , only yield a samight line if plutted against  $(1/\theta_1 - 1/\theta_2)$ . Values of  $v_0/v$  were determined from the data and the suggested curves plotted in the following manner.

From the curve in part A of Fig. 25 corresponding to v=650 atmost cm, one cone that a total absorption of 20.0 cm<sup>-1</sup> for the 1064 cm<sup>-1</sup> band corresponds to a \*\*\*recreture of 26°C, and a total absorption of 25.4 cm<sup>-1</sup>

corresponds to a temperature of  $36^{\circ}\mathrm{C}$ . From the curve in Fig. 20 one sees that total absorptions of 20.0 cm<sup>-1</sup> and 23.4 cm<sup>-1</sup> correspond to values of  $vP_0$ . We equal to 8800 and 10,900, respectively. The actual value of  $vP_0$  is not important, but the apparent effective concentration is increased by 10,900/8800 = 1.24. Similar calculations were made for 46, 56, 66, and 76°C and for the curves corresponding to v=1385 and 290 atmss cm. The ratio  $v_0/v$  is based on  $v_0=v$  at 26°C for all curves. The calculated values of  $v_0/v$  for the three different sets of samples were plotted in part C of Fig. 23 in the manner suggested above.

One sees that the points in part C of Fig. 23 form a straight line, within expected experimental error; and it was found that an effective absorber concentration can be calculated by the following equation:

$$v_e = v \exp \left[ 1950 \left( \frac{1}{299} - \frac{1}{6} \right) \right],$$
 (13)

where the factor (1950) was determined from the slope of the line in part C of Fig. 25.

Similar calculations were made for the 961 cm $^{-1}$  band; calculated values of  $w_0/w$  are plotted in part D of Fig. 2). From the slope of the line, it was round that the effective absorber concentration for the 961 cm $^{-1}$  band is given by

$$v_{\rm c} = v \exp \left[ 2040 \left( 1/299 - 1/4 \right) \right].$$
 (14)

The experimentally determined constants, 1950 and 2040, in Eqs. (13) and (14) can be compared with R<sub>1</sub>/k from R<sub>1</sub>. (11). If the energy value 1286 cm<sup>-1</sup>, which corresponds to the 1064 cm<sup>-1</sup> band, is substituted for E<sub>1</sub> in Eq. (11), one obtains 1850 in comparison with the experimentally determined value of 1950. Similarly, if 1388 cm<sup>-1</sup> is substituted in Eq. (11), one obtains 2000 in comparison with the experimentally determined value of 2040 for the 961 cm<sup>-1</sup> band. In view of the assumption made and the possible experimental errors, this agreement is quite good. One then concludes that the dependence of total absorption of these "not bands" on the perimental errors are increase in effective absorber concentration which is proportional to the increase in the population of the lower vibrational statue.

On the basis of this conclusion, one would expect that No. (9) and (10) could be used to predict the total absorption of sample. At temperatures between  $20^{\circ}\text{C}$  and  $70^{\circ}\text{C}$  by substituting for y the value of  $v_0$  from Eqs. (35) or (25) respectively.

### B. COMPARISON OF PRESENT RESULTS WITH THOSE OF KINARDS

Edwards has recently published results of measurements of total absorption of several CO<sub>2</sub> bands at temperatures from ambient up to 1390°K. The samples for which Edwards measured the total absorption of the 1064 and 951 cm<sup>-1</sup> bands contained relatively small values of w; the total absorption for the samples at room temperature was therefore small and the uncertainty was rather large. The equivalent pressures of many of the samples investigated by Edwards were considerably higher than the maximum value of 3000 mm kg used in the present study; therefore, there are only a few of his results which can be compared directly with those of the present study. The total absorptions of the samples which could be compared were in most cases approximately 30 to 405 less than the values predicted for these samples on the basis of the curves in Pigs. 20 and 22. Even this large difference was within the experimental uncertainty since the total absorptions being measured were small.

The ranges of temperatures covered by the two investigations of the 1064 and 961 cm<sup>-1</sup> band were so different that the results could not be compared.

Edvards also investigated the total absorption by the other CO<sub>2</sub> bands covered in the present study. His results for the 675-495 cm<sup>-1</sup> region will be compared to the present results in the next section. The results of the 5716 + 3609 cm<sup>-1</sup> bands and the 2350 cm<sup>-1</sup> band could not be compared since the total absorption of the smallest sample studied by Edvards was approximately equal to that of the largest sample used in the present study.

### VI. EXPERIMENTAL PROULTS FOR THE 575 - 495 mm - SPECIFAL REGION

the 875 - \$95 cm<sup>-1</sup> spectral region was divided into five sub-regions for the purpose of analysis. Tracings of many of the absorption curves are shown in Figs. 25 - \$2; the curves of Fig. 32 correspond to samples of CO<sub>2</sub> alone at various temperatures above ambient, while absorption curves in the other figures correspond to samples man embient temperature. The spectral slit-vidth indicated at various frequencies in Fig. 25 applies to the absorption curves in Figs. 25, ad 26, which were obtained with the Model 99 spectrumster; while the spectral slit-vidth indicated in Fig. 31 applied to the absorption curve in Figs. 27-32 which were obtained with the Model 21 spectrumster.

The data are presented in Table 4, which includes the values of the mean fractional absorption  $\overline{A}(v_1-v_2)$  for each spectral sub-region as well as the total absorption  $\int A(v)dv$  for the satire spectral region 495 - 875 cm  $\frac{1}{r}$  which was in our possible future by other workers.

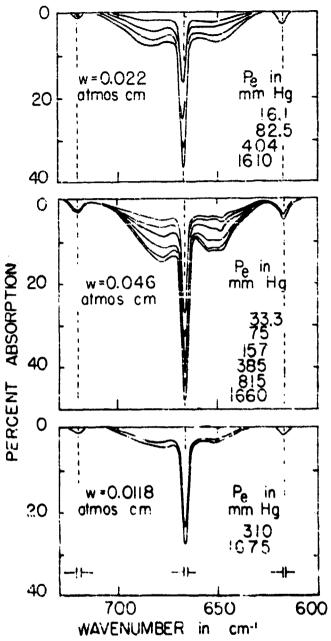
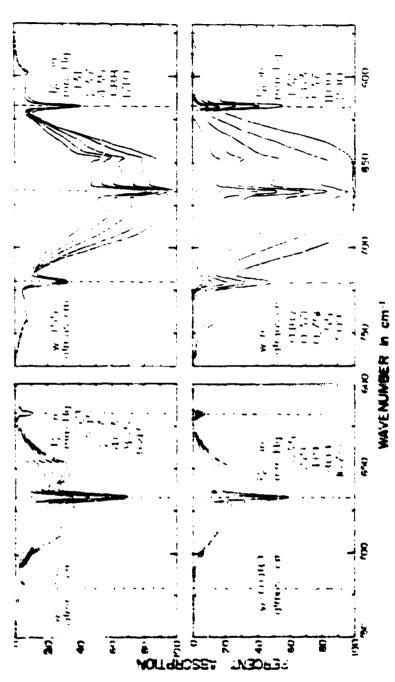


Fig. 24 ption curves of the 8/5-495 cm<sup>-1</sup> COp region. The spectral slit-vidths indicated in the lower portion of the facure apply to the curves in Figs. 84-26.



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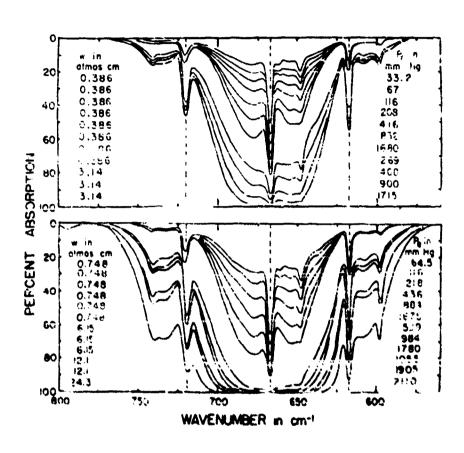


Fig. 96. Atsorption curves of the Milways reff by . Ion.

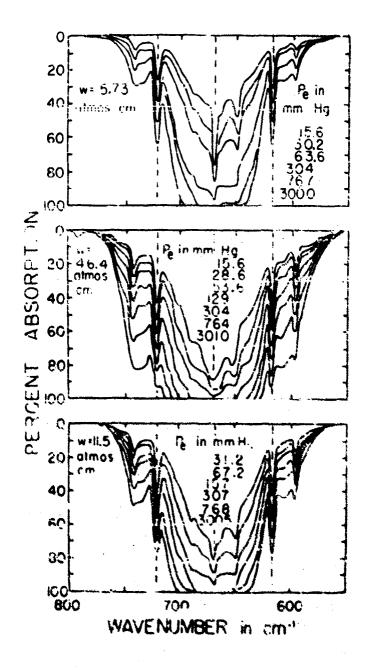
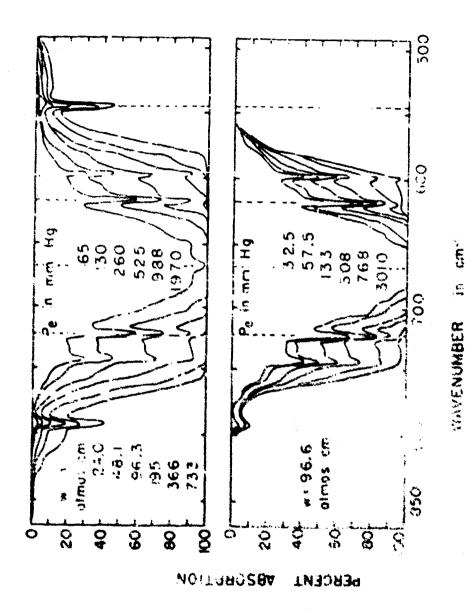


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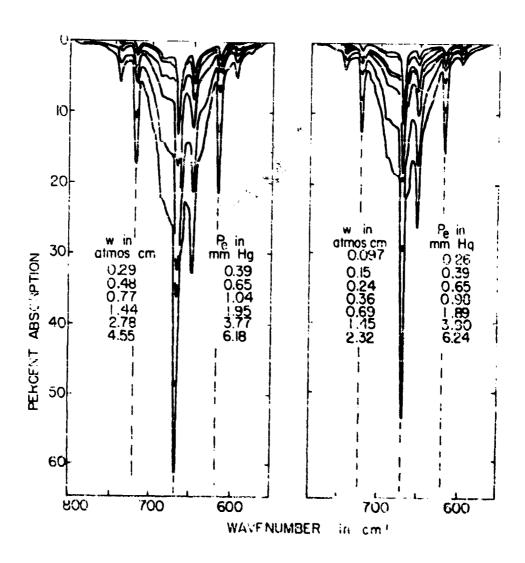


Fig. 29. Absorption curves or the Giray's cur's Cop region.

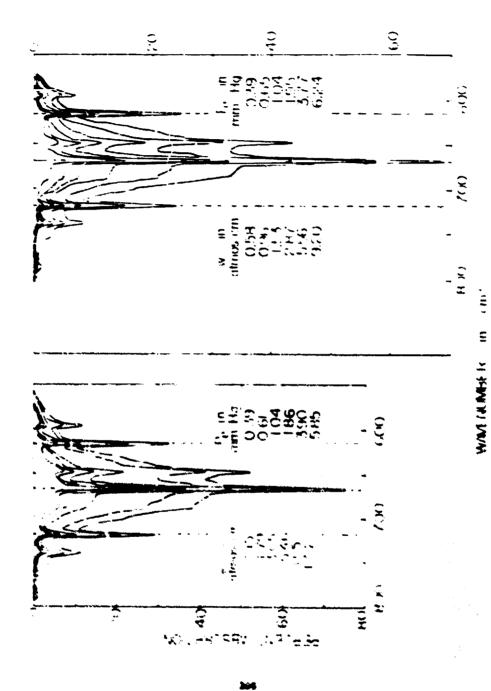
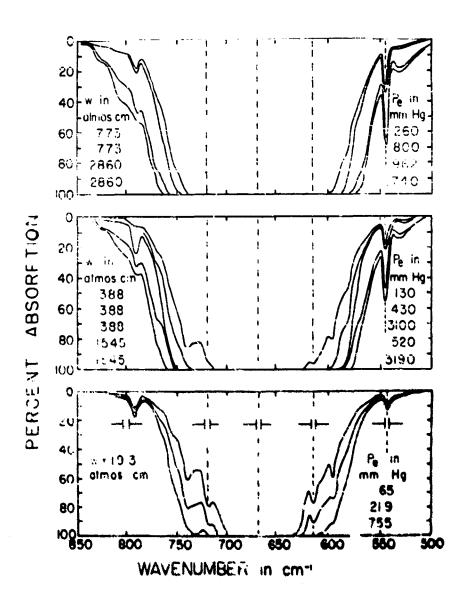


Fig. 27. Abor python nurves of the fifteen in the fitting of the president



Pig. 31. Absorption curves of the 674-495 cm<sup>-1</sup> CO<sub>2</sub> region. The spectral sub-wristed indicated in the later parties of the figure apply to the curves in Figs. 27-39.

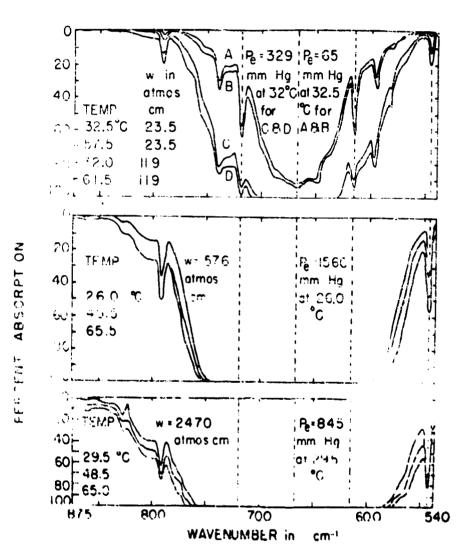


Fig. 32. Absorption curves of CO<sub>2</sub> scapies at elevated temperatures in the 675-695 cm<sup>-1</sup> region.

In the care of samples having small value of w, the major portion of the absorption occurs near the strong fundamental band at 667 cm<sup>-1</sup>.

In Fig. 33 are plotted six curves, each corresponding to a different absorber path length, which relate  $\overline{A(720-875 \text{ cm}^{-1})}$  to  $P_{e}$  for escaples of CO<sub>2</sub> alone. It is noted that the values of  $\overline{A(\nu_1-\nu_2)}$  which are plotted in Figs. 33-45 have been corrected to a temperature of 26°C by use of the curves of Fig. 46, which relate the increase in  $\overline{A(\nu_1-\nu_2)}$  to temperature.

A(720-875 cm-1) was plotted against P, in Fig. 34, where each curve corresponds to a given value of was indicated. The x's correspond to samples of CO2 alone and the fix to samples formed by adding N2 to the COo sample. The triangles on each ourve correspond to values taken from the ourves of Fig. 35; these values were obtained by first calculating for each cell path length the value of Fe which corresponded to the values of y listed on the curves of Fig. 54. The values of A(790-875 cm-1) corresponding to the calculated values of Pe were then determined from Fig. 33 and plotted in Fig. 34. By using the values represented by the triangles in Fig. 34 it was possible to extend many of the curves over considerably wider ranges of values of Pe. The sets of samples with w less than 11.5 atmos on could have been formed by starting with small excunts of CO2 in the Jabu-om cell and adding N2 to obtain the required value of Pg. This procedure was not followed since the initial partire pressure of CO2 for each sample would have been too low for accurate This procedure was not followed since the initial partial measurement. Further error might also arise from adsorption of CO2 on the cell walls; a phenomenon which is more serious at low pressures and which was discussed in some detail in Appendix I of Report 1.

The two curves in Fig. 34 corresponding to v = 5750 and 2945 atmos on represent samples C38-C41, which were investigated by use of the Model 21 with a MaCl prism. The other curves represent samples investigated by use of a KBr prism, with which it was necessary to use larger spectral slitwidths.

The values of  $\overline{A}(720-8/5~cp^{-1})$  for the NaCl data, must of which is not represented in Fig. 3h since it would result in overcrowding of the curves, were compared with values predicted on the basis of the KBr data. In most cases the difference was less than 5% with only a few values difference by as much as 8%. The differences are believed to be due chiefly to experimental errors, not to differences alreing from differing spectral slit-widths.

Values of \$\tilde{A}(720-875 cm^2)\$ were obtained from the curves of Fig. 34 and were plotted in Fig. 35 against a for the different vives of \$P\_0\$ indicated the "manched" curves of \$\tilde{B}(0)\$, 35 are helitived to cont represent the late, and the points which were used to determine these curves were noticed in order that values could be read from the curves curves were it is recommended that \$P\_0\$, 35 and the corresponding figures for the culture spectral sub-regions be used to determine values of \$\tilde{A}(v\_1-v\_2)\$ for more sense.

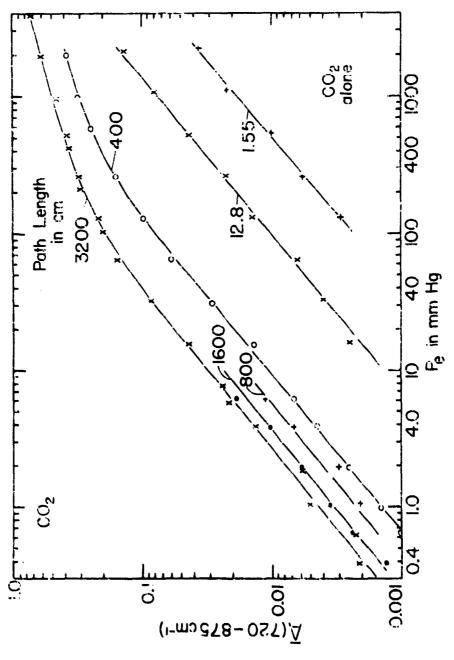


Fig. 35. A(720-875 cm'l) werens Pe for samples of COg alone.

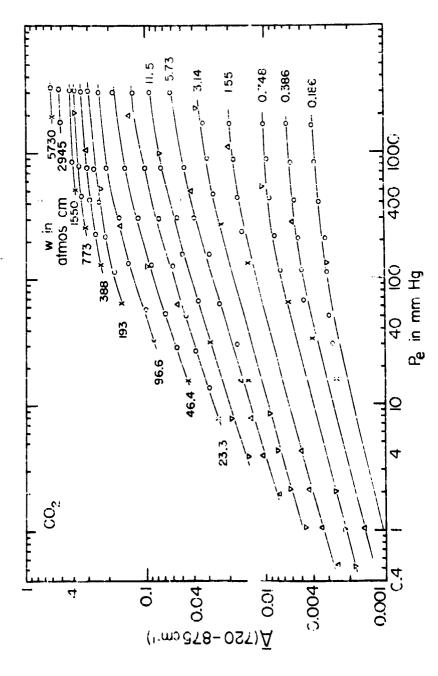


Fig. 34. \(\bar{A}\) (720-375 cm^1) versus Pe for various values of w. X's represent sumples of CO<sub>2</sub> alone; 0's represent samples of CO<sub>2</sub> and M<sub>2</sub>. Triangles represent values taken from curres of Fig. 35.

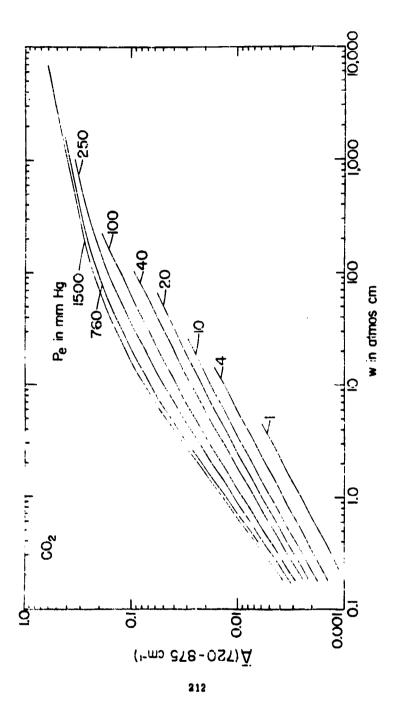


Fig. 35.  $\overline{A}(720-875~\text{cm}^{-1})$  versus  $\underline{\mathbf{y}}$  for various values of  $\mathbf{P}_{\mathbf{e}}$ .

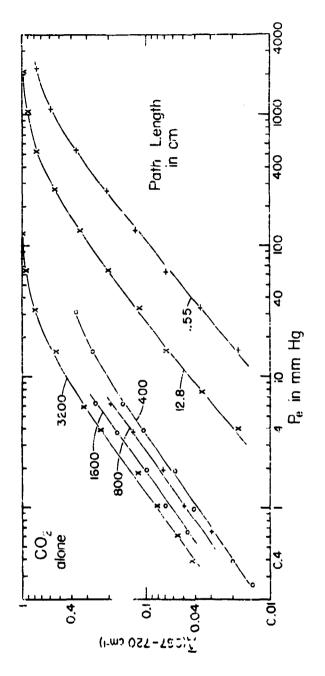
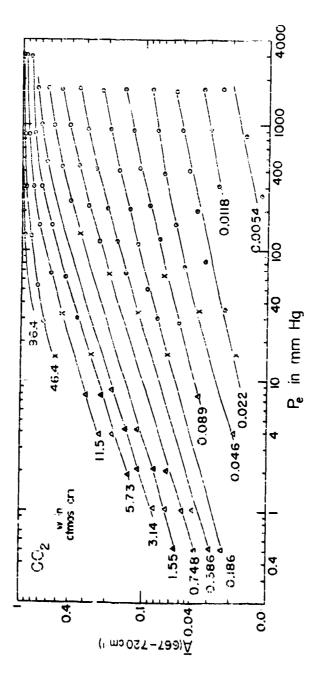
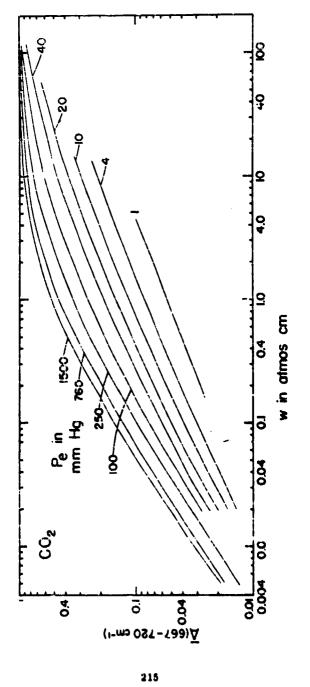


Fig. 36. A 661-720 cm<sup>-1</sup>) wersum Pe for samples of CO<sub>2</sub> alone.



Pig. 37. K(667-720 cm-1) wersim Pe for various values of y.



Pig. 38.  $\tilde{A}(667-720~cm^{-1})$  wersus y for warious wiles of  $P_{\alpha}$ .

Ì

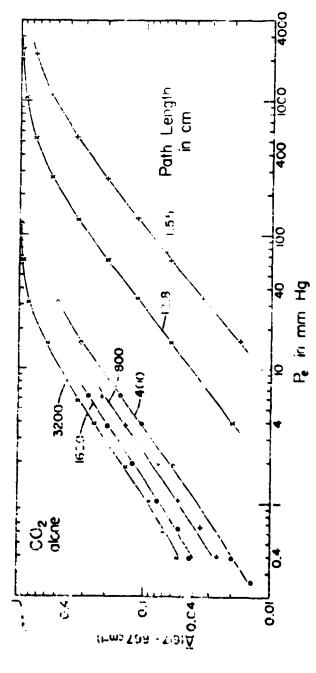
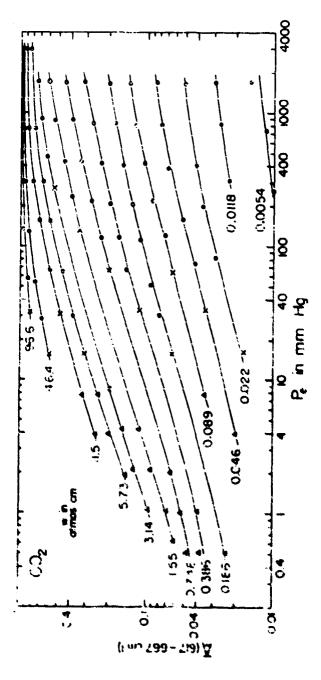
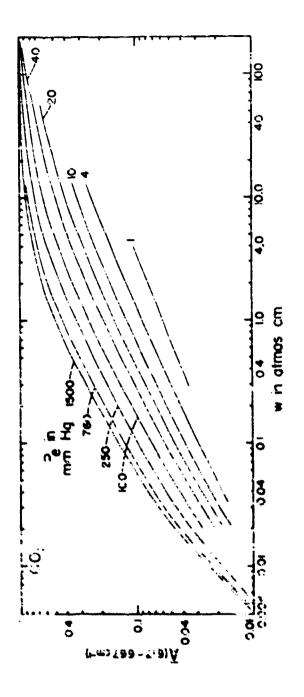


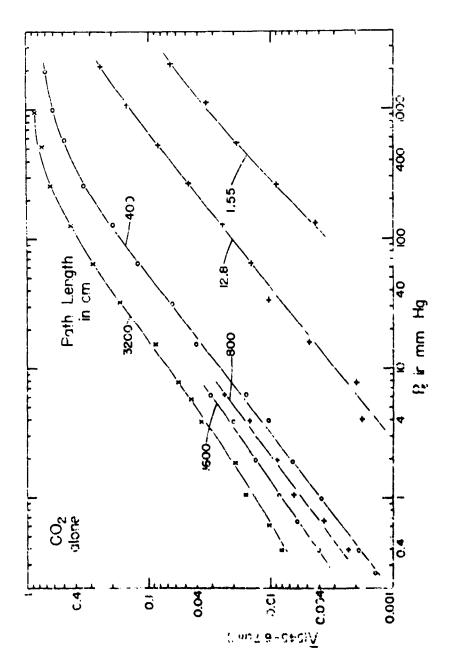
Fig. 39. A(511-567 cm<sup>-1</sup>) versue  $P_{\theta}$  for samples of CO<sub>2</sub> alone.



T.E. b.C. K. 617-657 ts-11 wersus Pg for warlous values of g.



The bir. A(617-667 cmt); wereas y for various values of Pa-



Pig. 32. A(\$45-617 cm<sup>2</sup>) were as for sumples of CO<sub>2</sub> alone.

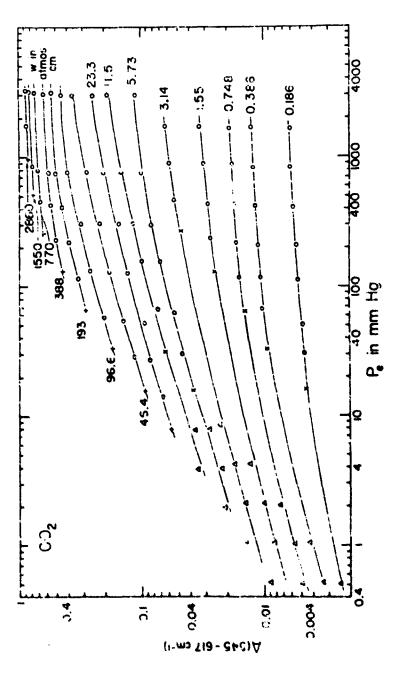


Fig. 35. K(545-617 car.1) versus Pe for natious values of v.

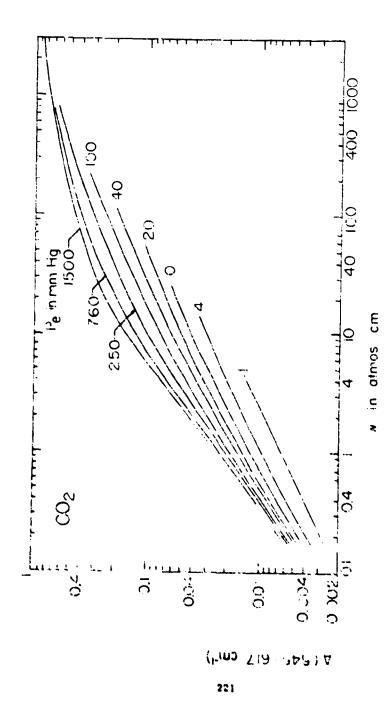


Fig. 44. 4.545-617 cm-1) wersum w for warlous values of Pg.

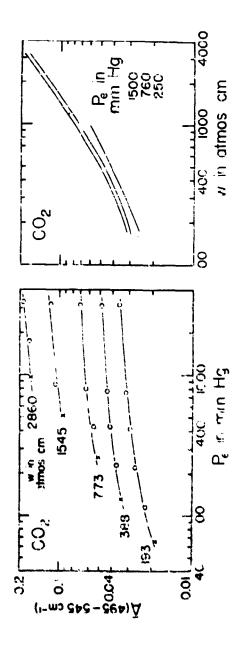


Fig. 45. A 191-545 cm<sup>-1</sup>) versus P<sub>e</sub> for various values of r in the list-hand portion: A(495-545 cm<sup>-1</sup>, versus w for three wiles of P<sub>e</sub> in the right-hand portion.

Curres corresponding to those in Figs. 35, 34, and 35 were drawn for the other spectral sub-regions and are shown in Figs. 36 -  $\frac{1}{2}$ 5. He figure was drawn for the 495 - 545 cm<sup>-1</sup> interval which related  $\Lambda(495$  - 545 cm<sup>-1</sup>) to  $P_{\rm e}$  for samples of  $CO_{\rm o}$  alone, since only the samples of large  $\psi$  produced a measurable amount of absorption.

# A. EFFECT OF TEMPERATURE ON MEAN FRACTIONAL ARRORPTION

Several of the absorption curves obtained for samples at elevated temperatures are shown in Fig. 32; and the results for these samples \$\lambda{lin} = \lambda{lin}, are given in incle 4. Only samples of CO\_ alone were investigated; the pressures of the samples were checked at the various temperatures and were found to increase essentially as for a perfect gas. The absorption curves for the samples at elevated temperatures were found to be erratic in the vicinity of the 495 - 545 cm<sup>-1</sup> sub-region; and these portions of the curves were therefore emitted from Fig. 32. The excessive scatter in the data may have been caused by the water vapor in the spectrometer, whose temperature increased as the absorption cell was heated, or by temperature gradients within the spectrometer. It is also possible that the amplifier gain was not sufficiently high for the recorder pen to respond properly. Regardless of the cause, the results for this spectral sub-region were much less consistent than those for the remainder of the absorption curves, and were emitted in making analyses.

The mean fractional absorption of each of the other four intervals is plotted against temperature in Fig. 46. The mean fractional absorptions of all four intervals increase with temperature. So attempts were made to compare the observed increase in absorption to predictions based on the changes in population of certain energy levels. Two or more bands contribute to the absorption in any of the spectral sub-regions, and a simple claculation such as that made in connection with the 1054 and 961 cm<sup>-1</sup> bands could not be made. Detailed theoretical calculations of the effect of temperature variation on the absorption by CO<sub>2</sub> in the C75 - 495 cm<sup>-1</sup> region have been made by several investigators including Gesamori.

## 3. COMPARISON OF RESULTS WITH PREVIOUS WORK

The HBM report contains the results of measurement of total absorption of all samples of CO<sub>2</sub> and CO<sub>2</sub>-N<sub>2</sub> mixtures in the 575 = .5 cm<sup>-1</sup>

"postured region. Three complex include absorptor constitutions in the range frus approximately 1 to 900 atmos on, and the results can be compared with the results of the present investigation. It is not possible to compare the total absorption or mean fractional absorption of each sub-region since only the total absorption of the entire region is listed

.

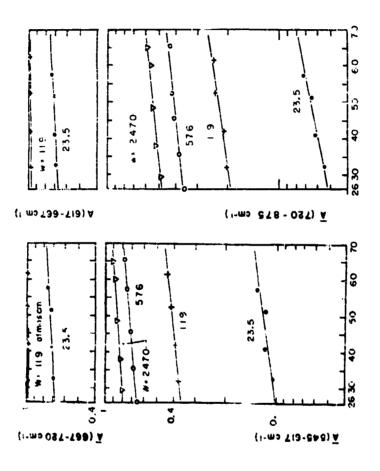


Fig. L5. M(66:-720 cm-1), M(617-667 cm-1), M(720-975 cm-1), and M(545-617 cm-1) were us temperature for four different samples.

TEMPERATURE

in the HBW report. Values of total pressure P and CO<sub>2</sub> partial pressure p from the HBW data were used to calculate the value of equivalent pressure P<sub>0</sub> for each sample by use of Eq. (5). This value of P<sub>0</sub> along with the value of y was used with the curves of Figs. 35, 38, 41, 44, and 45 to predict the mean fractional absorption produced by each sample in each of the five sub-regions. The total absorption of sach sub-region was then calculated from the mean fractional absorption by use of Eq. (1); the total absorption of the entire region was then found by adding up the results for each of the five sub-regions. This value of total absorption, which was predicted for each sample on the basis of the present work, was compared with the value obtained by HBW.

The samples having the smallest absorber concentration showed the greatest deviation between predicted value of total absorption and that observed by HBM. This is a result which one night expect since the CO<sub>2</sub> partial pressure, which is proportional to y, was too low to be measured accurately in the HBM study. Errors in the determination of y as a result of adsorption of CO<sub>2</sub> on the cell walls also contribute greater relative error in the predicted values of total absorption for samples containing a low CO<sub>2</sub> partial pressure (see Appendix I, Report 1). Uncertainty in the position of the zero absorption curve also contributes greater relative error in the measurement of total absorption of samples of small w whose total absorption is small.

Other than the fact that the largest deviations were associated with seasles producing low total absorption, there was no apparent correlation between the deviation and either of the parameters  $\underline{\mathbf{v}}$  or  $P_a$ . The predicted values of total absorption averaged approximately 2.0% less than the observed values; into phenomenon could arise from a number of things, such as systematic errors in sampling, errors in calibrations of the spectrometers, errors in the positioning of the zero-absorption curve, etc.

The root-mean-square deviation was found to be approximately 9%, a result which indicates the experimental error in the MMM work was somewhat larger for the 875 - 495 cm<sup>-1</sup> band (15-micron) than for the other bands.

The data reported by Edwards for the 15-micron CO<sub>2</sub> absorption include 16 samples at ambient temperature, whose values of w and P overlap those covered in the present study. The values of w, P, and n for Edwards' samples were used with the curves presented in this section of this report, in the same ranner as was fone for the HEW data, to predict the total absorption of each sample. With the exception of the sample producing the least total absorption, for which the edicted total absorption was 9.7% greater than the observed value, the deviation between predicted and observed values was less than 5%. The predicted walue averaged approximately 3% greater than the observed value; this result could be pertially explained by the fact that the Edwards results were obtained with samples at 200°K while the present data were obtained

at 299°K. If the predicted values had been corrected to 294°K, the deviation would have been even less; with the exception of the corresponding the maximum deviation would have been approximately 3 or 44, and the average deviation would have been very small. Thus, one concludes that the agreement is extremely good and is well within the estimated experimental error.

The Edwards data and the HBW data provided checks for the samples of the present study which have intermediate values of w; but no data were available to compare with the smallest and largest values. It should also be noted that the minimum value of equivalent measure, as determined by Eq. (3) for the Edwards data, was 385 mm Hg.

### V)1. BAND HELEMEITURS

It was demonstrated in Report 1 that under certain conditions the bend intensity  $\int k(\nu) d\nu$  of an absorption band, where  $k(\nu)$  is the absorber coefficient, can be determined from curves relating the total absorption to absorber concentration. The total absorption for samples for which  $k(\nu)$   $\nu <<1$  for all frequencies, is related to the band intensity by

$$A(v)dv = w / k(v)dv$$
 (15)

The conditions for the validity of Eq. (15) correspond to high pressures and small values of  $\mathbf{v}$ . For sufficiently high pressures and small values of  $\mathbf{v}$ , it follows from Eq. (15) that curves such as those in Fig. 5, for which the log of  $JA(\mathbf{v})d\mathbf{v}$  is plotted against the log of  $\mathbf{v}$ , should be linear with a slope of unity. The value of  $JA(\mathbf{v})d\mathbf{v}$  can then be determined directly from the linear portion of the curves.

In the present study, few, if any, of the samples had pressures sufficiently high and values of v sufficiently small so that the simple linear relationship given by Eq. (15) was completely valid. However, it was possible to extrapolate to smaller values of v the curves corresponding to the largest value of  $P_0$  in Fig. 5 and in the corresponding figures for the other bands. From the extrapolated partions of the only on it was then possible to estimate the values of band intensity  $(\kappa(v)dv)$ . The extrapolated curves were based on the fact that the alopes of the curves increase with decreasing v to a maximum value of unity, and v at all of the curves representing different values of  $P_0$  tend to converge.

It is recolled that values of mean fractional absorption  $A(v_1-v_2)$ , rather than total absorption fA(v)dv, were determined for the sub-regions in the 875-495 cm<sup>-1</sup> region. Therefore, in order to determine the intensity of the ffraction band, it was necessary to determine values.

of  $\overline{A}(v)$ dv from readings of  $\overline{A}(v_1-v_2)$  for the 617 - 667 cm<sup>-1</sup> and 667 - 720 cm<sup>-1</sup> sub-regions from Figs. 41 and 38 respectively. The absorption by the 667 cm<sup>-1</sup> tand is limited to these two sub-regions for values of w and  $P_0$  represented by portions of the curves used to determine the band intensity. Kaplan and Eggers have found that a small portion of the absorption in the two sub-regions given above arises from weak bands other than the fundamental band whose center occurs at 667 cm<sup>-1</sup>. The value of band intensity for the 667 cm<sup>-1</sup> band therefore includes a small contribution not due to the fundamental.

The intensities of the other bands in the 875 - 495 cm<sup>-1</sup> region, whose prominent Q-branches are seen in Figs. 25 - 32, were not determined because of overlapping of the different bands. The intensities of two of these bands have been determined from high-resolution spectra by Kostkowski and Kaplan<sup>9</sup>.

Because of the extrapolation which was involved in determining the values of band intensities, the uncertainty is approximately 20 to 25%, and is somewhat greater than for the results given in Report 1. If the determination of band intensities had been the primary rather than a secondary purpose of the present investigation, higher pressures would have been used and more emphasis would have been put on samples of small w.

The values of band intensity which were determined in the present investigation are listed in Table 5, along with the results of other workers for comparison.

### VIII. SUMMARY

Another get of curves shows the total absorption of each band and the monn fractional absorption of each spectral sub-region for various

equivalent pressures as a function of absorber concentration. These curves provide a means of predicting total absorption or mean fractional absorption for any values of w and  $P_{\rm e}$  included in the wide range of these parameters covered in the study. It was also possible to obtain empirical equations which relate total absorption to w and  $P_{\rm e}$  for certain limited values of the absorption  $\int A(v) dv$  and equivalent pressure  $P_{\rm e}$ .

Absorption curves for the 1064 and 961 cm<sup>-1</sup> bands as well as for the 875 - 495 cm<sup>-1</sup> region, were cotained for different samples of CO<sub>2</sub> at temperatures from faithent to approximately 70°C. The effects of temperature on the obserption were illustrated by the use of graphs. It was found that for the temperature range covered it is possible to predict one increase in total absorption of the 1064 and 961 cm<sup>-1</sup> bands by use of an elementary function which is based on the calculated increase in population of the lower energy level giving rise to the absorption.

Extensive tables which include all the CO<sub>2</sub> absorption data are presented at the end of the report for the use of other investigators who may be interested in an analysis different from that given in the present report.

It is hoped that the present results will be useful to investigators who are inversed in the many types of atmospheric studies involving the absorption and emission of infrared radiation by carbon dioxide.

Table 1. Data for the 3716 and 3609 cm<sup>-1</sup> CO<sub>2</sub> bands

The figure marker corresponds to the figure in which the spectrum of each sample is shown. Absence if figure number indicates spectrum is not shown.

					ja(v)dv ≃=-1		
smple To.	Fig.	v intervences	a. 2	5609 cm <sup>-1</sup> band	3716 2m <sup>-1</sup> bani	3716 - 3609 cm <sup>-1</sup>	Remarks
				1.55 cm Path Longth	th length		
1 4		0.(*)	82	1.5	2.5	4.0	Sample 1; 44.6 mm
A 62		0.(81	151	2.0	2.9	6.7	Mg of COp. Sample
• \ <b>∀</b>		0.(8)	123	7.7	3.1	5.5	2 - 6; No added to
*		0.(81	417	5.6	3.9	6.5	Sample 1.
4		0.(£1	830	2.7	ų, O	9.9	
<b>y</b> 6		0.(81	1555	2 9.	F. 1	6.9	
<b>A</b>		0.:65	117	3.5	4.2	1.1	Sample 7: 89.6 : E
€		0.165	SC:-	, ci	4.9	9.1	of Co. Samples 8 - 11
64		c.165	750	<b>4</b> .6	<b>6.</b> 2	10.8	Me added to Sample 7.
07		0.165	<b>8</b> 52	5.0	٦.٢	7.ZI	N
A ::		0.165	1620	6.0	1.7	13.7	
<b>A</b>		0.687	<b>69</b>	13.6	17.1	30.7	Bearple 12; 375 cm Ho.
<b>C</b> • • • • • • • • • • • • • • • • • • •		c.(87	ź	15.1	20.3	35.3	of CO2. Sample 15-14;
7. 4		0.687	1665	17.6	83	10.1	E2 added to Sample 12.
A 15	•	0.:78	210	7.1	9.5	16.6	Sample 15; 208 🖦 Ag
<b>9</b> 7 <b>V</b>	OI.	0.159	545	13.0	17.9	90.9	of CO2. Sample 16; 417 mm Hg.
11 7	. ~	1.17	1.360	24.3	X.2	\$6.5	of CO2. Sample 17; 808 am Hg
<b>81. ▲</b>	W	2.66	2963	43.8	51.3	6.4	of CO2. Sample 18; 157: see Eg
							۰ <b>د</b> ر <sub>ک</sub> ې

Table 1. (Cont'd)

1					IA(v)dv cm1		
i 、 中点 克		# - 17 mg	F. 80	3604 cm <sup>-1</sup>	3715 cm <sup>-1</sup> band	3716 + 3609 c=-1	Farmatics
				12.8 cm	12.8 cs futb length		
<b>5.</b> ◀	٠.	0.30k	14.2	<u>.</u> :	2.3	4.0	Sample 19: 10:9 mm Br
		0.15	23.6	٠.,	2.6	<b>.</b> . 4	of CO. Sample 20-25:
A 6.3		9.:3.	P. 1.	2.5	<b>3.</b> is	6,0	M. added to Semple 19.
N V	۲.	0.750	50	3.1		7.7	
A 2:		Ç. 1.)	22		5.5	0.6	
4		0.154	FO3	رب <del>ح</del> د		10.5	
₹.		9.14	4	۳.		11.9	
A : e.	^	\$1.0	ner:	5.5	8, 4	13.6	
	~	5	34.	1.2	0.4	7.1	Sample 77: 20.0 mm Ho
1.2	ı	9	4.0	÷.	, · · · ·	- 60	of Che. Sample 28.5%:
ŝ.	, 19	97.0	8	4.7	<b>6.</b> 3	11.0	M. edded to Samole 27.
A 23		97.0	673	5.6	1:1	13.3	
A 3.	i sa	ن. د . بو	419	6.6		16.3	
ii.		د و	1	<b>9.</b> 4	o,	20,4	
A 35	¥	51.5	1623	9.5	4.6	9:23	
4	(V	0.f 75	39.5	6.5	1.8	14.5	Sample 745, 45,7 mm [x
2 <b>4</b>	cu	2.6/2	. 1.5	<b>6</b> 0	0.1	, 9, H,	of 60h. Sample 35-30;
, e	W	2.672	23	11.0	*	4.42	W. added to Burnle 34.
·· ×	~	0.672	X	11.8	7.9	27.3	
40.4	•	0.672	3,20	14.7	6	34.1	
A 53	.,	0.672	15.5	15.7	5.5	6.6	

Table 1. (Comt'd)

					A(*)45. Car.1		
7 3	žė	* comp	7. A	1. E 63.	\$716 car 1 man	5716 ÷ 3639 cm-1	Rearts
			~1	2.6 ca 7sts	12.6 cm Noth Longth (Cont's)	7	
3	ŗ	<u>\$</u> :	2	<b>12.3</b>	34.6	% •	Semile to: or a man
14 4	*	1.75	Z.	15.1	10.4	**	OF CO. Sept. 13-14
;; ∢	~	3.	Š	10.6	٠. ئ	9.03	E. added to famile ho
~ ·	~	<b>X</b> :	ij	22.5	20°.	\$0.0	· · · · · · · · · · · · · · · · · · ·
\$ ~	^	×.	2	27.0	りま	61.0	
54.4	^	÷	×	8.0	9.0	31.8	Section 45: 107 mm 12
<b>3</b>	~	8.	4	r.	<b>%</b>	50.7	of CO. Samia Maria
<b>5</b>	•	ð	933	9	1,54	62.6	No sedded to temule 45.
3	^	8	22.	o. 1	8.3	8.5	
2	^	 8	\$28	9	9	8	
<b>9</b> .	.~	8	8	2.1	8	118.4	20 CO - 1
7 21	~	<del>2</del> .	1770	61.8	6.6	127.1	To added to Sample 49.
X 4	^	13.7	1070	75.0	70.5	146.4	Semale C: Role II II
\$	^	<b>1</b> .2	oý gr	61.1	7.7	155.0	of CO <sub>2</sub> . Semple 53;
\$ <		21.4	80%	4.16	9779	179	tions the 145 at 1890 and 186
							٠٠٠ دوم.

"Mile 2. Buts for the 2550 car 2 CO2 Bank

Mark also option curre as alwished at the head contert, 2950 cm<sup>-1</sup>. The values of total absorption of sail, while ware describing as and are tabulated below along with the total absorption of the sastir, hear. The figure is without a white the apericum of each sample is above. Also record of figure expected to apericum of each sample.

	Nesa rika		Sample 1; 2:9 mm Hg of 33;	Sample 2-9; No added to	Sample 1.	•						Sample 10: 6 m Me of CO.	Newple 11-16; No mided	to Sumple 10.	•							CO. Sample 29-25: E. added	3				
Zaci ra	Band		1.35	1.80	2.40	2.5	4.	ن <b>ن</b>	7.6	0.0	16.5	5.0	.e	<b>8</b> .0	(i)	8.F	1.1	15.35	6.	8.3	1.,	9.6	š	14.2	24.1	8	3.
À	> 2150 ce.1	1.5' on 18th Longth	2.3	0.75	8.7	1.15	1.9	2.5			× 0.	<b>5</b> .		0	ej.	£	2.5	45	*	10.15	0.	<u>د</u> د		Š	10.1	2.5	6.9
yb(v)42	1350 tm. > 2550 cm.1	37	8	6.1	9	1.75					3.5	8	7.6	÷.	*	۶.0	6.3	2.6	10.	2.2	-	1.75	5.73	 	13.0	15.9	£.
			3.6	13.1	3.	9.0	\$	ž	3	2	\$191	<b>61</b>	15.2	9.6	\$2.5	<u>5</u>	Ş	ŝ	Ž,	15.73	<b>5.</b>	21.12	53	ē	Š	Ç.	£73
y	4. THOS CO		C. 00/3	5 CO 13	C 0752	0.0732	J. C. 7.2	0.037	2000	J. C. X.2	5 C. C.	6.0.C	60.0.0	C. C.C.	6.0.0	5.00		6,10°0	0. AC	0 01.53	8	1100		0.0	0.0	0.037	0.423
3.5												Į.		ù		2		2	ï.					::	. : 	4	-
1	2			r <b>4</b>	•		,** &	•		•	<b>.</b>	3	. d	<u>"</u>	?	4.	\$1.4	3	17	<b>.</b>	2	Ö,	3.5			23	*

Inble 2. (Cont'd)

3	3,6	,	•	/A( v )&v	ķ		
Ö	2	atmos ca	Y E	< 2350 cm-1	> 2350 cm <sup>-1</sup>		Bearts
				L.55 CO Pat	1.55 on Web langth (Cont's	<b>9</b>	-
\$2.4		0. <u>17</u> C	027	26.3	16.9	45.1	S. W. S. W. S. W. B. W. C.
12 Q		<b>0</b> . 17c	82	S.	21.0	53.8	2
Ã.		D. 17C	ij	63.0		66.5	adderi to Bample 25.
23		0. 1C	655	13.7	99	72.6	
٥ <u>٢</u>		o. 17°C	1700	¥7.4	30.9	76.3	
12.4			4.2	8.0	0.50	1.35	31: 3.2 mg
2	#		7.9	1.80	1.25	3	ģ
5.	1		15.7	8	00.0	יא	13. 12 ]
J.	۱:4	0.0455	ĸ.		8	12.5	200 to 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1
2	ند		3,	15.4	0	25.3	, <u>;</u> ;
X:	7.7		177	8	18.9	18.3	10, 10,
F 2	=		E.	**	27.6	73.8	77: 211 mm Ex
<b>9</b>	7		ない	3	6.0	8.0	38: 411 BR
2	:1		.055	<b>*</b> •••	%	100.1	39: 812 <b>18</b>
ş	<b>1</b>		<b>5</b> 172	87.8	0.0 <b>4</b>	127.3	<b>6</b> 0
				2.6	12.6 to feth length		
111	9	ñ	×.0	3.1	1.75	<b>.8</b>	Sample \$1; 2:9 mm Te of CO
21 6		ŝ	1.1	·.s	5.0	7.40	beato: In added
514	23	5	3.5	6.5	8	9.01	Semple 41.
:		ń	<b>2</b> 9.9	8.2	5.4.5	15.7	
<b>5</b> 1	2	ŝ	5	; ;	. <del>2</del>	80.58	
.s.		8	218	15.4	n.5,	<b>%</b> 5.0	
·	2	0.0	E C	9 61	14.5	1.1	
	2	ŝ	910	23.5	17.75	41.2	
7		ş	1615	18	20.7	43.7	

Tatle 2. [Cont'd]

				val v lat	ħ.		
36.27	ار ن :	> 9 20 20 30 30 30 30 30 30 30 30 30 30 30 30 30	.9	< 37.0 cm <sup>-1</sup>	> 2350 cm <sup>-1</sup>	atire Brd	Rema. As
				12.6 cm P42	12.6 on Park Longth Cont'd	ন	
e. Ĉ	, <b>e</b> è	<b>8</b>	4.4		5.85	. t Uh	Sample 50; 5.7 mm Hg of 30;
45.00		*	<b>4</b>		4.65	9	Sample 51-58; M. added to
Ø.	14	<b>9</b> 6.3			5.5%	I. 5	Cample 50.
**			1,2		8.95	25.1	
读出	\·\$	8	91		11.6	26.9	
55 6		5.0g/	8		15.5	7.15	
*	^!	<b>3</b>	572		18.8	45.9	
ji.			07.61		23.6	55. k	
农		<b>38</b>	1,50	5.5	27.5		
\$	и	<b>39</b> 7 C	14.6		9	<b>38.</b> 6	Searche 59: 11.2 am Na of
25 8	¥	97.	2.45		0	23.5	<b>*</b>
19 1	31	57.0	3		11.6	10.2	to Sample 59.
į.	."1	0.14.0	<b>ئ</b> رن		15.1	6	
<b>8</b> 63	2	0.169	170		6.91		
3	?	9	8		25.9	60.3	
	3	C 165	69		27.5	3	
Š		0.169	15%		30.8	73.3	
19 6	;	-	Ž.	2.4	12.9	35.5	67: 25 mm Hz of
*. Z	** **	3.0	× ×			5	Same le 68-73; No added to
(÷ =	 ••		21.		21.7	57.6	Same Le 67.
£ 70	 	<b>1</b>	251		8.9	7:5	•
11.4	1	まい	Ŷ		8.6	81.4	
۳ ت	<b>;</b> ;	せんし	ĝ		£.9	99.6	
ا		まい	, 889		35.1	92.9	

Mails 2. 'Cont'd)

	Remark s		Sample 74; 45.5 = Mg of	Sample 75-79; No	added to Bample 74.	1			80: 85 mm Ke of	Semple 81-84: No	added to Remale 80.			e 85: 168 mm He of	COr. Herrile 86-88: In		•	Sample (99: 75) and Mr of	Semple 90.91: IL	added to Sample 89.	20 St. 7kt 18 Es of	CO2. Sample 97; H2 caded	and the page.	
			Semp)	8	<b>646</b> 4				ileme):	Ş	adden.			See L	ą			See	ģ	4	Herm	8	3	
	Pettre	ਰ	89.	73.7	<b>\$</b>	Š	8	81	8.5	· o	8	901	111	101	i	119	ź	ķ	2	<b>3</b>	147	125		
t,	> 1550 cm <sup>-1</sup>	2.8 cm Puth Longith (Cont's	2	<b>8</b>	¥.7	2.2	35.5	7.7	*0.	33.5	さま	35.	57.2	34.5	27.0	<b>8</b> 9	£2.2	40.3	+1.5	8. 8.	45.6	1.54		
va( v ) a l	< 2550 cm" > 1550 cm" 1	2.8 cm Nth	36.6	<b>1</b> 6.6	58	3.03	3.0	4.8	45.4		Š	70.	13.1	71.4	, A	29.5	85.6	96.1	30.5	90	861	100		
	. ž		<b>%</b>		210	'n	<b>1</b>	1990	Ħ	7	535	ŝ	386	216	\$	157	1470	9.	913	1680	Š	1780		
	AC souts		6.5.0	0.619	645.0	679.0	\$ .O	0.679	8	2	8:	8:	8	\$		2.55	2.55	8.	8	8.	1:.2	1:.2		
	7. 0.		1	1	#	7	#	7	2	3	2	2	2	22	2	2	ជ	я	`	2	2	2		
	No.		2	* 73	22	1.4	£	<b>61 a</b>	8	10	8	6	á	¥	, es	10	<b>98</b>	3	9	5.	9	<u>ب</u>		

-- -- -

Their 5. Into for the 1054 and 961  $cm^{-1}~{\rm Ch}_2$  Bands

The figure number corresponds to the figure in which the spectrum of each sample is shown. Abstnor-

				•	(1-1) (ca-1)	(cm-1)	
1	ż	4	(etmos cm) (mm list;)	( <b>**</b> ( <b>**</b> ( <b>**</b> ( <b>*</b> ( <b>**</b> ( <b>*</b> ( <b>**</b> ( <b>*</b>	1064 cm <sup>-1</sup>	951 cm <sup>-1</sup>	Remrks
				red is 1500 =	• 1600 =		
(	37	88	84	3.9	5.2 0.7.	<b>6</b> 1 10	Samples 1-7; 302 alone. Eressure of (0, equals
w =	¥ %	<b>X</b>	<b>3</b>	620	15.7	5.0	F./1.5 for all samples
\.# • U	2.1	0.93	1615	8	24.1	:8.8	of 302 alone.
, e	K	98	26%	526	9.74	33.5	
. <b>19</b>	2	26.0	\$141	0.6s	され	1.9	
-	3	<b>%</b> 0.	5395	<b>16</b> 55	9.09	<b>5</b> 4.10	
				Path Le ath .	* 400 cm		
•	4	2	9	9,1	1.2	7.0	Samples 6-12: CO2
	3.2	20,00	8	3,8	2.4	1.7	alone.
	2	X	8	950	æ. •	3.4	
	4	8	75	35	o. <b>6</b>	0.9	
2	l	<b>%</b>	21.	1935	17.5	12.5	
		8	8	733		5.1	Sample 17; 610 nm Eg
		1	8	•	9.1	9.9	of CO2. Semples 14-16;
		78	8		10.8	ю 0	Sample 15 beated to
<b>19</b>		56.5	8		<b>1</b> 5.	6. 6.	indicated temp.
		3	35.	57.55	29.8	23.7	Sample 17; 2875 vm Eg
- es			9	``	33.0	25.4	of CO2. Sample: 18-19;
20			1,00		8.0	89.0	Sample 17 ben'ed.

Table 5. (cont'd)

					I and a Marian	(car-1)	
	e di N	i i	( otens ma)	, E	1064 cm <sup>-1</sup>	961 cm-1	Pesarks
				Nts Jegst	Path Length - 400 on (Cont'd)	£'4)	
2 2	£ <b>5</b>	<b>5,5</b>	22	22/00	19.0 26.6	15.7	Sample 20; 1755 mm Hg of CO <sub>2</sub> . Samples 21-25;
3 2 2 E	<b>9589</b>	36.00 1.00 1.00 1.00	&&&&		23.52 6.00 e	21.0 25.2 27.5 51.0	Sample 20 betted.
			Æ!	Pets Leagth -	20.00		
<i>1</i> 8¢	1.1	25.0	\$ \$	67.7	6.÷	F. 9.	Sample 26; 79.5 mm Hg of 30. Brunden 27.30;
:889	H 11		ere)	,×25 ,×25		.v.v.ø .4.6.iv	M2 added to Sample 26.
ನ೩೭೩	777	%%%% %%%% %	2222	4 <b>88</b> 5	9.5 11.2 13.5 5.6	6.88 8.48 7.00	Sample 31; 163 mm Hg of CO <sub>2</sub> , Earples \$2-34; m <sub>2</sub> added to Sample 31.
288	222	488 266	५५टा ५५टा	2 <b>28</b> 52	21.2 21.2 2 <b>6</b> .0	13.3 15.9 20.4	Sample 55; 324 mm Hg of Co. Samples 56-37; Mg elifed to Sample 55.

Table 3. (Cont'd)

Sample	Ġ	i i	>	ù.	(A(v)dv (cs-1)	(c=1)	
	3	C. 00.6	(atmos cm)	(E)	1064 cm <sup>-1</sup>	×61 cm	Remrks
			2	Sength = 32	Path Length = 3203 cm (Conl.'d)	7	
Ж.Ж 00	<b>57 17</b>	% % .0 .0	295.5 24.5 5 4.5	1780 3240	1.6.3	34.5 36.1	Samples 38-39, No added to 763 am Hg of 60s,
33	11.11	%.5.°°	3730 3730	056 63K 1	57.2 61.1	50.9 55.1	Sample 40; 1495 sm Mg of CCp. Sarpie 41; Mp added to Sarpie 40;
<b>3</b> )	¥	9. 	002/17	3,800	75.3	72.9	Sample 42; 292 mm Hg

hable ... have for 80, absorption in the 815-495 cm<sup>2</sup> Region

semajes 100 were tarritigated by use of the semanted Notel 99 spectroseter; all others were staticd with the Notel 21. The figure number corresponds to the figure to against the spectrum of each sample is shown. Absence of figure number indicates spectrum is not above. Absence of a wise 1 indicates last this quantity was to small to measure accurately or that the Foult was delated because of some collines of figure indicates the transform to the critic spectral region [15-55] cm<sup>-1</sup>.

130 000 47		1,100	S	3	2000	100	12-22	" of JA what correspons to the entire specific tolden cly-by care.		
		į	,	•		, JA	A(v1-v2)		f 875 Af y blay	
•			(*12 mm)	(%)	720-375 ce ·1	%;-720 cm <sup>-1</sup>	617-667 cs	ç45-617 cm⁻¹	646	Remarks
1	<b>!</b>					Poth Singth . 1.55 cm	1.55 cm			
144		:.55	# # # # # # # # # # # # # # # # # # #	223		1.1 x 10 2 1.5 2.6	1 = 10 <sup>-2</sup> 1.5 1.5		2.1.08 1.40 18.08	Same. 1-3; 3.) = Hg of CO <sub>2</sub> plum H <sub>2</sub> enided.
* ~~	`	11 20 11 10 10 10 10 10 10 10 10 10 10 10 10 10 1	0.0016 0.0018 0.316	38£		0 mm 0 mm	2.5 3.0 5.0		9.9.v 8.4.v	Same, 4-6; 6.5 mm Mg of COp plus He anded.
3:33	* * 4 * 4			a <b>v</b> 888888		ရာလေလက္ခရာရာ က်လွှဲကိုက်သော လိုလ်	- 4 0 0 may 4 v		1.9 4.4.4.4.4 8.1.6 8.2.2.2.2	San. 7; 12.4 im Pg of 30; Seva. 9. 3; Hg added to San. 7.
<b>はずれまない</b> 日本ではない	21233 <b>8</b>	\$100 W 147	9999999 80666 80046	852888 2		7.7.4.8.8.10.0.0.0.0.0.0.0.0.0.0.0.0.0.0.0.0.	5.5 × 10-2 5.6 7.0 7.0 7.3 6.5		2.4.6.6.6.6.6.6.6.6.6.6.6.6.6.6.6.6.6.6.	Seas. 14, 25.6 um (6 of CO <sub>2</sub> Same. 15-19; 32 midded to Sam. 14.
2222242 44444	מממג ג	<u>ก็สิ่มสัมม</u>	######################################	65 : 219 219 605 605 605		केट खाले. कर्ड बंद स्ट्रेस	2.4.2.5.2.5.2.5.2.5.2.5.5.5.5.5.5.5.5.5.		9.98 9.88 11.6 4.51	San 20; 142.6 rm 145 of CO <sub>2</sub> . San 21-25; 152 added to San 20.
*****	********	द्रश्चित्रदर्श	45528 45528	55 25 25 25 25 25 25 25 25 25 25 25 25 2	3.5; # 10°2 0.53 1.1: 2.3 7.7	76.95 1.69.7 1.69.5 1.69.5	. 1. 5. 5. 5. 5. 5. 5. 5. 5. 5. 5. 5. 5. 5.	0.51 x 10 <sup>-2</sup> 0.50 2.1 4.2 7.0	8888	Samme. 26-30; 32 &lone.
						Part. 278gg = 12.8 gg	8:33 31			
	T 4 XXXX	**********	<b>2000年 100円 10</b> 0円 100円 100円 100円 100円 100円 100	2, 25, 25, 25, 25, 25, 25, 25, 25, 25, 2	200 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 × 107 ×	6.8 x 10° 6 7.8 x 10° 6 15.5 x 10° 6 15.4 x 10° 6 10° 4 x 10° 6 10° 8 x 10° 8 10° 8 x 10° 8 x 10° 8 10° 8 x 10°	6.2 x 10-2 8.0 9.4 19.4 13.1 15.5 16.1 20.2	0.49 x 10 <sup>-2</sup> 0.54 x 0.55 0.55 c c c c c c c c c c c c c c c c c c	3.75 8.90 8.90 8.51 8.55 1.55 1.55	Sam. (c), 1912 and Eng of COp. (c)anc. (20-45) Mg added 10 Sam. (3).

								A(v1-v2)			648	
	<u>.</u>	4 <u>2</u>		(etax	(Im Mg)		667-720 cm <sup>-1</sup>	17-667 cm-1	545-617 cs1	495-545 cm-1	Jugs A(v)dv	Reun rics
	i	! !	!				7 15	ength = 12.8 (c	ont '£.)			
		قد ٿا،	: , ,	7 d	53.2	0.42 × 10 <sup>2</sup> 0.48	11.2 x 10-2 14.5	11.3 × 10-2 14.5	1.1 x 10-2 1.1		55.0 4.53	(Am. 39) 25.5 rm Hg c
	.:	<b>.6</b> 3	, id 5	5) c) 第第	¥1.50	5.70	20.3	16.7 19.9	4		0.4.00	* 3 ded to Sem. 59.
	· .	'YR'	á	6	iş,	, o, i	25.0	33.1			90.0	
X         0.74         4.5         0.66 x 10 <sup>2</sup> 13.9 x 10 <sup>2</sup> 1.5 x 10 <sup>2</sup> 2.5 x 10 <sup>2</sup>	;;	EF.	z v	9 4 4 4 4 4 4	X SS	3 <b>6</b>	 	1.89.9X	111		35.6	
	ž	*8	×	9. T.	\$.5	0.66 x 10-2	19.9 x 10%	19.9 1 10-2	1.5 2 13-2		9.7	Sam 46; 49.6 an Ag of
	<u> </u>	X1 Y	, 1, X, 1	φ <b>σ</b> (-) (-)	3 E	3. d	÷.	4.5.4	۲.۲		3. 5. 2. 5.	102. 30ms 17.51; M2
	įį	i XI	XX	9	ķ	D: 0	** **	53.2	1.8		37.5	
1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5    1.5	22	XI XI	ಸಬ	0.7.6 8.7.0	Q Ž	 	41.2 1.3.3	 <b>.</b>	2.0 2.0		<b>2</b> €	
			ĸ	15.1	161	1.5 3 10-2	34.2 x 10-2	¥. × 10-2	2.5 x 1.3"2		3.5.5	Sam. 52; 103 mm Mg of
10   10   10   10   10   10   10   10	. S.		ķ	,	7	- u - i	4.0	9.	80.0		1,5	CC2. Sam: 53-55; da
15   170   2.7   1.0   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1   5.1			r s	i	÷	1.1	2. 5.		מי ויי סי מי		Ж. 6	added to San. 52.
35.         36.         3.14         369         2.1 x 10 <sup>-2</sup> 35.1 x 10 <sup>-2</sup> 4.9 x 10 <sup>-2</sup> 4.9 x 10 <sup>-2</sup> 4.9 x 10 <sup>-2</sup> 4.1 x 10 <sup>-2</sup> 36.1 x 10 <sup>-2</sup> 4.1 x 10 <sup>-2</sup> 36.1 x 10 <sup>-2</sup> 4.1 x 10 <sup>-2</sup> 36.1 x 10 <sup>-2</sup> 36.2 x 10 <sup>-2</sup>	22				12.		4.6	50.1			89. 69. 6	
25	24	<b>19</b> 1	<b>2</b>	3.1	*	2.7 × 10.5	55.1 x 10°	55.6 × 10°?	4.9 × 2072		3.5	34 m. 57; 207 ms Hg of
25	22.2	R Xi X	R.R.		88	ų eni Nani	. e.		, o		. 8. 8. 6. 6. 6.	added to Sam. 57.
25	ŝ	æ	o,	ă A	1115		· ·				96.76	
25 35 12.1 1555 8.E 10 <sup>2</sup> 93.6 x 10 <sup>2</sup> 16.5 x 10 <sup>2</sup> 16.5 x 10 <sup>2</sup> 18.5	₩ ₩ ₩	X.#3	ลืดต	9999 9999	.8 <b>4</b> 8	4.7 ± 10°2 5.5 5.6	76.4 x 10°? 31.8 83				163.0	Stem. 61; 4C7 m: Hg of Chy. Sams. 62-55; No added to Sam. 61,
75 5.254	1 0.0	55 % 55 %	z.r.	12.1	1905	8.E + 10-2 9.7	93.6 x 10°2 92.4		16.3 x 10°2 19.9		118	Sam. 54; 811 um Bg of
75 2.5 24 25 25 25 25 25 25 25 25 25 25 25 25 25	;		:		,							
27 27 27 27 27 27 27 27 27 27 27 27 27 2	\$5.5	ì	۲۲	e di di di di di j	7.7		5.5 5.5 5.5	5 * * * * * * * *			 	
27 27 27 24 25 2.5 1.1 2.10 <sup>-1</sup> 36.6 x 10 <sup>-</sup> 35.8 x 10 <sup>-2</sup> 3.6 x 10 <sup>-</sup> 3.5 x 1	š n	3	~,	ŝ			0-1	bih Length = 40	C + 17 C		3	
21 21 5.7 5.9 5.2 4.7 5.5 5.9 5.1 4.7 5.5 5.9 5.1 4.7 5.1 5.0 5.1 5.1 5.0 5.1 5.1 5.1 5.1 5.1 5.1 5.1 5.1 5.1 5.1	65 6		2		3.51	7.4	-C1 4	31.1 × 10-2			5.45	Sam. 69; 32.0 an Eg o
27.5 5.7 1.56 5.7 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0 6.57 5.0	2 Z		3.3		X X X X X		55.5	5, 73, 61 t.	1.4 2.3		4. P.	CO. Sames 7C-75; No same 69.
2	2 E		2. <b>18</b>	~ <u>}</u>	ደይ		L	1. v.	0 m		 	
	2		18.4	, A.	5		73.5	97.6	10.0		 183	

	Peaksilas		Sem. 75; 24.0 m; fig of	added to Sas. 76.			3ams. 82-94; CO2 alora.										Sens. 95-105; CO., alone.										Ham. 197; 9.0 mm Eg of Hy added to cell, then 190, added to term measure	of 12.0 ms lig. He was adde	of 502 on cell wells. Sees 108-113; My addid to Sen.	107.
675	J <sub>LS5</sub> A( v)dv		5.92	6.79	33	r)	# 8 	8	4.45	N N N N		97.6	13 151	176	2 <u>3</u>		2.66	2 8 2 70	8.50	*` **		7.3	, eo (%)	13.1	, e			ğ	ខ្មែន	
	495-545 cm-1														5.1 x 10° 7.1						•									
	545-617 cm.1	Cont'd)	6.4 × 10-2	0,	1.4	-0.0	0.1 x 10°2	,	410	37.	9		35.0		70.1	51	0.26 x 10-2	<b>83</b>	0.78	2.4	8	0.41 x 10-2	96.	. 0	3.1	8	4.5 # 10-2 5.9	-0,5 6,4	6.69 1.5.55	
1(41-42)	617-66? cm <sup>-1</sup>	Rth lergth = 400 cm (Cont'd	43.6 x 10-2 53.4		87.5		1.4 x 10 <sup>-2</sup>	9		10.3	,	x 10-5			38	Path Length - 500 cm	5.6	v v v	1.6	21.4	Path length = 1600	¥ 10-2	7.1			Peth langth - 3200		1012	}	
	667-720 cm-1	1 414	#6.4 = 10°2	11:1			1.2 x 10 <sup>-2</sup>	) + \ 	9.0	. Q. =		71.1 ± 10-2	35.2	8	88	ĕ	2.1 x 10-2	. o.	7.2 12.6	19.5	Ä	2.3 x 10-2	o, o,	9.6 17.0	33.5	빏		81.5 = 10-2	18. 5. 18. 5.	,
	720-875 cm-1		3.0 x 30-2	· <b>60</b> (			0.02 x 10 <sup>-2</sup>	0.1	0.0	).4. F		6.0 x 10-6	, o.	34.5	31.0 36.5		0.04 × 10-2	8 %	78.0° 50.00	1.1		0.19 = 10-2	, K	0.39	2.0		2.4 × 10°2 5.1	1 1 1	3.4.5	•
•	(200		51.2	5	<b>3</b>	, 55.	, g	0	Ŝ.	8.6	:	\$	3.8	K.	<b>8</b> 65		\$	53	12	6.18		0.3	1		<b>7</b>		äĶ?	ន្ត្រី	돈	
	(a.za (a.)		. TT	1	isi	 <b>=</b>	<b>16.0</b>	3	9.0	2 1. c 2 2. i	v,		, K	Ž,	ÁP.		67.0	o.∵	2.78	i di		<b>9</b> ,7	3.7 \$20	ei e.	9.50		33.	iää	111	
	و نو د د د		5.13		289	ę	0	`\.	۲,	in:	•	\$			r C:		<b>%</b>	Ŕ	<b>1</b> 7 17	S.		Xi X	î	:14	T.		266	 	6 K	
,	2		1;+	71	( K. (	-	83	i Ri	r í	133	G	21	8	*	R R		8	C R	取用	R		2.2	2	22	8					
	į		¥.t	2	. 3	5	84	66	2,	151		8	25	2	er er		8:	e e	4.1 - 0	8		4	9	9 5	200		1000	31	ee 되고	

The second secon

		1					A(v,-v2)			2184	
a de	žź	ž,	(a) som	H.	7.3# 618-012	651-120 cm	617-667 cm <sup>-1</sup>	545-617 cm <sup>-1</sup>	1,95-5-5 cm-1	A(v)dv	Rent rits
		}					Seth Length = 3200 cm (30at'd)	(20at'd)			
		ŧ	1 62	7.6	x 10 <sup>-2</sup>			5.7 x 13-2			548. 314: 6.0 mm Nr of
: i		T.	23.50	13.8	٠.٠			2.6			CO., Same. 115-121; H.,
; <b>A</b>		ĸ.		ж; •	<b>-</b> .			9,0			saided to Sam. 114.
		en e	, c		15-15 -8-4		15.8 x	. o.		8	2(36)-720 cm-1) and
: :		×	2.5	ă	1	67.5	1			110	from table because of ex-
2			2	Ş	20.0	ĝ.	31.1	20.3		125	ressive error due to 305
17: 6		•;	25.5	8	<u>፡</u>		33.0	7.47		135	in stm apperic sath.
	k	*	77	1,51	2.30.	57.9 1 10-2	56.9 x 10-2	9,4 x 13°2		71.5	100 100 100 0 mm Re of
	, h	, e	4	9.		ŧ	39	•		85.6	Co. Stud. 127-126; No.
	k	K	4.3	÷3.6	eri Vii	5.1	76.3	7.61		7.16	tidbed to Sam. 192.
	K	i,	<b>Y</b>	ζ.	an .	O. ≠	÷.	18.5		11	
	i, I	Y, Y	9	ğ,			0.1	61.9		129	
	5 K	÷ ::	•	<b>!</b> :		, 4 , a	<b>.</b>	7.79		1,5	
	:	:								3	
	R'	11:		2.5	, b, x	79.6 x 112	2.01 H 6.55	17.3 z . 5°		011	34m. 1294 25.0 rm Bg of
¥;	RI				e :	7.7	, ,			27	1.2. Jack. 139 W. N.
	RY			Ş		Ņ,	Ç	9.9		, i	cided to : a
	178		,	3 28	21.2	- <b>8</b>	2	5.0		35	
	R	<b>8</b> .5		9.0	22	8	38.6	6.14		172	
		8.8		*	15.5 x 10.2	24.0 £ 10-2	9.7 x 10-2	39.6 x 10-2	1.6 x 10°2		1.43. 1.45. 50 mm Be of
		. k. s		1							Co. Sarat. 136-140; No.
21	<b>7</b> .	Š.		5.9	. I. i	<b>8</b> .7	9.9	7.6	2.5		201ed to See. 175.
	:	Ç				È		, t	v. c		
	٠.	ž	5	Ž	, C. 3	12	38	53.6	` • ·	<u> </u>	
	2	×.		9	2 30.2	, 10° s	97.4 x 10-2	14.2 x 10"6	2.9 x 19-2	158	SAR. 34 : 100 mm He of
2	l	, i	製	2			8	1	t	સુ	n. ners. 140-145; M.
	ス	×.5		Ş			81	55.4	3.6	161	#.13ed to Sam: 14.1.
1	i			C	Š.		8	57.6	9.0	£67	
3	Z,	r.					8		٠.۶	500	
911 0	×	\$ 5		93	31.3 1 10.2	135 x 10-2	100 x 10-2	64.4 x 10-2	1.4	301	. 843. 2463 200 mm Eg of
		ř.	£	(5)	1.3	81	8	6.9	8,4	52	.03. Sarm. 1.7-149; W2
	.,	Š		e,e	e	130	27	70.6	5.5	515	edded to San 146.
7		ı. K		9.	æ. %	81	130	74.5	6.1	ŝ	
951	ಷ	<b>3.</b> 5		š	J 10-2	2-31	130 x 1072	75.8 x 10-2	e)	8	Sea. 1503 400 ox Hg of
151 a		<b>1</b>	£.	ž	<b>k</b> 1.0		100		9.5	£	102. Same 15t-132 #2
	¥	 		<u>s:</u>	6.3		82	8.0	01	2 7 2	edded to Sam. 150.

					-		(24-11);			57.5	
į	ž	:	,	( <del>)</del>	.2 \$17. ±°.	1-20 cs-1-1-	6.7-507 ca.1	\$+5-617 am	.95-345 3∎" <sup>1</sup>	ADV AN SIA	Permits
						अंधनः प्रमु	Park length 4 7930 cs. (Cont. 6)	( p. 10)			
			198	438	6.0 x 1,76 -7.1	) 10 10 10 10 10 10 10 10 10 10 10 10 10	501 × 021 301 × 021 301 × 021	95.7 × 10°. 96.7 86.7	14.5 x 10-2 14.7 15.6	25.6	Sca. 153; 740 um 1g of 315. Samm. 154-155; Ng wided to Sam55.
********	2222		2423	50 11 2 kg	10 A H G (((((((((((((((((((((((((((((((((((	4 7 8	× + C = 0 × + C = 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	0.8 x 10°3 1.0 1.6 2.0 5.1		20. UI 19.00 80.00 4.00.00	Rame, 156-151; JO <sub>2</sub> 8, 03#
22424	ххя	*****	erer Anna	2 kg 44	\$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$	88883 8883	60 60 60 60 60 60 60 60 60 60 60 60 60 6	ያልጜዾች 6 ዲዕተ-ጎ		\$25 £55	Na. 162: 950 am kit of No. Samo. 165-166; Na. 212 heaved. Resulta Na. 21 A(495-54) cm. 7 for Na. 164 bockur of cm. Included bockur of cm.
な程まった	R P R	Sugar Sugar	2884x	8.9 8.90.	\$1.00 m \$1.00 m \$1.00 m \$1.00 m \$1.00 m \$1.00 m	88888	48888	88 17 7.7 8 8 4 4 6 8		80 20 20 20 20 20 20 20 20 20 20 20 20 20	Sam. 167, 1200 mm Rg of 17p. Same. 163-171; Sam. 167 beaved.
200	<b>N</b> R	E E	꿪춖쮸끖	សូទិ: សូទិ: ស	21.2 21.0 25.7 25.9	* 9 77 3. RRRR	97.58 8.17.68 1.18.69	4 0 <b>0</b> 0 2 2 2 2 2 2 2 2 2		155 168 156 156	isan, 172; 255 we Mg of 100-, Same, 173-175; Essa, 172 besteri.
2116	<b>X X</b>	na par	7 5,512, 2,4,4,4	7) 5; A 75; A 76;	90 W 1 W	292 292 293 293 293 293 293 293 293 293	68.9 70.4 71.0	10.5 13.0	;	988.7 98.1. 99.1.5	15an, 176; 50 mm Hg of 15g. Shams, 17779; lass, 176 benthed

Table 4. (Cont'd)

Sample No.	Fig. No.	Temp. C. deg.	v (atmos em)	(ma Ag)	Ā(720-875 cm <sup>-1</sup>	Remarks
				Longth -	200 ca	
c 26 c 27 c 28 c 29 c 39	17 17 17	25.0 25.0 25.0 25.0 25.0	305 305 305 305 305	203 236 332 767 3070	20.2 x 10 <sup>-2</sup> 24.4 26.0 29.1 33.5	Spectra or Same. C 26 - C &2 were obtained in connec- tion with the in- vestigation of the 1064 and 961 cm <sup>-1</sup> CO <sub>2</sub> tands. A HaC: prism was used and
C 31 C 32 C 33 C 33	17 17 17	25.0 25.0 25.0 25.0	685 685 685 685	212 389 798 3075	70 0 32 4 34.8 36.6	frequencies less than 670 cm <sup>-1</sup> could not be investigated Sampling procedures for Same. C 26 - C h2 are given in Table 3.
C 35 C 36 C 37	17 17 17	85.0 85.0 25.0	1245 1845 1245	2150 985 #87	37.1 40.0 43.7	
c 38 c 39	17 17	25.0 25.0	2945 2945	1780 1240	51.3 53.0	
C 40	17 12	25.0 25.0	5750 5750	1990 33 <b>8</b> 0	61.0 62.7	
c 42	17	25.0	11,800	1800	71.7	

Table 5. Band Intersities

The values of jk(v)dv are in units of atmos<sup>-1</sup> cm<sup>-2</sup> S.T.P.. Values given for the present study correspond to a temperature of 26°C, and the estimated uncertainties are indicated.

Creev. set.	∫±(v)dv
3716 cm <sup>-1</sup> Bank	
Present Study Eggere and Crawfore <sup>10</sup>	54 ± 10 39
1609 cs-1 Band	
Present Study	57 ± 8
8359 cm <sup>-1</sup> hand	
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#### PART C

# ADSORPTION BY WATER VAPOR AND NITROUS OXIDE

bу

Darrell E. Burch, Edgar B. Singleton Wilbur L. France, and Dudley Williams

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#### ABSORPTION BY WATER VAPOR AND NITROUS OXIDE

#### I. INTRODUCTION

This is the third and final report dealing with the results of F research program concerned with the absorption of infrared radiation by various aumospheric gases. The contents of the three reports are as follows:

REFORT 11: Absorption by nitrous oxide in the 2224 cm<sup>-1</sup> region; absorption by carbon monoxide in the region of its fundamental (2145 cm<sup>-1</sup>) band and its first overtone (4260 cm<sup>-1</sup>) band; absorption of methane bands at 3020 cm<sup>-1</sup>, 1550 cm<sup>-1</sup>, and 1306 cm<sup>-1</sup>.

EMPORT  $2^2$ : Absorption by carbon dioxide bands at 2350 cm<sup>-1</sup>, at 3609 cm<sup>-1</sup>, at 3716 cm<sup>-1</sup>, at 961 cm<sup>-1</sup>, at 1056 cm<sup>-1</sup>, and in the range 875-895 cm<sup>-1</sup>.

FINAL EXPORT (herewith): A study of the relative pressure "broadening shillties" of all of the ebscribing gases studied in the research program, and of various inactive gases. """ption by water vapor in the region of its viprution-rotation bands at 1595 cm<sup>-1</sup>, at 5352 cm<sup>-1</sup>, and mear 3700 cm<sup>-1</sup>: absorption by mitrous exide bands at 2563 cm<sup>-1</sup>, 2461 cm<sup>-1</sup>, 1265 cm<sup>-1</sup>, 1167 cm<sup>-1</sup>, 692 cm<sup>-1</sup>, and at 589 cm<sup>-1</sup>.

A THE THE RESERVE THE DATE OF THE PROPERTY BY THE PROPERTY OF THE PROPERTY OF

A portion of the inta us the scinive breadening chilities of inmit're gases, and on the N<sub>2</sub>O, CO, and CO<sub>2</sub> bends, was reported previously in Fh.D. dissertations by two of us.John Mosever, additional data included in the three reports outlined above were not included in these dissertations.

The present research program is a continuation of two earlier studies made in this department, in 1940 by Chapman and assured, 5 and in 1953 by Housel, Bureb, and Villiams (horseful referred to as MEV).

The experimental methods used in gas sampling, in obtaining the elemption curves, and in analyzing the data were described in Suport 1 and will not be repeated, with the exception of the method of sampling water vapor which is described below in Section IV:

‡

The original place included a study of unter vapor absorption in the MO-1000 ca<sup>-1</sup> region; however, this portion of the otaly was post-posed and will probably be done as a part of amount fature program in this laboratory.

#### Symbols und Units

Values of absorber partial pressure p, total pressure P, and equivalent pressure  $P_q$  (defined by equation (5) are expressed in units of mm Fg absolute. Absorber path langths are in on. Values of absorber concentration q of water vapor are in units of precipitable continueters (abbreviated pr, cm) which represents the depth in continueters of the water that would be obtained if all the water vapor in a column containing the beam were condensed out. The values of absorber concentration of other gas samples are expressed in units of atmost q, and are found by multiplying the absorber path length in the absorber partial pressure in atmospheres corrected to standard temperature.

Programatics and values of total obsorption  $\int A(\nu) d\nu$ , where  $A(\nu)$  is the observed fractional elementies at frequency  $\nu$ , are in units of  $cm^{-1}$ ; and band intensities  $\int k(\nu) d\nu$ , where  $k(\nu)$  is the absorption coefficient at frequency  $\nu$ , are in units of atmost  $cm^{-2}$  S.T.P.

#### II. RELATIVE PROADERIES ABILITIES

A method of measuring the "self-broadening ability" of an absorbers gas relative to the "broadening ability" of the non-absorbing gas Mg has been given in Report 1. It was shown that the analysis of total absorption data obtained from simples containing Mg along with an absorbing gas was simplified by use of an experimentally determined ratio of the broadening abilities of the gases. The broadening effect of both gases can then be expressed in terms of a single variable, which is called the equivalent pressure and is defined below in Mg. (5); and the total absorption can be expressed as a function of two variables; absorber concentration, and equivalent pressure. In this section all the results of telf-broadening measurements control out in the present research investigation are summarised, and these results are discussed in some detail and compared with those of other workers.

Aside from its importance in the enalysis of infrared absorption data, a inculate of the relative breadening shilities of games provides information conserving the collision erose sections of the unlessles, which are determined by various furces of interestion between the nelection. For the same reasons that a bacadeign of calf-breadening is important, so is a knowledge of the breadening shilities of numberoting games other than Rg, such as Qp, Ro, Ar, Rg. The latter portion of this section deals with the relative breadening shilities of these numberoting games, so well as with some other games such as Q0, Q0, and Q0, which were used as "furnign" breadening games in the study of absorption bands coourring of frequencies where these games do not absorb.

The half-width  $\sigma$  of a collision-breakend spectral line is graper-timal to the collision frequency F of each absorbing molecule. From binetic theory it can be seen that the half-width  $\alpha$  of a collision-breakend line is given by the relation

$$a = \frac{p}{2\pi} = \frac{1}{4\pi} \frac{\pi}{2} \, a_2(a_{0,2})^2 \, \left[ \operatorname{ann} \left( \frac{1}{2\pi} + \frac{1}{2\pi} \right) \right]^{2/2},$$
 (1)

where  $E_{i}$  is the number of unlexalor of the i-th type per unit volume,  $b_{i+1}$  is the sum of the systeal collision diameters of the absorbing unlexale and a unionis of the i-th type, it is interests's constant, I is the absorbing televale, and  $b_{i}$  is the absorbing unlexale, and  $b_{i}$  is the axes of a unlexale of the i-th type.

The appearance of M<sub>1</sub> and M<sub>2</sub> is (A) indicates that different passes have different "reddening" affects on operated line. Since most of the gas complex employed in the experiencial wark, existed of binary mixtures of absorbing gauss and mitragen, which was used as a number-printing "trundening gas," it is destroble to consider (A) for the

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special case of a binary mixture of an absorbing gas a and a broadening gas b. The summation in (1) will in this case have two terms: a self-broadening term in which i = a, and a second term for the effects of broadening by the inert gas i = b. Equation (1) that assumes the form.

 $\alpha = \frac{1}{18} \left[ 2 \pi \kappa^{\frac{1}{4}} \right]^{1/2} \left\langle x_{a} \left( D_{a,a} \right)^{2} \left[ \frac{2}{N_{a}} \right]^{1/2} + H_{b} \left( D_{a,b} \right)^{2} \left[ \frac{N_{a} + N_{b}}{N_{a}} \right]^{1/2} \right\rangle (2)$ 

By recalling that the partial pressure  $p_1$  of gas of the 1-th type is given by  $p_1=R_1kT$  and by assuming that collision diameters are independent of temperature and pressure, one can write (2) in the form.

$$\alpha = \frac{1}{kT} \left[ \frac{2T}{kT} \right]^{1/2} C_{a,a} p_a + C_{a,b} p_b , \qquad (5)$$

where  $C_{B,R}$  and  $C_{B,R}$  are constants involving the optical collision diameters and masses of the absorbing and broadening gases. Since the total pressure  $P = p_B + p_B$ , one may obtain an expression for G in terms of P and  $p_B$  by rewriting (5) in the form

$$\alpha = \frac{1}{4\pi} \left[ \frac{2\pi}{kT} \right]^{1/2} \quad C_{n,b} \left\{ P + (B-1)p_n \right\} \quad , \tag{b}$$

where B =  $C_{ab}/C_{ab}$  is called the self-browlening coefficient of the absorbing gas and represents the ratio of the "self-browlening shillty" of the absorbing gas to the "broadening shillty" of the somebrorbing gas. Nothods for determining B for various absorbing gases relative to No will be described below. The term is brackets in (4) is called the equivalent pressure  $P_a$  of the gas sample.

$$P_{q} = P + (B-1)p = p_{Bp} + 2p$$
 (5)

where p is the partial pressure of the absorbing gas and P is the pressure total pressure of the comple due to the absorbing gas and  $R_{\rm p}$ . It is seen that  $P_{\rm p}$  reduces P in the case of p <C.

Fina (\*) one one that for a given temperature it is proporticall to  $P_{\alpha}$ , and once the value of P has been determined it is possible to describe the "broadening effect" of the two games as term of the " $e^{i\alpha}$  presents  $P_{\alpha}$ . It was shown in Report 1 that the fractional absorption A(r) and the total absorptionA(r) are functions of only two parameters,  $P_{\alpha}$  and absorber concentration  $\alpha$ .

The experimental determination of B involves a stud of the transmission of radiation through cells of different long, a in which the absorber concentration y is the same. The sample in tar short cell, called the "reference rell," consists of a pure sample of the

shaorbing gas at pressure p; for this cell Pref. = pref. and the equivalent pressure is given by

The sample in the longer cell, called the "sample cell," initially consists of a sample of the pure absorber at the partial pressure p required to produce an absorber concentration w equal to that in the reference cell. Initially, the absorption of radiation in the sample cell is considerably less than that in the reference cell, since the partial pressure and total pressure are less than those in the aborter reference cell. Various amounts of H<sub>2</sub> are then added to the reference cell corresponding to equivalent pressures

$$P_{e} = p_{R_{2}} + E_{p} \quad (Sample Cell), \tag{7}$$

where  $p_{RO}$  is the partial pressure of the broadening gas. For the situation in which the absorptions in the sample and reference cells are most nearly equal, it is assumed that the values of  $P_0$  in (6) and (7) are equal. The value of the self-broadening coefficient can then be determined from the relation

$$B = \frac{B_{Lot}^{-D}}{b^{165}} \tag{9}$$

where pg is the partial pressure of H<sub>2</sub> in the longer sample cell when the absorptions in the two cells are equal. If the reference cell also contains H<sub>2</sub> at a partial pressure pgg. it can easily be shown that

$$P = \frac{p_{\rm HC} - p_{\rm HC}}{p_{\rm ref}} \qquad (9)$$

The value of 8 was determined for several of the checrytica bands included in the present research program by making use of the double-beam feature of the Fertin-Liner House & spectrumeter. In some cases a 5.15-cm absorption cell was used in the sample beam with a 1.55-cm cell in the reference beam; in other cases a NOD-cm reference cell was used with either an 800-cm or 1600-cm remain cell to investigate samples having large values of y.

Typical spectral comparisons which were obtained in this manner for the 200% cm<sup>-1</sup> 8gO bank are shown in Part I of Fig. 1 or various values of total pressure in the sample cell. The spectrometer recorded the ratio of the transmission of the sample cell I(cam.) to the transmission of the dashed line at

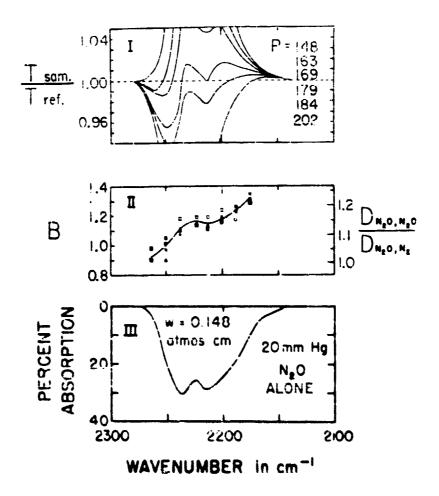


Fig. 1. The colf-breakming coefficient B for the 220t on 1 Bg0 is both port I: Spectra obtained with w = 0.50 across on 1 Bg0 is both reference and couple colds; twent pressure in reference cell = 16h mm Mg. Total pressure in camble cell for each opertrum to indicated.

Part II: Flot of B and collision cross contion ratio. The four different geometrical points correspond to the results of the cost of spectra chann in Part I as well as to three other acts of comparison being different values of y.

Part III: A typical absorption curve which is about for comparison.

T(sam.)/T(ref.) = 1.00 represents the recorder pen position when the two beams were "balanced," i.e., the transmissions or both samples were the same. It is seen that the ratio of the transmissions decreased with increasing pressure in the sample cell as would be expected, and that some of the curves representing the ratios occur show the balance line for come frequencies and below the line for others. Therefore, the total pressure corresponding to balance was not the mass for all frequencies; the pressure corresponding to balance for any given frequency was determined by interpolating between the two successive curves which occur on opposite sides of the balance line.

Since the value of pm, corresponding to equal transmission of the cells was found to vary with frequency, it is apparent that the value of B varies throughout the band. Points corresponding to the experimentally determined value B were plotted against frequency in Part II of Fig. 1. The collision cross section ratio which is shown by the right-hand scale in Part II of Fig. 1 can be shown by use of (2), (3), and (4) to be related to D by

$$B = \left[\frac{2N_0}{N_0 + N_0}\right]^{1/2} \left[\frac{D_{N_0,0}}{D_{N_0,0}}\right]^2 \tag{10}$$

where Ma is the mass of the Ma solecule Men.

A typical absorption curve of an HgO sample is shown in Part III of Fig. 1 in order that the "attracture" of the absorption curve can be correlated with the structure of the curve relating 8 to wavenumber.

In Pigures 2, 5, and 4 are above curves of 8 vs. wavesumber, and the related rollision cross section ratio, for the 3000 cm<sup>-1</sup> and 1506 cm<sup>-1</sup> CM<sub>0</sub> bands and for CO<sub>0</sub> in the 657 cm<sup>-1</sup> region, respectively; a typical absorption curve is shown in each figure for comparison. The curves of 8 for these bands were chosen as illustrations since the value of 8 changes very definitely throughout the bands. The value of 8 was found to be more nearly constant for other bands investigated by the method described above.

Is should be emphasized that the recurrecy of the value of 3 determined by this method depends atrongly upon the accuracy to which the absorber consentration y is unde the same in both the reference and sample cells. If y were not the same in both cells, a false value of 3 would be determined; the shape of the curve relating 3 to savenumber would also be sodified.

In the process of measuring 3 for each band, some of complex baving different values of y and I were used; complex of at 1 y were used primarily to determine 3 at frequencies of strong absorption, while samples of larger y were necessary to levestigating the wings of

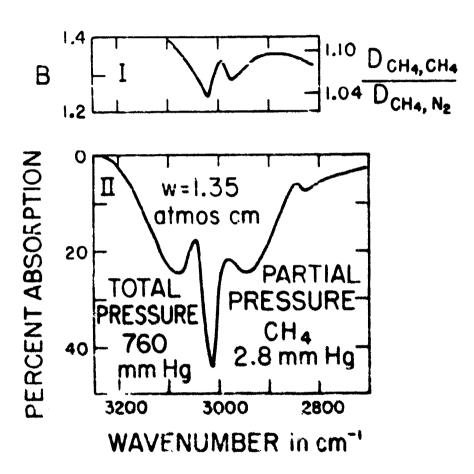


Fig.2. The self-breakening coefficient 3 for the 5050 cm<sup>-1</sup> Ch bank.

Fort I: Not of 3 and the sollision cross section ratio.

Fort II: A typical obscription curve which is shown it comparison.

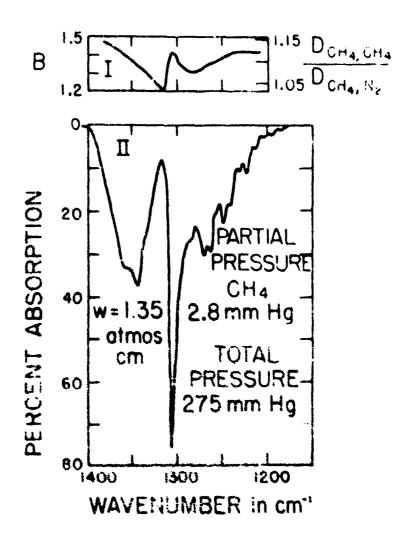


Fig. ). The pair-branching confidence b for the 1306 on A CR, band.
Fort I: Plot of B and the callinger cross section to be
left I: A typical absorption curve which is obout for comparison.

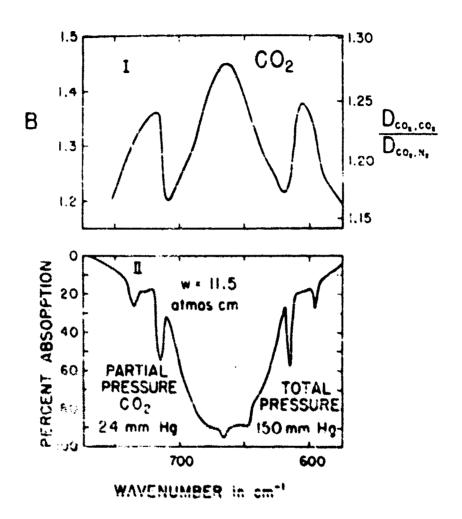


Fig. 1. The self-breakening coefficient 3 for CO<sub>2</sub> in the obj cm<sup>-1</sup> region.

For i: Plat of Find the collision close certion ratio.

[Mill: A Uplant absorption curve which is shown for comparison

the bands. The samples were formed so that the fractional apsorption of the samples being compared was usually between 0.3 and 0.9 in the spectral inverval being investigated. For smaller fractional absorption the spectrometer was less sensitive to small changes in pressure, and thus, the total pressure corresponding to equal absorption in the two cells could not be determined as accurately. For values of fractional absorption greater than approximately 0.9, the power generated by the detector was too low to properly drive the servomechanism which comprotes the comb in the reference beam; a false reading might ther be obtained unless extra care was taken. This difficulty could be overcome to some extent by increasing the amplifler gain and increasing the scanning time; however, as the fractional absorption approaches unity, the effect of scattered light of frequencies other than that being studied might become important and give rise to false readings. If the scattered light were equally intense in both the reference and the sample beams it would have no effect on the reading of the spectrometer, but it is possible that the intensities of the scattered light in the two beams were not equal; and values of B were not calculated for samples at frequencies at which the fractional absorption was greater than approximately 0.9.

Total pressures were also restricted to values less than approximately 1000 mm Hg, since for larger pressures the spectral lines become smeared and the dependence of absorption on pressure becomes less.

For the ranges of pressures and fractional absorption of the samples investigated, no dependence of the measured value of B at a given frequency on either of these parameters was found.

Some of the late were obtained by use of a single-beam Perkin-Elmer Model 99 spectrometer which was sounted in a vacuum tank\*. This spectrometer was used to determine B for the 4260 cm<sup>-1</sup> CO band, since this band could not be studied by use of the Model 21, which was limited to frequencies below 4000 cm<sup>-1</sup> because of the prism used. The Model 97 was also used to investigate the total absorption and to measure B for some etrong Gap and HgO bands.

Since the Model 99 is a single-beam instrument it was not sourtble to compare simultaneously the fractional absolution of two samples in different chils by the method described in connection with Fig. 1. Instead, "curries of growth," of total ebsorption were obtained for various samples with two or more different cell lengths. The "effective" value of 9 for the entire band one determined from the curves by finding the total pressure required for a sample in the longer cell to produce the same total absolution as a sample in the charter call which had the rate absolute concentration. A sample calculation for this method of determining B was given in Report 1 in connection with the 4260 cm<sup>-1</sup> CO band.

<sup>&</sup>quot;This apparatus is described in some detail in Appendix T of this report.

The present research program deals primarily with total absorption rather than with fractional absorption at certain frequencies. It was therefore desirable to obtain a nominal value of B which could be used for an entire band. The method of determing B by use of the single beam instrument, of course, gladed such an effective value of B since it was determined on the basis of total absorption. In the case of bands studied with the double-beam spectrometer, a nominal value of B for the entire band was determined by taking a "weighted average" value of B from curves such as those in Figs. 1-4, with extra weight given to the frequency ranges which contribute most of the absorption.

The values of B determined for each band included in Reports 1 and 2, as well as in this report, are listed in Table 1, and wherever possible they are compared to the results of other workers. Except for the case of H<sub>2</sub>O, the tabulated values are believed to be accurate to +6% as an effective value for the entire bands.

It is somewhat more difficult to determine B for H<sub>2</sub>O backs than for the other bands for two reasons: (1) the value of shorber concentration w of an H<sub>2</sub>O sample cannot be found as accurately because of the adsorption of the H<sub>2</sub>O on the walls of the absorption cell. The method of determing w for H<sub>2</sub>O samples, which involves finding the dew point, is discussed below in Section IV; (2) the range of available partial pressures of H<sub>2</sub>O is limited to values lass than the vapor pressure.

In the investigation of E<sub>2</sub>O, a multiple-traversal cell whose path length could be adjusted by external controls, was used; different path lengths corresponding to 4, 18, 16, and 52 traversals were therefore easily obtained. From curves of growth of total absorption for different samples investigated at different path lengths it was possible to determine B for the H<sub>2</sub>O beands. Because or the limitations stated above with regard to the accuracy of determining B for the B<sub>2</sub>O bands, the value may be in error by as much as 25 to 30%.

It is apparent from the curves of Figs. 1-4 that a nominal value of a sudder integral, inage whightly with total absorption, since for large total absorption the major portion of the growth occurs at different frequencies than for small total absorption. Since the nominal value was obtained by "seighting" the value of B at different frequencies according to the centribution to the growth, the best effective value for the entire band would therfore change with total absorption. However, it is seen from the curves of Figs. 1-4 that B varies by only a few percent for the different frequencies, and any change in the nominal value of B for the entire band would be negligible.

It was stated above that the measured value of B at my frequency did not apparently change with pressure or with fractions, absorption for the limits over which those parameters were varied. Honedist and Painer10, in an article on the shoorption by RgO in the 20 micron region,

Table i. Self-Broadening Coefficients

i

deri	Investigator	Self Fromdening Coefficient B
222k cm <sup>-1</sup> , 1265 (m <sup>-1</sup>	Present Study	1 13 + 6.17
and 16, cu		10:0 + 37:7
1285 cm-1	Souty and Mormell <sup>7</sup>	1.27 ± 0.3
1167 cm-1	Sooty and Wormell?	1.35 ± 0.07
2.24 cm-1	Cross and Daniels	1.2)
8		
2.45 cm-1	Present Study	1.02 ± 0.06
1,:50 cm <sup>-1</sup>	Present Study	1.08 ± 0.05
+ 10		
3020 cm-1	Prosent Study	1.30 ± 0.08
1506 cm <sup>-1</sup>	Present Study	1 38 ± 0 08
3716 cm-1, 3609 cm-1, 2350 cm-1, 1054 cm-1, 251 cm-1	Present Study	1.70 ± 0.08
bends and the 815-495		
Same CC2 bands	Bt.on. 139	2 + 0,50
		(where p = Ch <sub>2</sub> partial pressure).
120 120 120 120 120 120 120 120 120 120		
	Fresent outdy	5 # 1.5
20 increase region	Palanti	6-11, increasing linearly
4,325 & cm-1 HgO 1836	Vesiloveky and Deporant <sup>11</sup>	with fractional absorption 6
4 times in 500-60, cs-4 region	Instill	3,6-5,5

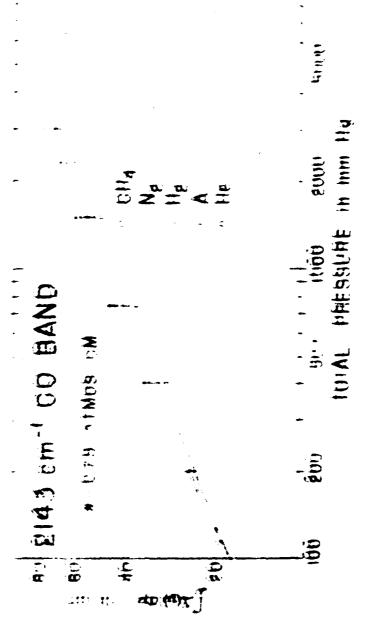
"B was not seasured for 2565 tm-1, 2161 -m-1 and 589 cm-1 bands; but 1.12 was used in the analysis of the bunds.
\*B was not seasured for 1550 tm-2 band; 1.36 was used in analysis of the band.

have suggested that the value of B should increase with fractional absorption. This phenomeonon arises from the fact that the spectral lines do not have a true Lorentzian chape, and that a colf-broadened line absorbs more in the wings than does a foreign-broadened line of the same halfwidth. Since the line centers are virtually opaque in the case of large fractional absorption, the only growth with increasing pressure results from the wings of the lines, where the beli-broadened lines are stronger and the effective value of B is increased. It was not possible to observe any increase in B with increasing total absorption in the present investigation of H<sub>0</sub>O, possibly because of the limited values of w and p which were obtained. Although, according to Benedict and Palmer<sup>1O</sup>, the value of R at a given frequency in an HoO band might vary from approximately 6 for small fractional absorption to 11 for large fractional absorption, one might expect that the effective value for measuring total absorption of an entire band might not deviate appreciably from 6 since the major portion of the growth of the band occurs at frequencies for which the fractional absorption is not large. This result, which is based on theoretical calculations involving the shapes of individual lines, has been confirmed by Benedict, 13 to a large extent by calculations based on the results of total absorption measurements of MgO Sands reported by HBW. Benedict found that the HBW data yeilded a value of b equal to approximately 5 the same as the value determined in the present investigation. It is noted that MEW used a "weighted" pressure (P+p) for the 850 bands; this corresponds to B=2, a value which is believed to be too low in view of the more recent experiments and theoretical calculations.

Values of self-broadening coefficients for individual 8.0 lines have been determined by two different sets of workers using different methods. Vasilovsky and Reporant<sup>11</sup>, by considering the total absorption of a single line, found the broadening ability of 8.0 relative to that of 8.1 to be such as to correspond to 8.6. Inatt<sup>12</sup> found the measured half-width of four self-broadened n.0 lines in the 500-600 cm<sup>-1</sup> region to be approximately 5.6 - 5.5 times the calculated half-width of 8.2 broadened lines; this result corresponds to a value of 8 varying over approximately the same range, 5.6 - 5.5.

#### foreign-Gas Tituestening

The relative broadening abilities or two mert gases can be found by comparing the partial pressures of each gas which is required to produce the same total absorption when added to equal complete of an absorbing gas. The method used in the present investigation can best be described by referring to Fig. 5. A sample of CO laying a partial pressure of LOV me kg was introduced into a 6.55 cm cell, and the total absorption of the 2145 cm<sup>-1</sup> CO hand was measured. The measured value of total absorption was plotted at the lower end of the curves of Fig. 5. One of the foreign gases, he, for example, was then added to the cell to total pressures of approximately 200, 500, 700, 1500, 2200, and 5000 mm Hg; the total obscription was measured at each pressure and the results formed the points for the table of Fig. 5.



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This procedure was then repeated for each of the other hrusdening gases; ears was always taken to start with, as nearly as possible, the same amount of CO for each set of samples. The total absorption of each sample of CO slone was assesured before adding the broadening gas; the commistency of this value of total absorption provided some indication of how meanly equal were the values of absorber concentration for each set of samples.

From the curves of Fig. 5. At can be seen that a total pressure of 960 am Hg is required to produce a total absorption of 50 cm<sup>-1</sup> with Hg used as the broadening gas; a total pressure of 1450 am Hg is required to produce the seem total absorption with He used as the broadening gas. Since the partial pressure of CO is 200 am Hg in both cases, it is comcluded that 1350 am Hg of He are required to produce the same broadening as 860 am Hg of Hg.

A quantity called the foreign-broadening coefficient P of an inert gas b was defined by

$$T = y_{1/2}/y_{\rm b}, \tag{11}$$

where  $p_0$  is the partial pressure of gas b required to preduce the same line broadening as a partial pressure  $p_{0,0}$  of  $N_0$ . Thus, from the example above, the foreign-broadening coefficient of He is 860/1350 = 0.64. Unless otherwise states, all values of foreign-broadening coefficients  $F_1$  as well as self-broadening coefficients  $F_2$ , as well as self-broadening coefficients  $F_3$ , listed in this report are relative to the broadening chility of  $N_0$ .

Values of P were determined in a similar manner for the other broadening gases included in Fig. 5. Calculations were made at several different values of total arsorption for each gas; and it was found that there was no significant dependence of P upon the total absorption, which, of source, increased with total pressure. One, therefore, concludes that an unambiguous value of P can be found for each broadening gas for a given absorption band. But that its bulbe has at nost a slight dependence on total absorption or pressure for the ranges of pressure included in the pressure study.

Several other absorption bands, the could me, and 1095 mm, and 1095 mm, and 1095 mm, and the special could be such of the special could be such that such is the special could be such that the special could be such that the special could be such that the special pressure of the absorbing gas and the range of total pressures over which calculations of F were made. Also included in table 2 are values of B, 1/B, m, which can be absurabled in table 2 are values of B, 1/B, m, which can be absurabled. (1), and the information calleving th, to be related to F, the foreign-broadening coefficient of gas b, by

$$\frac{D_{0,10}}{D_{0,10}} = p^{1/2} \left[ \frac{23 + y_0}{y_0^2 + y_0} \frac{y_0}{28} \right]^{1/4} , \qquad (12)$$

where 28 is the molecular weight of Ho.

Table 2. Foreign-Broadening Coefficients and Relative Molecular Cross Sections

Absorber and Band	Partial Pressure of Alesorber*	Total Pressure Range (mm Rg)	Broadeuser 0	p++ (men)	Da, No
₩20 2221, cm-1	10	25- 1000	H <sub>e</sub> C <sub>2</sub> H <sub>2</sub> CH <sub>4</sub>	0.73 C.83 O.78 1.23 1.08	0.58 0.95 0.25 0.64 0.94
<b>H</b> gO 1285 cm <sup>-1</sup>	700	1000 180-	H <sub>0</sub> H <sub>2</sub> A O <sub>2</sub> CO <sub>2</sub> CO	0.70 1.21 0.85 0.72 1.17 0.97	0.57 0.64 0.67 0.67 1.14 0.99
57#3 cm-7	106	30)- 300/	H, A Ho CH <sub>h</sub>	0.64 0.78 0.85 1.12	0.56 0. <b>92</b> 0 <b>.56</b> 0.98
00 <sub>2</sub> 2350 cm-1	50	80- 700	He Og He A	0.49 0.81 1.17 0.76	0.52 0.92 0.62 0.93
Colle 5000 cm <sup>-1</sup> Base:1	10	30- 1000	H <sub>a</sub>	0.52	0.51
2020 um 1	<b>4.</b> 1	100- 2500	A CO <sub>O</sub>	0.56 v.82 1.25	0.56 0.95 1.15

\*Cell length \* 5.35 cm
\*\*Tabulated values of F are believed accurate to less than 10%.

Since both the self-broadening coefficients B and the fereign-broadening coefficients F are based on the broadening ability of  $N_{2}$ , it follows that Eq. (5) could be generalised for a sample containing several inert broadening gases by to yield

$$P_0 = B_p + p_{E_0} + \frac{\Sigma}{i} P_i p_{0i}$$
, (15)

where  $P_0$  is the equivalent pressure, p is the partial pressure of the electring gas, and  $P_1$  is the foreign-breedening coefficient of the 1-th gas whose partial pressure is  $p_{\rm b1}$ .

A STATE OF THE PARTY OF THE PAR

Some of the present results on the measurement or collision erose section ratios can be compared with the results of previous workers. Cross and Daniels determined values of the ratios for a few of the same inert games as were used in the present study by observing the fractional absorption at a given frequency in the 2145 cm<sup>-1</sup> CO bund and the 2224 cm<sup>-1</sup> 350 band. This method is in contrast to the writed of measuring total absorption which was used in the present study. The results of Gross and Baniels are compared with the present results in Table 3.

Table 3. Comparison of Ratios of Collision Green Sections With Results of Cross and Daniels  $\theta$ 

ъ	4 - # 22% er		Da,b/Da, Mg	a = 0	) Journal	
	Cross and Pagiols	Present Study	\$ Dar.	Gross and Designs	Present Study	DIEL:
Be	0.63	0.58	ė 0	0.6A	0.₹6	12.7
ka	0.70	Ů.04	6.7	0.68	0.56	11.5
O <sub>2</sub>	0.9%	9.93	4.1	0.95		
<b>A</b>	0.98	0.95	5.6		0.92	
CIL,					υ. <b>9</b> δ	
C2#6	1.09			1.09		

Benesch and Elder weed a high-resolution spectrometer to observe a quantity proportional to the line width of an individual rotation line. The results of these workers, which were based on the July line of the 3020 cm-1 CB<sub>k</sub> hand, are compared as far as is applicable in Table 4 with the present results based on the total absorption of this same City band.

Table 4. Comparison of Broadening Abilities and Collision Cross Septions with Results of Benesch and Elder<sup>14</sup>

	Broadening Abili	ty•	Collision Crois-Sections		
<b>b</b>	Beneach and Elder	Present Study	Benesch and Elder	Present Study	
£.	1.00	1.00	2,00	2.00	
A Oz	1. <b>39</b> 1.47 1.73	1.44	3.24 3.62 2.28	3.31	
8. <b>8.4</b> %	1.60	1.76	3.82 3.51	3.56	
<b>20</b> 2	1.79	1.91	5.7	4.11	

<sup>&</sup>quot;Strondening ability of He normalized to 1.00. "Gullision cross section normalized to  $D_{n_1, 2\sigma} = 2.00$ .

## Experimental Results for the 8g0 Pands at 2563, 2461, 1205, 1107, 692 and 539 cm-1

Absorption data were obtained for the 2563, 2861, 1265, i167, 092, and 589 cm<sup>-1</sup> E<sub>2</sub>0 bands, by use of a Perkin-Elmar Nodel 21 double-seem spectrometer; an NoCl prism interchange unit was used for all but the latter two bands, for which a EEr unit was used. As in previous reports, the data are presented in tabular form and by showing representative absorption curves; and the observed values of total absorption are related to the two presentars, absorber concentration y and equivalent pressure P<sub>0</sub>, by the use of curves. Value: of equivalent pressure were calculated for each sample by the following equation:

$$P_0 = p + 0.12 p_1$$
 (14)

where  $\overline{r}$  is the total pressure due to  $B_{\overline{p}}0$  at partial prevsure p and to tan  $B_{\overline{p}}$  in the sample.

In Fig. 6 are shown representative absorption curves of the 2563 and 2461  $\rm cm^{-1}$  bands; the values of y and  $P_y$  corresponding to each extre are indicated. The absorption data for these two bands are listed in Table 5.

In the left-head portion of Fig. 7 the total absorption of the \$500 cm<sup>-1</sup> \$40 bead is plotted against \$70, with each curve corresponding to a consumer value of y. The quartal features of curves similar to those presented in this report were assumed in considerable detail in Supert 1 and will not be reported. Values of total absorption were taken from the curves in the left-head parties and repletted epilost y in the light-head parties and repletted epilost y in the light-head parties. It is believed that the "emortuse" curves in the right-head parties of Fig. 7 can be used to predict the total absorption of a known \$70 sample to \$15, and the curve of theal absorption less than approximately \$0 cm<sup>-1</sup> for values of theal absorption less than approximately \$0 cm<sup>-1</sup> for values of theal absorption less than approximately \$0 cm<sup>-1</sup> for values of theal absorption less than approximately \$0 cm<sup>-1</sup> for values of theal absorption less than approximately \$0 cm<sup>-1</sup>.

A station set of curves were drawn for the 3461  $\cos^{-1}$   $\theta_00$  band in Fig. 6.

final, and lightly have scatter total absorption of the 1865 and 1167 on 1 200 backs in a manner similar to that used in to present about. It is rether difficult to compare the findings of those various summand process causian since their were not invaluant. We figures, which were similar to fine 10 and 11, were presented to relate the total

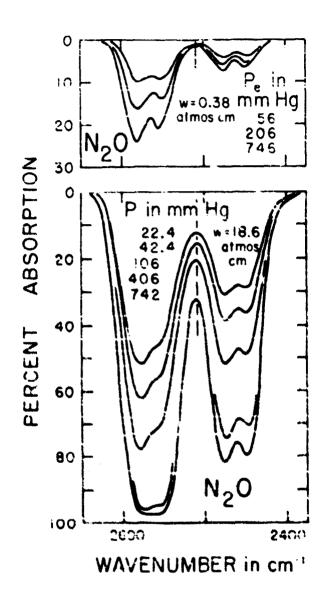


Fig. 0. Absorption curves for the 2565 earl and 2501 curl B<sub>1</sub>0 ends.
The vertical line at api0 earl represents the frequency at which
the banks were divided. Spectral alit vidits equals approximately
21 earl

made 5. Data for the 2563 and 2461 cm<sup>-1</sup> MgO Sands\*

	P	P	P <sub>e</sub>	*	JA(v)3	v in cm <sup>-1*</sup>
no no	m Bg	m Ng	⊷: ¥£	atmos cm	256 cm	Beng S401 c
		7	ath leng	uh = 6.35	ca.	
ī	10	140	741	0.075	3.2	0. <b>8</b> C
2 3 5 6	70	<u>5</u> 0	56	Ŭ. <b>5</b> 8	5.8	2.38
3	50	500	206	o. <b>38</b>	9.7	3.22
•	50	740	746	0.38	13.8	4.03
5	103	100	113	0.75	11.6	5.0
6	100	400	11.12	0.75	16.7	5.9
7	100	740	752	0.75	19.1	6.0
8	SOC	Súu	224	1.5	21.8	9.4
9	<b>\$0</b> C	740	764	3.5	31.6	12.0
		3	beh leng	<u>r</u> - joo o	<b>3</b>	
LU	12.4	12.4	13.9	5.9	17.5	9.4
17	12.4	25.4	26.	4.9	21 6	12.0
2	12,4	65.0	<b>54.</b> 5	5.4	29.5	15.2
13	12.4	196	198	5.4	43.5	23.0
. L	12 4	763	7.0	7.3	59.8	31.9
		•	est level	<u>13 - 300 g</u>	•	
15	20	<b>2</b> 0	22.4	18. ù	37.3	۲.)
16	20	<b>6.3</b>	42.4	18.6	8.44	26.7
17	<b>20</b>	10h	106	18 é	57.1	97.7
1.75	•	NO.	<b>~</b> 36	10.6	17.6	>3.5
i Ģ	£ 3	160	742	18.6	81.7	58.0
₽¢	:•	50	46	16. T	ns. B	47.5
P1	<b>~O</b>	1700	106	RN T	77 6	44.0
72	₩	we f	406	46.	82.7	71.9
?1	12	· * .	1 A 8 4	Mark.	14.7	Tie, a
		5	ere franc	r - 1600	<u>ca</u>	
٨.	w	40	44.5	76.5	73.9	54.5
<b>2</b> }	44	163	166	76.5	86.6	72.8
et.	w	200	810	74.4	20.3	7.0

offer unconstant bands never thread at 1537 and approximately by percent and approximately approximately than the second approximately app

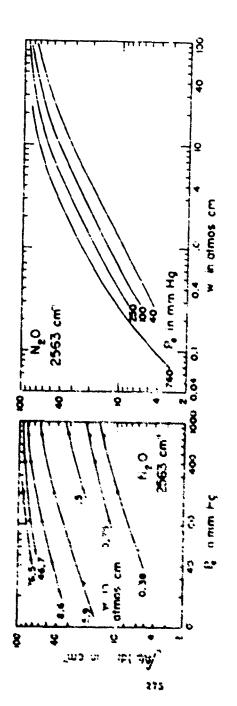


Fig. . The total ebsemption of the 2563 and MgO bend versus P, and M like a pedicte in the left-band purties correspond to samples of MgO and Mg.

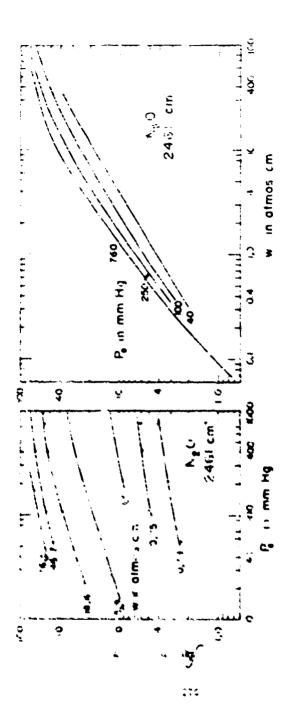


Fig. 1. The total absurption of the Biblica's mid bend versus P<sub>0</sub> and t. The Total Bush is the Least poster Correspond to samples if High about, the circle points to comples of High about, the circle points to comples of High and High.

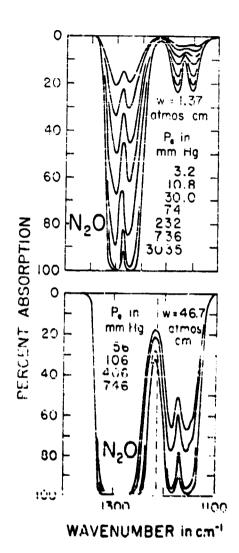


Fig. 9. Absorption curves for the 1265 cm<sup>-1</sup> and 1167 cm <sup>1</sup> No bends.
The restinal bears line w. Lety cm<sup>-1</sup> represents the requescy at which the bands were divided. Spectral slit widt, equals approximately 10 cm<sup>-1</sup>.

Table 6. Data for the 1285 and 1167 cm<sup>-1</sup> NgO Fands\*

			_	v	IA(v)di	in cm-1
Sample W∧	p man Hg	P mm Hg	P. Hg	ecesta ca	1285 cm-1	1167 cm-1
		Pa	th Length -	<u>6.35 cm</u>		
1	100	100	112	0.75	36.4	5.4
a 3	1.00	200	575	0.75	46.0	6.6
3	700	400	412	0.75	35.5	7.2
4 5 6	100	740	752	0.75	61.4	8.2
5	100	(40	752	0.75	60.3	7.5
6	100	1475	1485	0.75	64.2	8.0
7	100	2290	2500	0.75	65.4	<b>8.</b> 0
8	170	<b>305</b> 0	<del>306</del> 0	0.75	66.8	8.2
9	100	100	3 (2	0.75	36.7	5.5
.O	100	740	752	0.75	60.5	7.6
.1	100	1470	1410	0.75	61.7	<b>7</b> - <u>7</u>
2	100	2290	57/0	0.75	65.0	7.8
3	100	250	3010	0.75	Şt. ş	8.2
		Pa	th Length -	*00 cm		
•	2.9	2.9	• •	1 **	15.8	5.9
.5	2.9	5.9	6.2	1.77	17.0	3.4
.6	2.9	10.5	lu.đ	1.57	20.9	3.6
.7	2.9	27.7	·C.0	4. 37	50.1	5.8
.9	2.9	**:	7	1.37	÷0."	7.4
9	2.9	5 32	5.25	1.57	57.5	9.9
10	3.0	736	730	1.51	71.*	11.9
1	8.9	50.55	3035	1.57	76.0	12.5
2	5.8	5.8	6.5	2.45	24.1	4,5
3	5.6	12.5	15.2	4.77	52.6	6.0
•	5.8	28.5	29.2	2.75	41.6	õ. <i>≩</i>
,	; <b>.</b>	73	***	***		11.5
Ġ	5.8	<b>?</b> \$**	524	4 11	90.7	16.3
ŧ	5.6	in in	14.	ě	<b></b>	44.4
ð	5.8	\$050	3050	2.75	8° #	23.3
•	12.4	12 -	13.0	5.9	44. 6	9.4
Č.	12.4	25.2	26.7	• •	53.3	14.5
1	12.4	63	6 h 5	5.9	5.0	17.0
2	12.6	126	196	4. 7	82.1	25.6
3	12.4	749	756	5.9	90.5	•

Table 6. (Continued)

			_	٧	LE(v)AL	10 cm-1
Sample No.	ma Hg	an Ĥg	ner He	etmos CM	1205 cm-1	1167 cm-1
		Pat	h length -	800 cm		
بلو	10.0	10.0	13.2	9.3	47.2	11.6
	10.0	20.0	21.2	9.3	57.6	14.8
36	13.0	40.0	41.2	9.3	56.8	18.8
37	10.0	100	101	9.3	80.3	25.0
35 ኤ 37 <del>5</del> 8	10.0	400	401	9.3	90.2	36.0
39	10.0	740	741	9.3	92.6	43.0
μū	10.0	2310	2310	9. 3	95.6	8.A.
<b>41</b>	20.0	20.0	22.4	9.5	74.2	83.8
42	20.0	<b>60.0</b>	bC. b	9.3	01.4	23,6
43	20.0	<u>ين،</u>	106	9.3	91.0	39.0
**	20.0	707	406	9.3	99.6	55.8
45	20.0	740	742	9.7	\$7.6	62.4
46	50	50	5ô	46.7	96.k	46.3
47	<b>5</b> 0	100	106	46.7	100	55.8
¥£	รบ	400	AL SE	46. /	107	73.6
49	50	740	7+0	46.7	109	78.6
		ME	h length =	1600 34		
50	<b>40</b>	<b>40.</b> 0	<b>%</b> 5., <b>6</b>	75.6	101	54.6
<b>51</b>	<b>\$0</b>	165	799	75.6	110	76 9
52	NO.	206	211	75.4	111	77.6
ر،	100	760	112	بهر	116	84.6
5Å 55	100	MOS.	414	189	152	96.
12	lou	\$0.1	لندر	16,	1.72	34.7

The two shourytion burds were divided at 1815 cg-1, office uncertainty in observed total absorption is approximately speccent for values greater than 95 cm-1, and concents greate for smaller values.

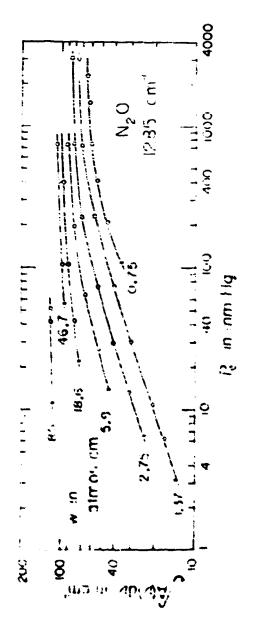
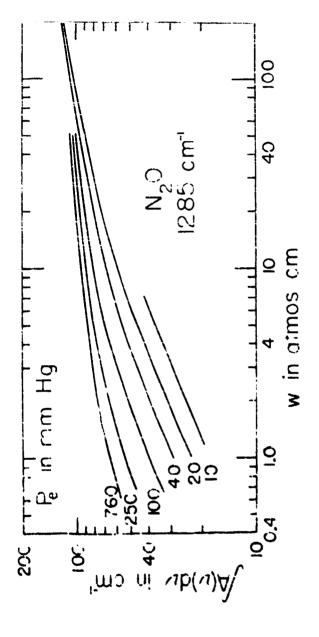


Fig. 10. The total accomption of the LMC, cut. M.O band versus for the life by points agricultured to samples of MCO alone, the sirries prints to samples of M.O and Mg.



 $15_6,\ 13.$  The total absorption of the 1285 cm  $^3$  MgC tank versus  $\underline{\bf y}_*$ 

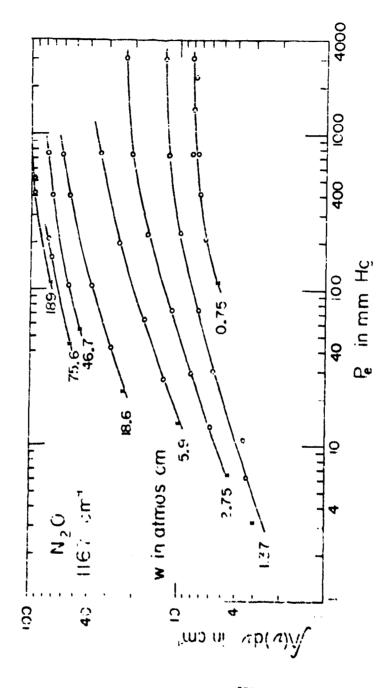


Fig. 12. The total absorption of the 1167 cm<sup>-1</sup> H<sub>2</sub>O band versus P<sub>e</sub>. The x points correspond to samples of H<sub>2</sub>O alone; the circle points to samples of H<sub>2</sub>O and H<sub>2</sub>.

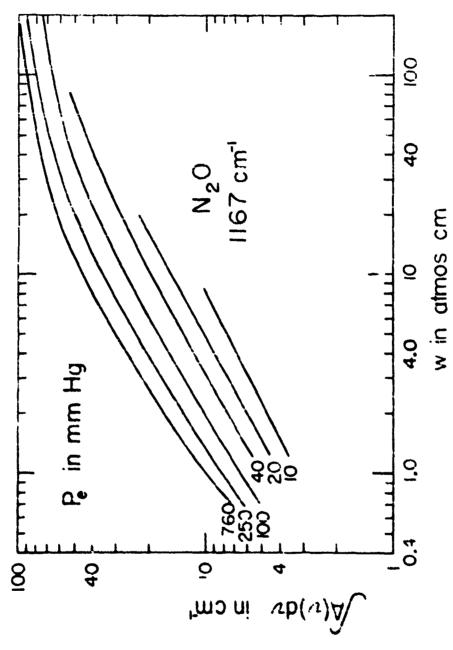


Fig. 15. The total absorption of the 1167 cm<sup>-1</sup> 4,0 band versus  $\underline{\mathbf{y}}$ .

absorption of the 1285 cm-1 band to equivalent pressure and to a quantity which is proportional to absorber concentration. Values of total absorption were expressed in units of microus instead of cm-1. These figures in the article by Goody and Wormell had been reduced so much that it was difficult to real values of total absorption more closely then approximately 1 2%. However, the values of these assermation of the outrement hypothetical samples were predicted by using the curves of Goody and Vormell and by using the curves of Fig. 11. The first sample contained an absorber concentration of 4.01 atmos-on at an equivalent pressure of 1:10 mm Mg. The surves of Goody and Wormall were used to predict a total absorption of approximately 63 cm<sup>-1</sup>, while Fig. 11 predicts 65.5 cm<sup>-1</sup>. Similarly, Goody and Worsell's curves predict a total absorption of approximately 29 cm-1, which can be compared with 33 cm-1 predicted by Fig. 11 for a sample of y = 4.01 and  $P_0 = 10$  mm Mg. Considering the fact that the curves cannot be read very accurately, one eccelules from these uso commerisons that the two sets of rusults are in at least fair egreement.

No curves were presented by Goody and Mornell to relate the total absorption of the 1167 cm<sup>-1</sup> band to absorber concentration or equivalent pressure; therefore, no comparisons could be unde between their results and those of the present investigation. The calculated band intensities from the two investigations are compared below in a subsection on band intensities.

In Fig. 1b are shown representative absorption curves for the 500 and 500 mm<sup>-1</sup> kp0 banks. The data for these two banks are listed in Table 7; only the samples of relatively large y produced a measurable absorption in the region of the 500 cm<sup>-1</sup> bank. The observed values of total absorption for this bind were tabulated, but were not plotted against y or P<sub>0</sub> since only a few netwood was consisted. The environmental bank was divided at the center. 500 cm<sup>-1</sup>; the total absorption of each pertion was necessared and listed in Table 7 along with the total absorption of the autire hand.

A large parties of the 589  $\rm cm^{-1}$  H<sub>2</sub>O data were plotted in Fig. 15 with total absorption shows as a function of P<sub>0</sub>. As for the other banks, values were takes from the curves of Fig. 15 to provide points for the empotion curves of Fig. 16, in which the total absorption was plotted against y for various values of F<sub>0</sub>.

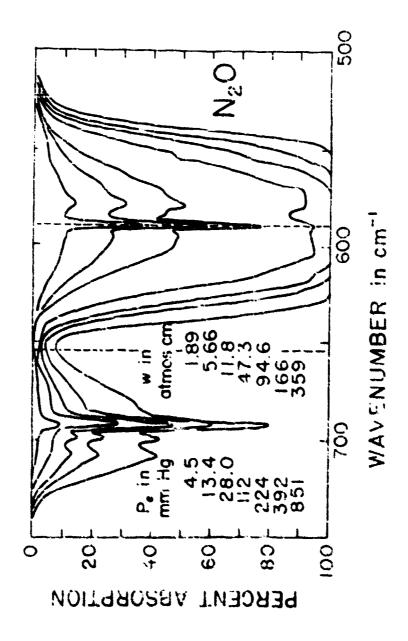


Fig. 14. Almosphion curves for the 969 cm<sup>-1</sup> and 692 cm<sup>-2</sup> B<sub>2</sub>) bands.
The vertical broken line at 555 cm<sup>-2</sup> represents the frequency of witch the brake were divided. Spectral alice width equals appreciablely 6 cm<sup>-2</sup>.

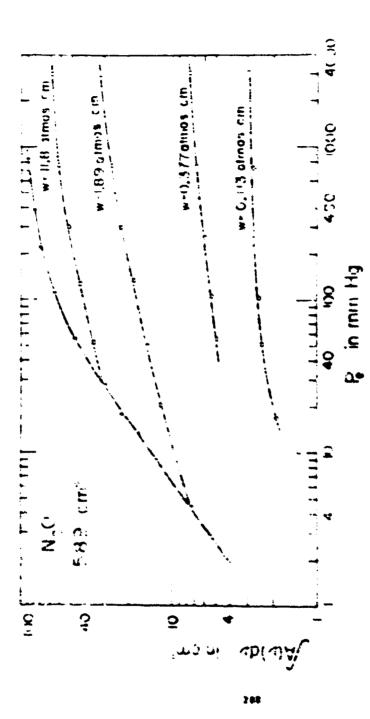
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7	2	3	6.11	7.55	ŝ	2.5	25.	
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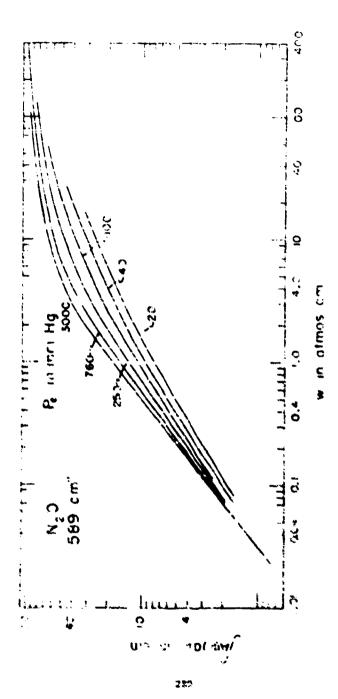
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<b>.</b>	K	ă	E	6.7	4.18	65.3	, 0
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The enerthiary is ensured total obstrytion is epiroximately z 5 percent for values greater than  $25 \cdot x^{-1}$ , and convening greater for emiller values.



is. 1. The facts spencytism of the 1977 card Mid book versus in the A prints correspond to manuface of Mid alone, the civile prints to seasile of Mid and Mg. Manufacture curve copiecents a can stant white of M. The soll increase presents of M. The soll increase presents



## Pand Intensities

is was stated in the part 1 that for complete of railing catty with pressure and less accorden rendentration,  $k(v)v \ll 1$  for all frequention where k(v) is the absorption coefficient, and the band increasing [Fig]2. Is related to the total absorption by

$$\int h(v) = -\frac{1}{v} \int \Lambda(v) dv. \qquad (2.5)$$

inder one conditions for which (c) is valid, the total absorpt the independent of equivalent recessors and is proportional to x. The slopes of the curves is the right-hand portions of Figs. The correspond to P = 7000 cm Hz are seen to increase toward in the x-in x value approximately equal to unity. It result which the conditions are satisfied for Eq. (iv) to be approximately time. It the most accurate values of land intensity were to be foles, it would be essential to obtain core data on the total absorption of surples at high equivalent pressures and low we however, by essentially extense error to complete with an alph-pressure curves of Fig. 7 and 8, as well so the corresponding righted for the other lands, with the alopes increasing an upper limit of unity, it was possible to estimate the lend interestive of four of the NgO bands reported here. The college contains with the lands with values of other workers to comparison.

Table o. MgO, this intensities"

		•		
Band	Present 1 day	W: 150% et. 145	Goolg sign horaell	Regers and deswrota
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2461	Elip # 2.5			
1295	•	<b>46.</b> J	•	. •
1101	12.0 : 1.0		a de e	J.
90)	23 4 %	4:,		28.9

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in du y marilly based As 1979 The particle pressure of \$100, five which y rould easily be calculated, was determined by measuring the description of the sample in the absorption cell. This was accomplished by the ere of a "well" made of a thin place of "income." Uning, which was closed on one and and was suddered to the wall of the absorption cell, projecting approximately, " Labor into the interior of the cell. Ether was put into the tube and \$5 was bubbled through in order to cool it until condensation formed on the cutaids washed the tubing, which could be observed through a vinion in the absorption cell. A thermometer interest in the other was used to detarmine the temperature at which condensation started.

The accuracy with which the dew point could be determined deprecial primarily upon how accurately one could tell when condensation started. and also upon how accurately the thermometer read the temperature of the surface on with. U. B.O candenced. One would report that tembelence of the other, which was caused by the subbling of the nitrages, would test to minimise temperature gradicate in the citer; it was found that the reading of the thermomies varied by not more than approximately # 2 % to i 0.200 when held at different positions in the other. It was found that. as the ether was couled slowly, a temperature was reached at where a very thin film would appear on the surface of the tubing; if this tamperature were maintained, no further condensation seemed to occur. Hemever, if the temperature were decreased by approximately 0.2%, the condensation increased. In the case of samples of HoO alone, the partial pressure could be measured with a senometer and compared with the partial pressure determined by use of standard tables relating temperature at which emidensation occurred to the partial pressure. It was found that the values of 350 partial pressure determined by the two mothous agreed to approvimately 2 2 percent if the temperature used was the one at which one east. tion would increase, rather than the alightly higher temperature at think a thin film first appeared on the surface of the tholing.

It is believed that the values of which were determined are accused to 12 or 13 percent for samples whose dow-point was greater than 10°C, and to 23 or 16 percent for samples having lawer day points. Paviations in what approximately 1/2 percent could be observed. RgO was introduced to the absorption cell by admitting a measured amount of liquidition into a small evacuated bulb which was connected to the absorption cell through a value and a short piece of support tubing. The value and tubing connecting the bulb to the absorption cell were heated sently to avoid condensation in these parts; the value was then opened and the H-O allowed to evaporate into the previously evacuated all. The evaporation could be openeded up by gently heating the liquid dpd.

by considering the volume of the absorption (all, it was possible to calculate the greatment that would be expected from a snown quantity of liquid light from a snown that would be expected from a snown quantity and note were adsorbed on the case valls. In most cases to was round that the present present was only approx sately \$0 to tO percent as great as the calculated value, a result with indicated that approximately

1/2 the HgO was adsorbed on the cell walls since care was taken to avoid condensation. If the sample was left in the cell, the pressure was found to decrease gradually for a few hours, thus indicating that more adsorption was occurring.

In the process of obtaining data No was added to the HgO to produce samples at various values of total pressure less than one atmosphere. The day point, from which the HgO partial pressure could be istermined, was measured periodically, at least once or twice after adding each sample of Mg. It was round that the day point remained essentially constant if the HgO had been allowed to remain alone in the cell sufficiently long for the absorption in the complete before Mg was added. A major portion of the investigation took place during the summer months when the temperature of the absorption cell, which remained near ambient temperature, varied by as much as 6 or doc during a given day. As the temperature of the cell increased, a portion of the adsorbed gas apparently was "driven off" to cell valls since the HgO partial pressure increased at a rate considerably greater than one would predict on the basis of the perfect gas laws. The fact that adverted across are "driven off" solid surfaces by heating is, of course, will know.

Various coatings on the interior of the cell were tried in an attempt to decrease the absorption; these coatings included automobile lacquer, glyptal varnish, and a filr of vacuum pump of the so of these surface coatings seemed to change significantly the amount of 500 adsorbed; however, more time was required to "is-ass" the call by evacuation when use of the show coatings was used. Therefore, the call was used with the bare steel expused. Little difference was observed between the adsorption of the steel cell when it was an abserved between the adsorption of the steel cell when it was an abserved. A porcelain coating or given lintag would probably refuce the cisorption, but this would have involved considered expense and was not cannidered essential since the MSO partial presence could be mon tored.

As a result of the investigation of the adsorption of EgO on the value of a steel cylindor, it can be concluded that accurate values of y causes se determined from a measurement of liquid EgO introduced into the cell; in general, values determined by this method would be too high. In the work by EDN, values of y for EgO samples were determined by measuring the EgO partial pressure before Eg was mided. It has then assumed that the EgO partial pressure remained constant as different amounts of Eg were added and the absorption numbers seem obtained. On the basis of the investigation of adsorption described above, it can be concluded that the EgO partial pressure was probably not changed because of Eg displacing the EgO which had been adsorbed on the walls of the adsorption hall. The greatnet causes of variation in the EgO partial pressure was probably changes in temperature, or further adsorption built with have occurred after the partial pressure was measured. Buring the major portion of the absorption wall was nearly or stant, increfore, any error in the determination of y tecasis of temperature variations was probably small.

1/2 the H<sub>2</sub>O was adsorbed on the cell walls since care was taken to avoid condensation. If the sample was left in the cell, the pressure was found to decrease gradually for a few hours, thus indicating that more adsorption was occurring.

In the process of obtaining data N2 was added to the H2O to product samples at various values of total pressure less than one atmosphere. The dev point, from which the H2O partial pressure could be determined, was measured periodically, at least once or twice after adding each sample of M2. It was round that need to remain alone in the cell sufficiently long for the adsorption to be complete before M2 was added. A major portion of the investigation took place during the summer months when the temperature of the absorption cell, which remained near ambient temperature, varied by as much as 6 or dOC during a given day. As the temperature of the cell increased, a portion of the adsorbed gas apparently was "driven off" the cell valls since the H2O partial pressure increased at a rate considerably greater than one would predict on the basis of the perfect gas laws. The fact that adsorbed gases are "driven off" solid surfacer by beating is, of course, well known.

Various costings on the interior of the cell were tried in an attempt to decrease the absorption; these costings included automobile lacquer, glyptal varnish, and a film of vacuum pump of . Home of these surface costings seemed to change significantly the amount of Eq0 adsorbed; accever, more time was required to "is-gas" the cell by evacuation when one of the above creatings was used. Therefore, the cell was used with the bare steel exposed. Little difference can observed between the adsorption of the steel cell when it was nemerical rusty as crepared to the cell immediately efter it had been sand-blasted. A porcelain cesting or glass lintag would probably reduce the discrption, but this tould have involved considerable expense and was not scanidered assertial since the EgO partial pressure could be monitored.

As a result of the investigation of the adsorption of MgO on the valle of a steel cylinder, it can be openimied that accurate values of w carnow on de trained from a measurement of liquid 120 introduced into the cell; in general, values determined by this method would be too high. In the work by E'd, values of y for MgO samples were determined by seasuring the MiO varial pressure before it was middle it has then assumed that the MgO partial pressure remained constant as different escents of Mg were added and the abs aption curves were chimised. On the tasis of the investigation or adsorption described above, it can be concluded than the MGO partial pressure was probably not changed because of My displanting the MyO which had been adsorbed on the walls of the ad-The greatest causes of variation in the MgO partial surption ---pressure were probably changes in temperature, or further edsorption will with have occurred after the partial pressure was pervised. During the main portion of the investigation by MM the trup's atturn of the about prich walk was nearly constant, unerefore, any error . the detertime! I of w cocause of temperature variations was probably small.

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Errors arising from further adsorption after the HoD partial pressure was measured were probably not excessive since investigation of the results shows that the total absorptions of rew samples did not change appreciably after having remained in the cell overnight.

Comparison of the present results with those of HBW which dovered the same range of values of y and % indicated that the two sets of results agreed within expected experimental error.

### Experimental Results

Absorption curves were obtained for the 5532, 5700, and 1595 cm<sup>-1</sup> H<sub>2</sub>O tends by use of a Perkin-Klaer Model 99 spectrometer enclosed in a vacuum tank which is described in Appendix I. A LiF prism was used for the 5532 and 5700 cm<sup>-1</sup> bands and an HaCl prism for the 1595 cm<sup>-1</sup> band. The method used to replot the recorded spectra on a linear wavenumber scale and to measure the total absorption is the same as that described by HBM.

The data and results are presented in a minner similar to that described in the previous section for the  $M_2O$  bands. Some of the curves which are presented are based on both the present results and the HBW results. Values of equivalent pressure  $P_{\phi}$  were calculated for each sample by the following equation:

$$P_n = P + bp, \tag{16}$$

where F is the total pressure due to HgO at partial pressure p and to the Mg in the sample.

In Fig. 1; are snown representative absorption curves of the 555% cm<sup>-1</sup> MgO band; the absorption data for this band are listed in Table 9. The estimated uncertainty of the tabulated values of total absorption is approximately 2 6 to 8 percent for values greater than 50 cm<sup>-1</sup>. The relative uncertainty is somethat greater for smaller values of total absorption. The socuracy of the tabulated results for the other MgO bands listed below to approximately the same.

Representative portions of data for the 5332 cm<sup>-1</sup> band were plotted in Fig. 18 with total absorption versus P<sub>0</sub>, where each curve corresponds to a given value of y. It is noted that in This 9 the camples are all groups of approximately equal y; in order to pack the results of these data in Fig. 18, an average value of y was obtain for each group of semples and the observed values of total absorption which are listed in the table was concreted to the average y. The correction was obtained from plots of total absorption versus y where each curve was detarmined by four points comes (possible to the four different path lengths at which each gas of voters was estanted.

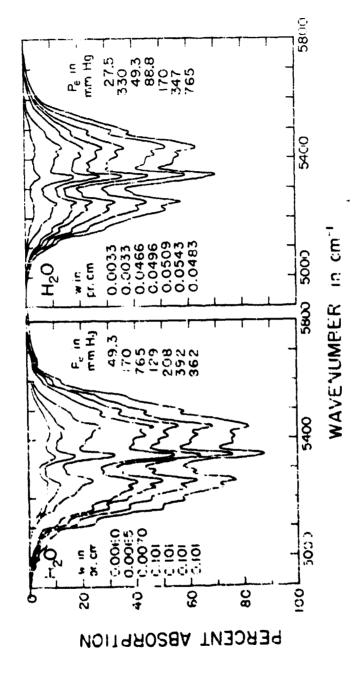


Fig. 17. representative absorption curves of the 5%52 cm  $^{1}$  HgO hand, approximately 20 cm  $^{1}$  .

Table 9. Data for the

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1 2		Pats Length	9 (R <sub>2</sub> 0)	5 te 1	. Z	ar.	\[ \langle A(\nu) \dv \tan \tan \tan \frac{1}{5332}  \tan \frac{1}{2}  \text{Band} \]
	8.0	62.5	5.5	5.5	27.5	1	11.7
~	0.6%	Ą	, , ,	0.2	いま		13.0
~	Š	æ	5.5	21.5	43.5		16.1
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~	\$. \$.	Ŗ	5.5	83	747		27.0
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t·-	٠ <b>9</b>	ž	9.;	<b>3</b> 2	<b>3</b> 82		51.:
ಖ	<b>3</b> .c	Ŕ	6.6	6.6	49.3		26.3
C	26.5	8	₹.Q	2.5	8		
9	77.5	8	9	47.C	9.88 8.6		35.7
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7	<b>98</b>	<b>&amp;</b>	<b>10.</b> 6	Š	法		56.3
2	o. <b>⊀</b>	Ą	2.01	724	765		9.47
4	8.5	25%	31.6	Š	13		87.6
<b>5</b>	0.43	०६टा	01	72.	165		109
91	35.0	26.45	9.0	6.9	49.3		53.1
1.7	8	24.5	£0. \$	27.2	68.2		21.0
<b>8</b> 1	i.	2445	ខា	6.7.	89.8		6.69
ć1	2 <b>9</b> .5	24.5	:. S	ጃ	170		8.0
-	٠. ون	24.45	1	Š.	K		116
17	<u>ئ</u> .د	2 <b>64</b> 5	3.01 3.01	124	765		<b>†</b> T
3	25.5	<b>48</b> 75	C. 10	9.9	₹.64		98.6
2 4	2 K	\$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$	š.	27.5	(S) 28		75.1
•	· · ·	* U.T.	ें	?	3		3

Sable 5. (Continued)

Serrelk Ek.	į ર	Part. Length ca	38 32 2 (15/2)	5.57	e #	7r. ea	)A(v)dv in cm-1 5352 cm-1 Band
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9,0	86.5 .5	, 1 <b>0</b>	7.11	ጁ	, i	C.0543	797
12	34.0	-875	ĭC.2	73.	765	5.048₹	702
92	3.0	÷875	8	2.5	1113	0.109	15.
8	20.7	±375	21.12	45.0	621	0.101	15.
2	3	đ T	21.7	ន្ទ	80 80 80	101.0	186
ĸ	٩.٠ <u>۲</u>	+375	21.8	ጳ	<u>%</u>	3.101	762
9.	¥.5	+375	8 E	11.	<b>8</b> 8	0.101	<b>562</b>
*	χ; ο.	*	5.17	7:33	15.9	0.00192	6:
ス	o, ₹,	Š	6.7	, ,	33.5	0.C)1605	17.0
x	X.	38	9.5	3.3		0.3356	87.0
ķ	<b>X</b> .0	Š	12.*	* . *	o. 8	0520010	33.7
×	23.0	Š	14.6	11.5	75.0	0.00850	36.0
92	\$2.0	ş ilga	5.17	3.17	15.9	0.0150	30.6
Ç,	<b>%</b>	, <u>19</u>	5.1	6.7	33.5	0.0316	48.8
ş	;;	, <b>6</b>	9.5	٠ <u>٠</u>	0.94	が き。 。	76.9
7	35.0	£24	15.4	٠. در	0.39	0.0586	88.6
ņ	<b>%</b>	19	14.6	91	73.0	1990.0	104

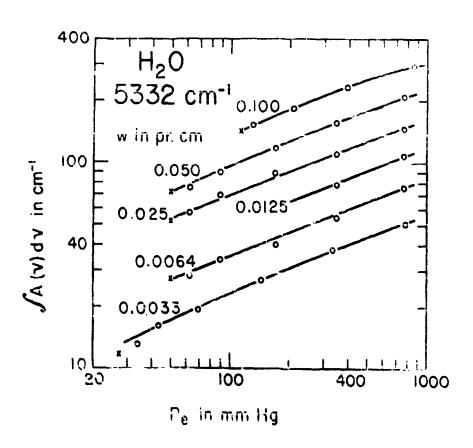


Fig. 10. The total absorption of the 5552 H<sub>2</sub>O cm<sup>-1</sup> bond totals P<sub>2</sub>.

The x points correspond to sample of H<sub>2</sub>O and H<sub>2</sub>.

points to samples of H<sub>2</sub>O and H<sub>2</sub>.

The curves of Fig. 19, in which total absorption was plotted against where obtained by using both the present results and the HBW results. Values of botal absorption were caken from the curves of Fig. 18 and from similar curves which were drawn for the HBW data but are not presented in this report. These values were then used to provide points for the curves of Fig. 19, in which the points were omitted in order to facilitate reading the curves.

Over a small range of values of withe How data overlapped the present data and provided a convenient comparison. Curves drawn through each of the two sets of data without consideration for the other set would have agreed to within 5 to 10 percent for most of the different equivalent pressures illustrated in Fig. 19, as well as for the corresponding figures sham below for the other bands. This agreement is good when one considers that the HDW data being compared represented samples of small w for which the accuracy was pointed. Comparison of the present data with those obtained in the early work of Howard and Chapman? Was much loss fewerable: the results of total absorption for similar samples often differed by as much as 25-40 percent.

In view of the improved methods which were described above, the present results were given extra "weight" in determining the positions of the curves of Fig. 19, which are believed to best represent all of the data; and values obtained from these curves are believed to be accurate to it 6 percent. The same accuracy is estimated for the corresponding curves of other NgO bands shown below.

Reproductative absorption curves of the spectral region near 3700 cm<sup>-1</sup> are shown in Fig. 20. The data are lifeted in Table 19, which includes the values of the total absorption for the three spectral intervals: 2009-3540 cm<sup>-1</sup>, policyfown cm<sup>-1</sup>, and 1000 fine cm<sup>-1</sup>. The sum of the total absorption are intervals, which together are referred to as the 5700 cm<sup>-1</sup> hear, is also given for each sample. The values of the total absorption for the 2000-3540 cm<sup>-1</sup> legion, which corresponds to the 3.50 manh in the 2000-3540 cm<sup>-1</sup> argion, which corresponds to the 3.50 manh in the 2000-3540 cm<sup>-1</sup> property.

Figures 21 and 22 contains obver existing the total shorption of the 3700 cm<sup>-1</sup> band to P<sub>c</sub> and y, respectively. These curves were obtained in the same names as were the corresponding on sea for the 5552 cm<sup>-1</sup> and:

Figures 25 and 24 show represent tive absorption curves or the 1595 cm<sup>-1</sup> hgC band. The data are listed in Table 14, the values of total absorption are included for the 1250-550 cm<sup>-1</sup>, and 15 0-2100 cm<sup>-1</sup> intervals and for the sum of these two intervals, which is called the 1595 cm<sup>-1</sup> ago band. The data for the hand were treated to the name ammor as were the data for the 5532 cm<sup>-1</sup> and 5700 cm<sup>-1</sup> bands. Figure 25 compation of the present data to P<sub>6</sub> for carbon values of w, and the curves of Fig. 20, which is set total data.

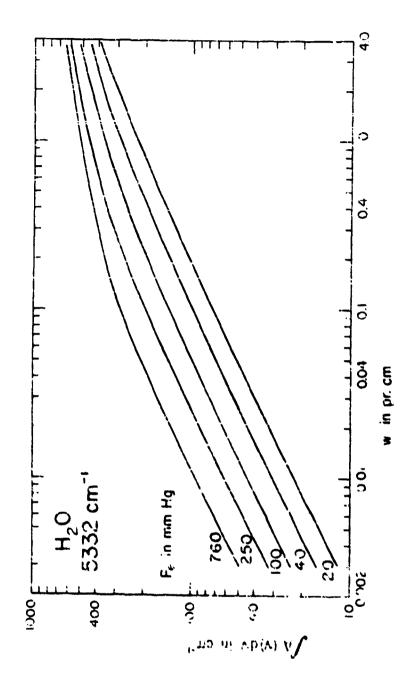
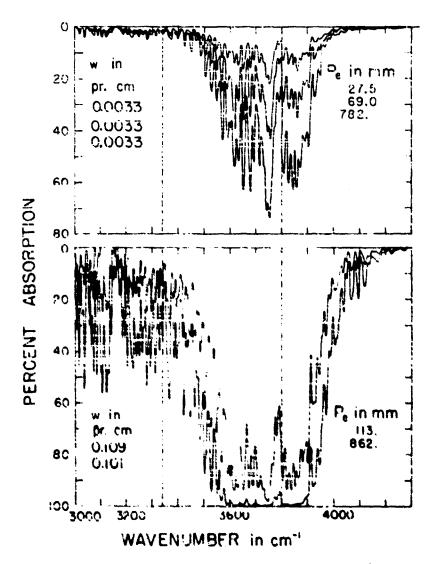


Fig. 19. The total absorption of the 5552 cm<sup>-1</sup> H<sub>2</sub>O band vernus y. The curves are based on HM data for larger y and on present data for smaller y.



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Table 11. (Continued)

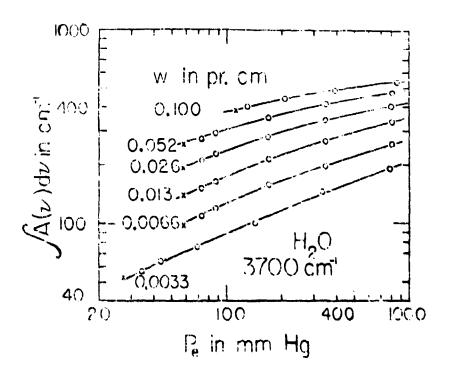
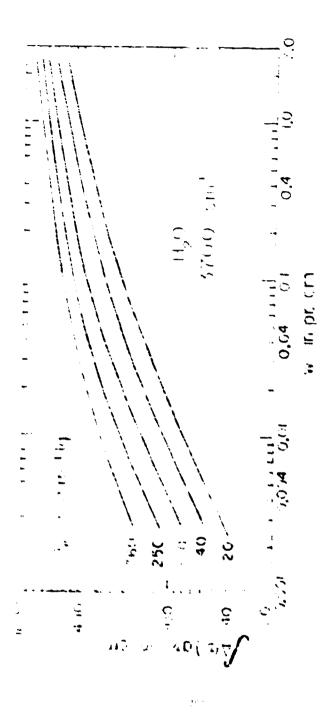
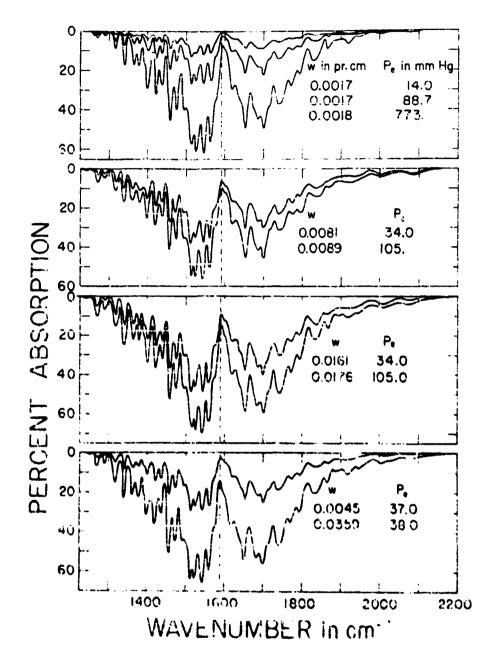


Fig. 21. The total absorption of the 3700 cm<sup>-1</sup> H<sub>2</sub>O band vertue P...
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Cap. 23. A conscive absorption curves of the 1505 cm<sup>-1</sup> foo band. Special 23th older equals approximate a cm<sup>-1</sup>.

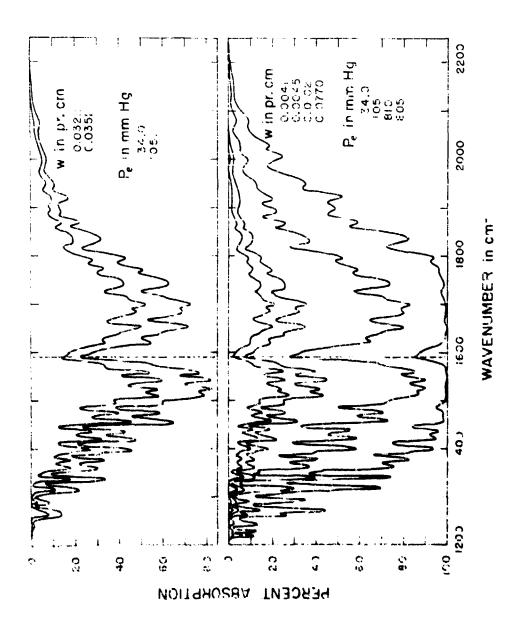
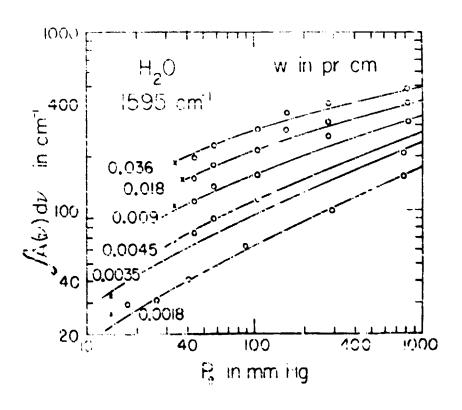


Fig. 2h. B. presentative above for curvet of the life of 2 20 band. Species, slit width erads approximately of curl.



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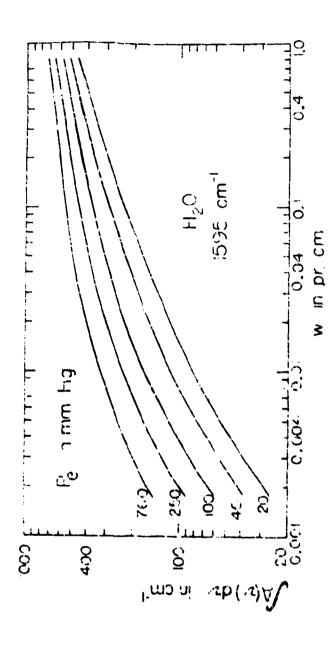


Fig. 10. The Unit mborption of the 1999 emit 9.0 and versus y. Inc curves at passed at 48. data for lable of y and on precent data for data for lable of y.

Oktle 11. Data for the 1595 cm<sup>-1</sup> H<sub>2</sub>O Ban3

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#### SUMMARY

The self-broadening coefficient has been determined for saveral bands of N<sub>2</sub>O, CO, Ch<sub>4</sub>, CO<sub>2</sub>, and H<sub>2</sub>O; thus coefficient, which is defined in the text, is a measure of the relative broadening ability of the absorbing gas with respect to N<sub>2</sub>. By the use of the experimentally determined self-broadening coefficients it was possible to define an equivalent pressure  $P_{\rm e}$  which served to simplify the analysis of the data by accounting for the broadening by the absorbing gas and by N<sub>2</sub> with the use of a single variable. Values of B used in the analysis of all the bands included in the present research program are summarized in Table 1.

The relative broadening abilities of various inactive gases have been investigated and, where possible, the results have been compared with those of previous workers. The results are shown in Tables 2, 5, and 4.

Absorption curves were obtained from which the total absorption has been determined for the N<sub>2</sub>O bands at 2565, 2461, 1285, 1167, 692, and 589 cm<sup>-1</sup>. The data are listed in Tables 5, 6, and 7, and curves have been drawn to show the functional relationship between total absorption and the two parameters, absorber concentration  $\mathbf{v}$  and equivalent pressure  $\mathbf{P_e}$ . Values of the band intensities of part of these bands have been determined and compared in Table 8 with the results reported by other investigators.

Total absorption measurements have also been made for the  $\rm H_2O$  bands at 5332, 3700, and 1595  $\rm cm^{-1}$ . The results, which are presented in tebular form and in graphical form with their absorption plotted against  $\gamma$  and against  $P_{\rm e}$ , supplement earlier work by Howard, Burch, and Williams, and serve as a check against the results of these workers for values of  $\gamma$  and  $P_{\rm e}$  included in both. There is good agreement between the earlier results and the present results, which were obtained by improved methods.

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#### APPENDIX I

#### A VACUUM TANK FOR A PERKIN-ELMER Sangale-Bram medicidingmanyon

The need for evacuating the optical path of a spectrometer used for investigating strong atmospheric absorption bands was discussed in Report 2. Enclosing the Perkin-Rimer Model 21 Jouble-beam spectrometer in a vacuum tank semmed to be impractical; therefore a vacuum tank was designed and constructed to contain a Perkin-Elmer Model 99 single-beam monochromator which was aveitheble in the "aboratory. This tank, which is illustrated in Fig. 27, is discussed below.

The trial is feedback by the shape and inexpensive construction and by its versatility. The main part of the tank consists of a piece of 22-inch o.b. steel pipe with 3/8-inch thick valls, and the end plates are circular pieces of 3/2-inch thick steel plate. The monochromator and other components are set on an 15-inch-wide steel plate which is spot-welded inside the tank. A portion of the top of the tank was cut away and the hole covered by a lid of 3/4-inch thick plexiglass, which makes it possible to see all the enclosed parts. An endianal piece of pure gum rubbon and serves as a gasket between the plexiglass and the steel tank. Rubber channel gaskets provide seals between the cylindrical tank and the flat end plates.

Electrical connections are provided through homsetically sealed porcelain fittings. In order to avoid excessive noise which might arise from a fitting in the circuit cetween the preamplifier and the amplifier, the cable connecting those two components was put through a hole in the end-plate which could easily be sealed with a putty-like sealing compound. This cable was covered with an air-tight plastic conting; but it was necessary to open the plastic coating and put soft was around the wires in the cable to prevent leakage through the cable from one end to the other. It should be noted that small leaks can be supported aince the vacuum pump is left on during operation of the ampediturester.

Mechanical adjustments can be made through one of three "rotating seals" ande of a piece of cross atook which was soldered over a hole in the tank and drilled through to accept a piece of 3/6-inch drilled. A circular groove was machined but of the lines surface of the brass piece to hold a small "O" ring to form a seal around the drilled. These seals were found to be virtually leak proof, and make it possible for adjustments to be said by a sutating motion or by a linear motion by sliding the rod through the wal. The acchanical consections were used to make minor an intensity of the mirrors, to operate a sautter as items of the entrance sitt, to operate a valve on a small unsurface well, in mights the sitte, and to drive the Litterov mirror. The gear box and recay system for the Litterov drive were removed from the

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Fig. 17. Browse, of victims that containing a Perido-Education Note: J. protection for apparent is do without in Appendix I of Alia repair.

external means were consisted to the mesself the salars through the second means means were consisted to the mesself after through the second related to the control of the second related to the control, it was possible to use a smaller tank, to reduce the second control of electrical confidences of through the tank, and to avoid the shift minht make from operating, the small bitter widelve mater and the relays in a vector.

A useli gran-wow inside the tank cakes it possible to make four emporate adjustments with only one usel. A spur gens on the rotating elections be moved to mesh with may one of the outlier of a power which are connected to the various adjusting mentantams.

The optical dirgram above the main provide of Fig. 21 is a convenient arrangement for the one of a sharter accomption cell contained inside the various back. See can be admitted to the vell through a piece of copper tubing which extends through the wall of the bank.

The option are unparent in the coin partition bag, 27 to convenient for the one of the injector case rength multiple-traversal cell und in obtaining data described pressonally. The end-plate of the vacuum tank true against the end-plate of the ensorption cell; a 1/2 tash endless gaz rubber gasket is held in place between the two end-plates. The end-plate of the vacuum tank is cut away appeare the windows of the automation cell.

A 5:2 magnified large of the Normat glover is formed inside the multiple traversal call the image size is maintained during the resolutions in the call and then reduced by 2:1 at the entrance alit of the monochromator. The light beam then rills the optics of the monochromator.

No serious difficulties were excountered as a result of operating the monochromator in vacuum. The disper mechanism apparently needs labeled to more frequently when operated in vacuum, although this was not narticularly troublessme. Material tends to sublime from the glower and form tain films on nearby objects; it was therefore necessary either to keep the source neveral inches from any windows, or to shield these parts from the glover. The two rather minor difficulties described shows could probably be avoided by evacating one lank with the instruments off, and then filling it with No before operating the lastingest.

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